

Properties of nuclear masses for heavy and neutron-rich nuclei

Hiroyuki KOURA(小浦寛之)

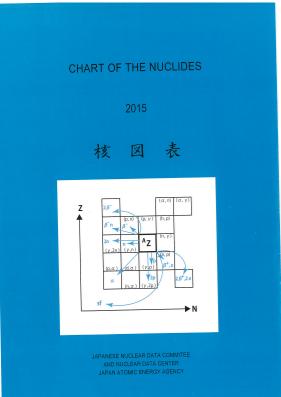
Advanced Science Research Center
Japan Atomic Energy Agency (JAEA)

- I. Introduction
- II. Bulk properties of nuclear masses
- III. Nuclear mass formulae
- IV. Application to r-process
 - i) Mass-model dependency
 - ii) Effect of fission
- V. Recent study on beta-decay
- VI. Summary

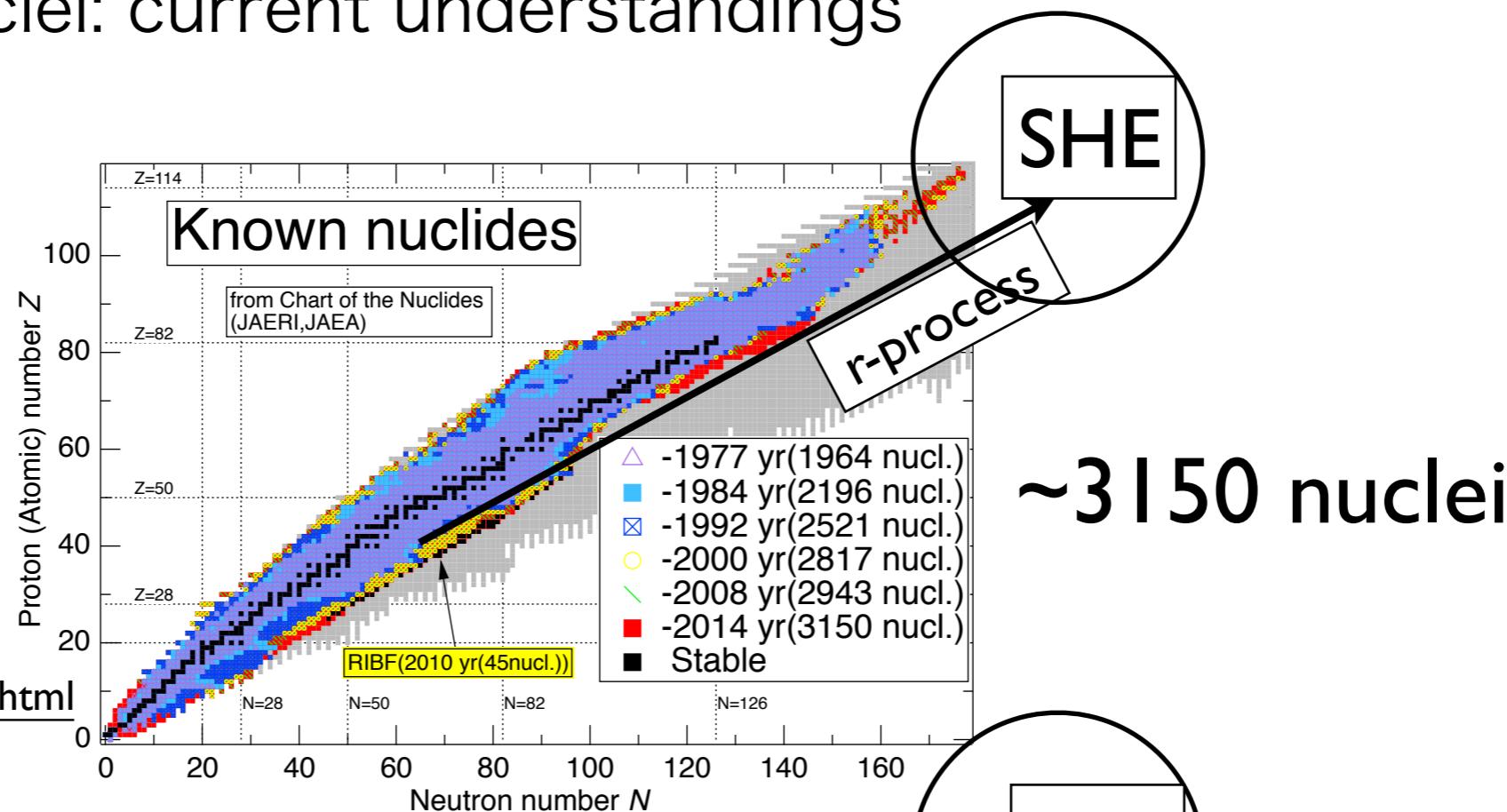
I. Introduction

Search of nuclei: current understandings

Identified



taken from Chart
of the nuclides by
JAERI and JAEA
(HK, et al., 2015)

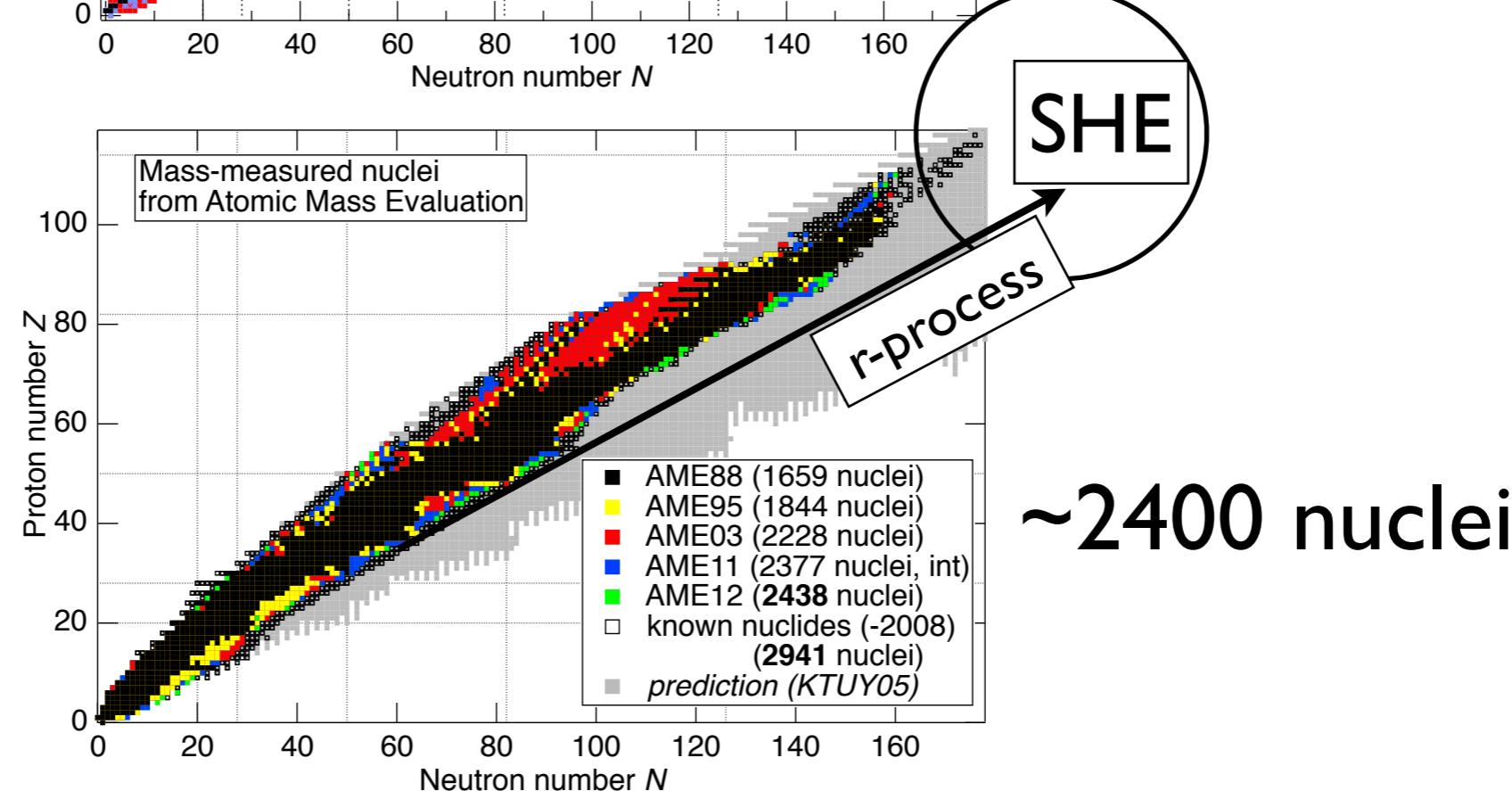


Mass-
measured



Atomic Mass
Evaluation is updated
as AME2016

amdc.impicas.ac.cn/



II. Bulk properties of nuclear mass

Weizsäcker-Bethe semi-empirical atomic mass formula

$$M_{\text{WB}}(Z, N) = Z m_{\text{H}} + N m_{\text{n}} - B(Z, N)$$

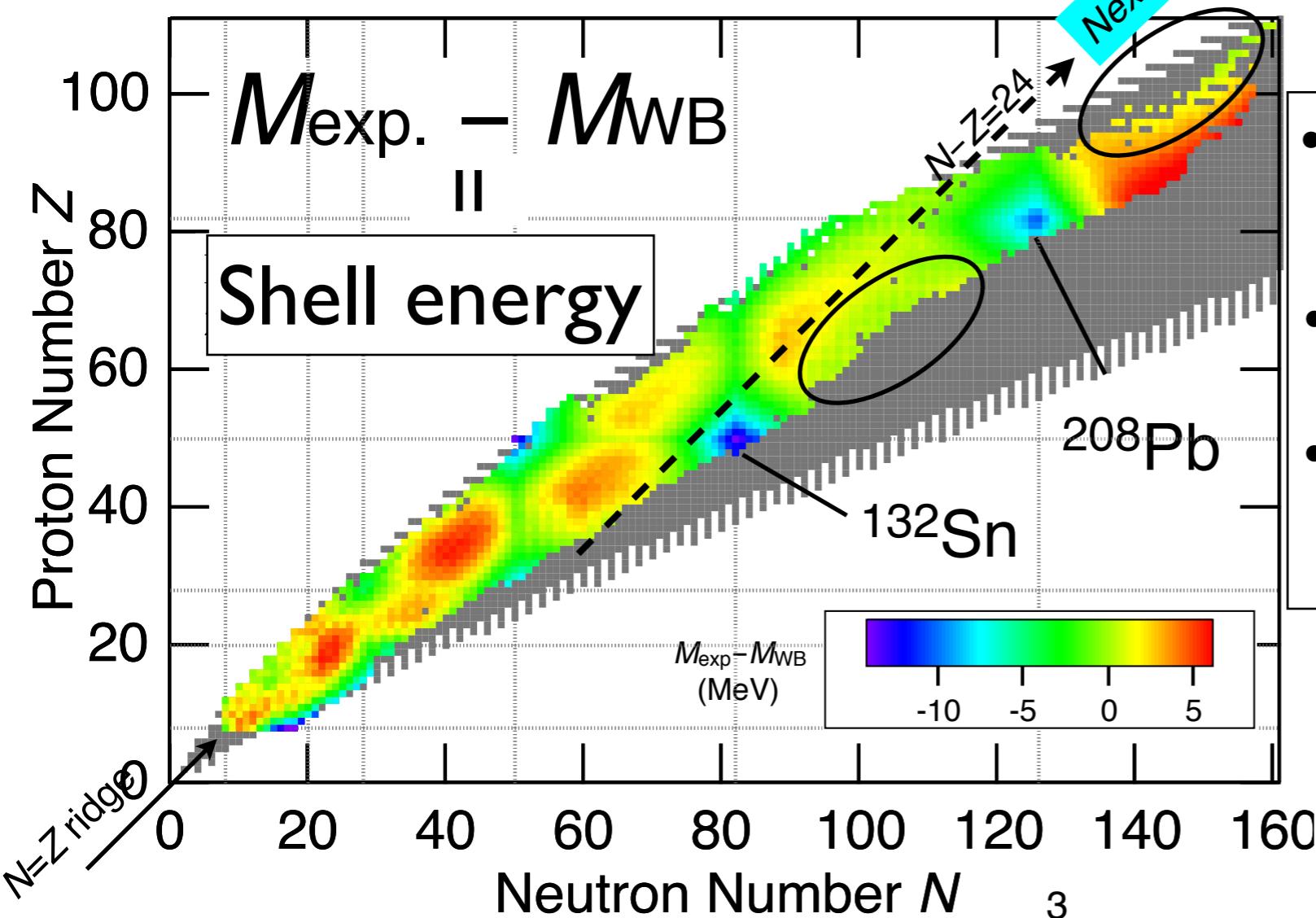
$$= Z m_{\text{H}} + N m_{\text{n}} - a_V A + a_S A^{2/3} + a_I (N-Z)^2/A + a_C Z(Z-1)/A^{1/3} + \delta_{eo}$$

a_V	a_S	a_I	a_C	a_{eo}
15.604	17.472	22.99	0.7073	12.338

(MeV)

$$\delta_{eo} = \begin{cases} -a_{eo}/A^{1/2} & \text{for even-}Z \text{ and even-}N \\ 0 & \text{for odd-}A \\ +a_{eo}/A^{1/2} & \text{for odd-}Z \text{ and odd-}N \end{cases}$$

Next figure



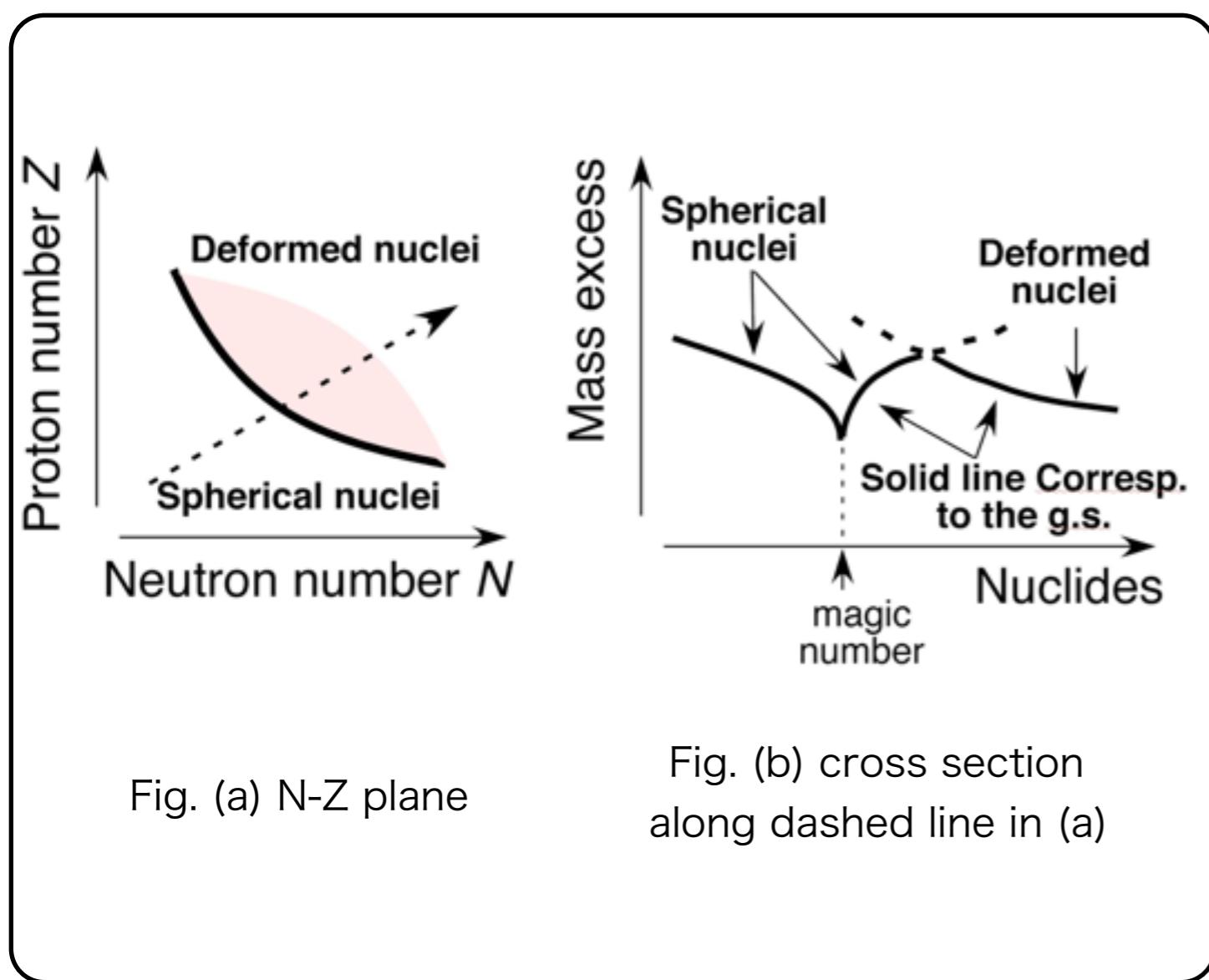
From the 'exp.' shell energy:

- Existence of magic number
 $N=28, 50, 82, 126$
 $Z=28, 50, 82$
- Wigner energy
 $N=Z$ ridge
- Depression due to the deform.
rare-earth, actinide

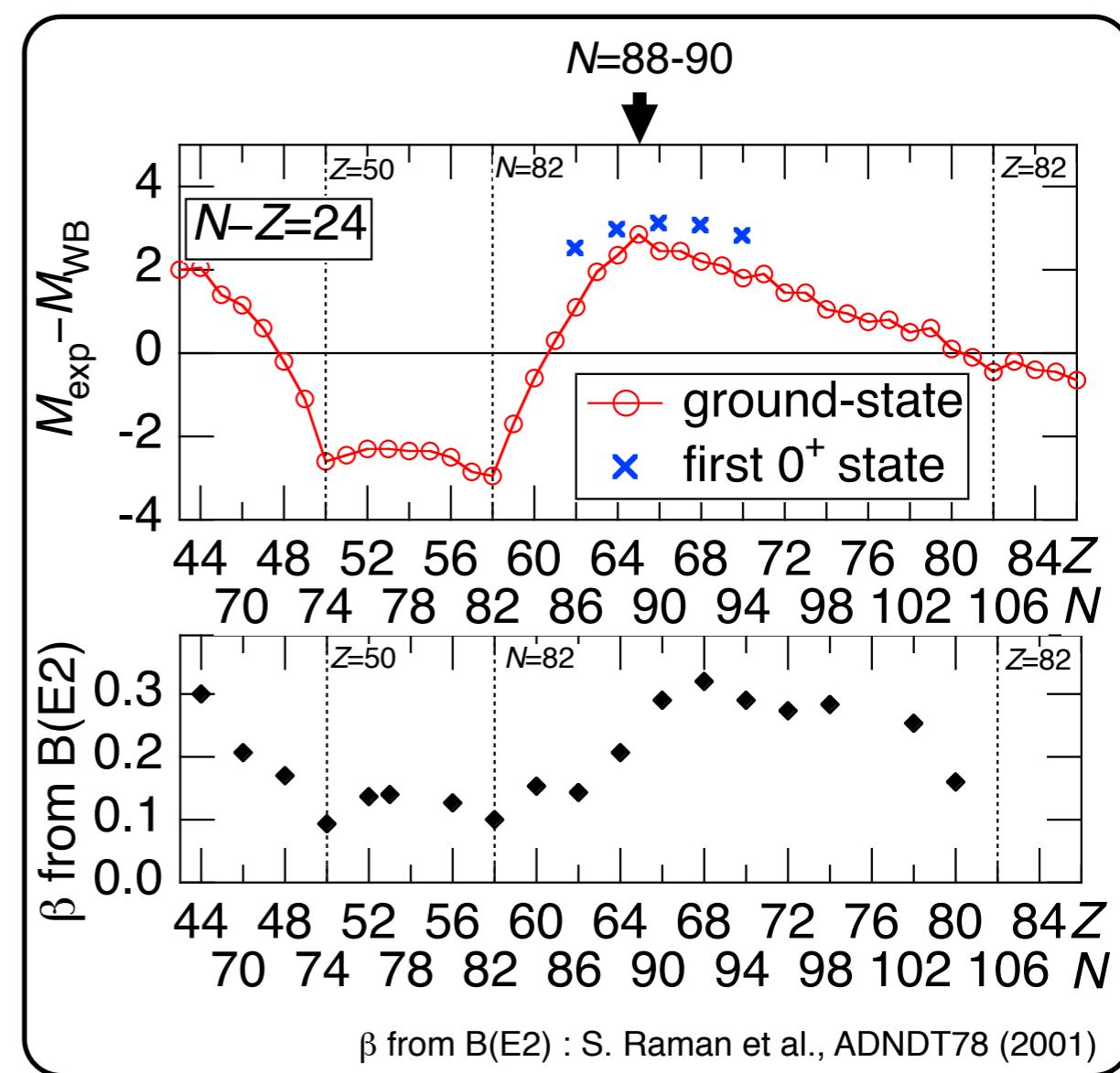
Mass data : 2012 Atomic mass evaluation
(M. Wang, G. Audi, A.H. Wapstra *et al.*)

Trend in MeV-order

- Schematic -

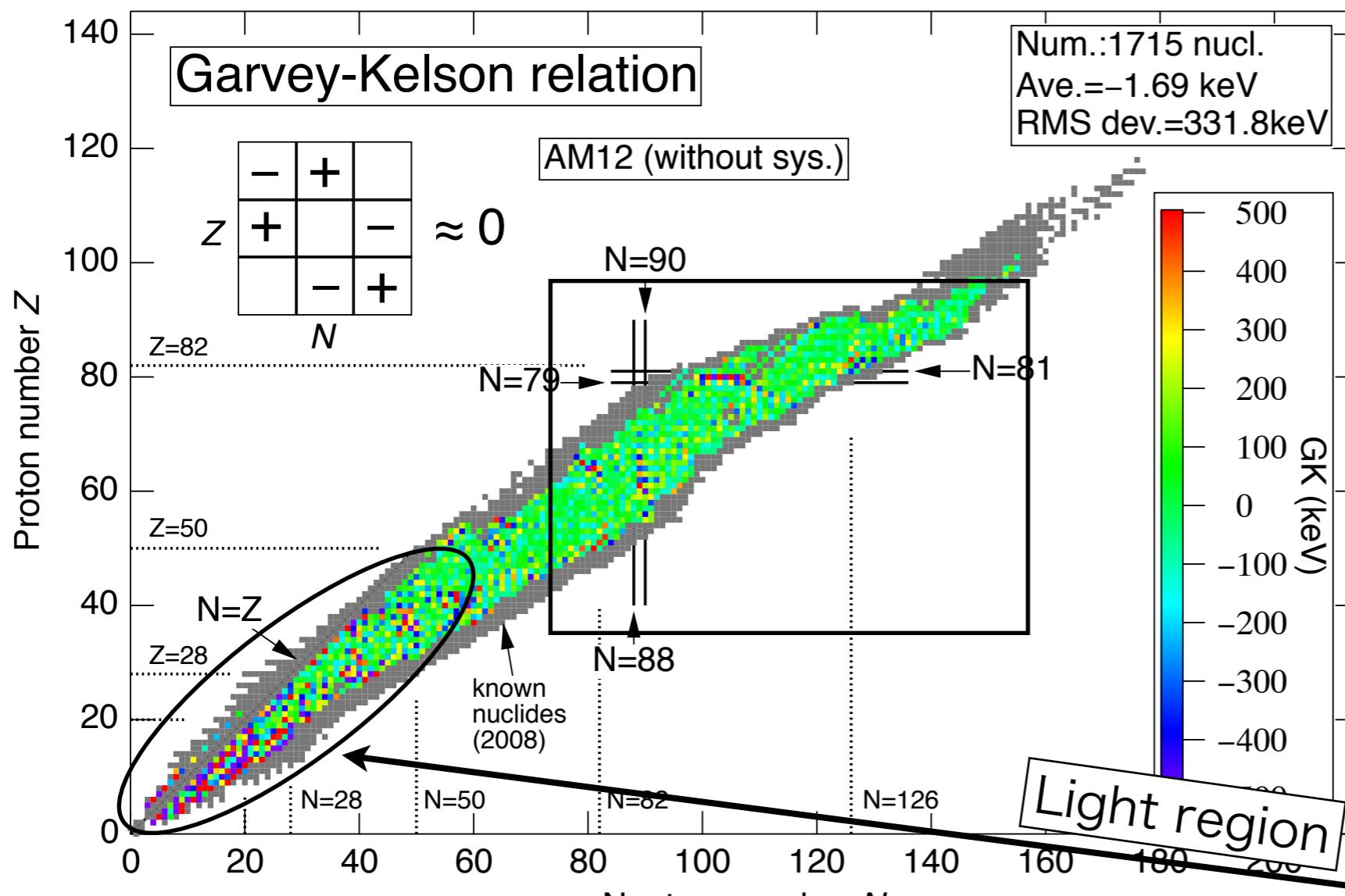


- Experiment -

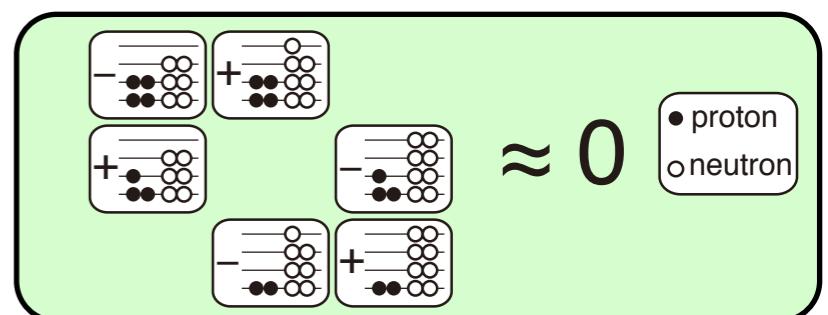


Notable feature on discontinuity of derivative of mass values

- $Z=50$, $N=82$ and $Z=82$ discontinuity of derivative: Spherical single-particle shell closure
- $N=88-90$ discontinuity: Shape transition



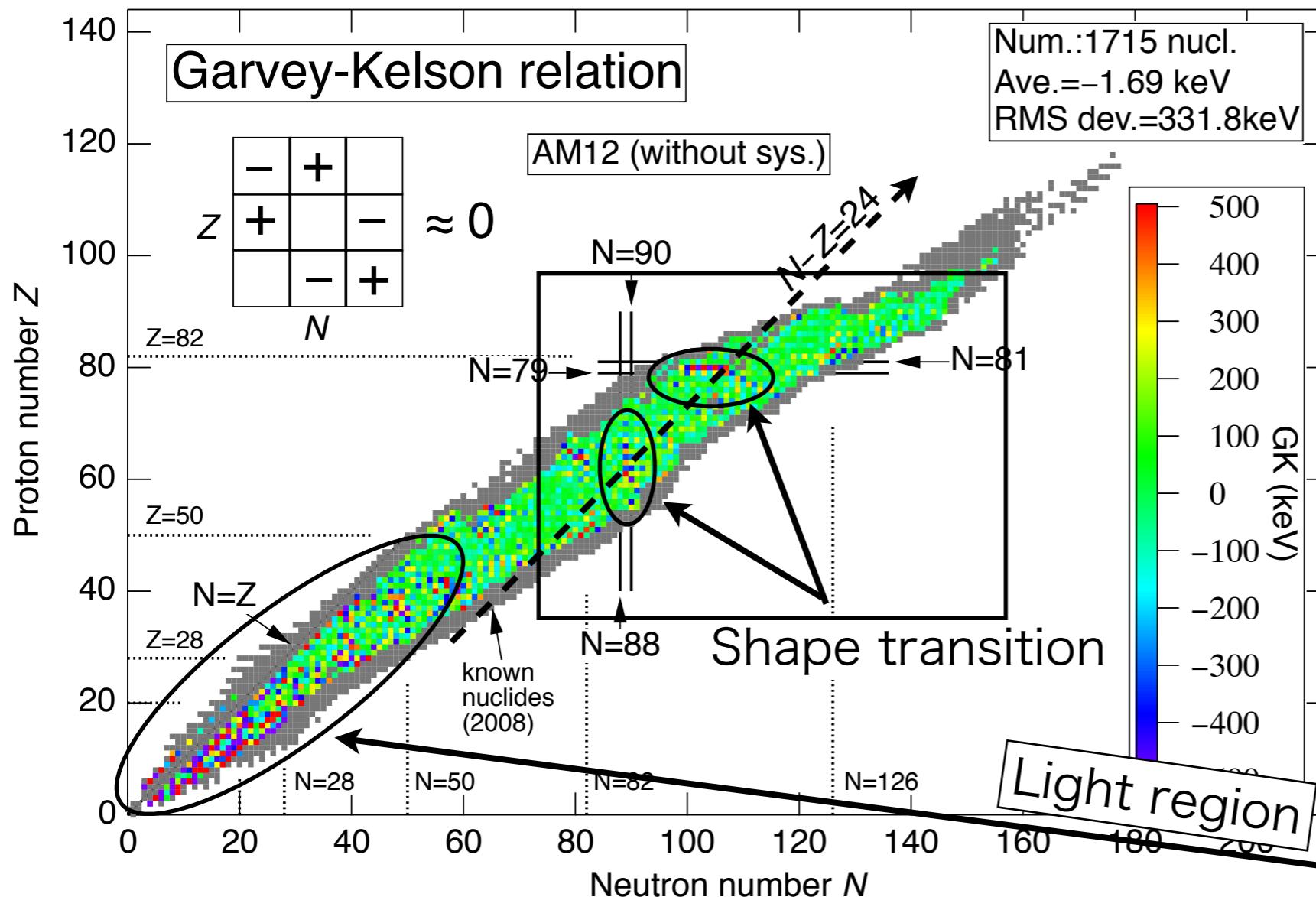
A consideration of cancellation of core + valence nucleons (based on the shell model)



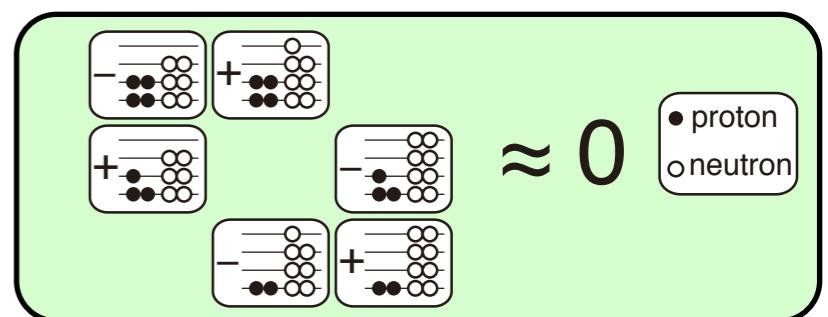
- Assumption: Cores among related six nuclei are the same.

Region	num.	Average	RMS dev.
All	1715	-1.7 (keV)	331.8(keV)
$A > 100$	1202	-0.03	161.2
$A \leq 100$	513	-29.3	554.1

larger



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larger

GK type mass formula:

Comay-Kelson-Zidon (CKZ₁₉₈₈)
Jänecke-Masson (JM₁₉₈₈)
Masson-Jänecke(MJ₁₉₈₈)
(ADNDT39, 1988)

In heavy region, the GK sys. gives some transition as changes of structures, especially nuclear shape.

III. Mass formulae

KTUY mass model

KUTY (KTUY) mass formula (Koura,Uno,Tachibana,Yamada (00,05))

$$M(Z, N) = M_{\text{gross}}(Z, N) + M_{\text{eo}}(Z, N) + M_{\text{shell}}(Z, N)$$

$M_{\text{gross}}(Z, N)$: gross term

H. K, et al., Nucl. Phys. A **674**, 47 (2000)
H. K, et al., Prog. Theor. Phys. **113**, 305 (2005)

$$M_{\text{gross}}(Z, N) = M_{\text{H}} Z + M_{\text{n}} N + a(A) A + b(A) |N - Z| + c(A)(N - Z)^2/A + E_C(Z, N) - k_{\text{el}} Z^{2.39}$$
$$a(A) = a_1 + a_2 A^{-1/3} + a_3 A^{-2/3} + a_4 (A + \alpha_a)^{-1}$$
$$b(A) = \quad \quad \quad + b_4 (A + \alpha_b)^{-1}$$
$$c(A) = c_1 + c_2 A^{-1/3} + c_3 A^{-2/3} + c_4 (A + \alpha_c)^{-1}$$

$M_{\text{shell}}(Z, N)$: shell term

Spherical nuclei

Calculated from **Spherical single-particle potential**
for any nuclei (includes the BCS paring, reduction)

Deformed nuclei (Spherical-basis condieration)

Obtained by **an appropriate mixture** of the above
spherical shell energies + liquid-drop deform. energies

KUTY mass formula

(Koura, Uno, Tachibana, Yamada (2000))

$$M(Z, N) = M_{\text{gross}}(Z, N) + M_{\text{eo}}(Z, N) + \color{red}M_{\text{shell}}(Z, N)$$

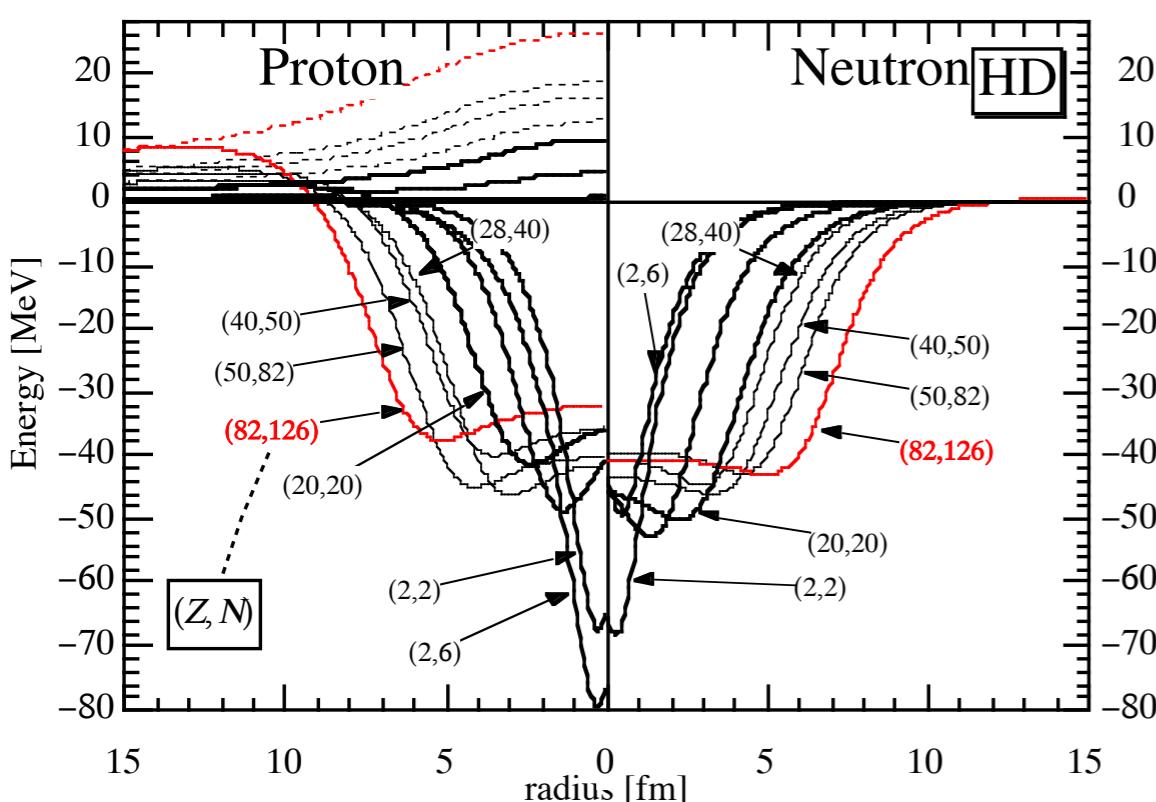
- $M_{\text{shell}}(Z, N)$: shell term

H. K. and M. Yamada, Nucl. Phys. A 671, 96 (2000)

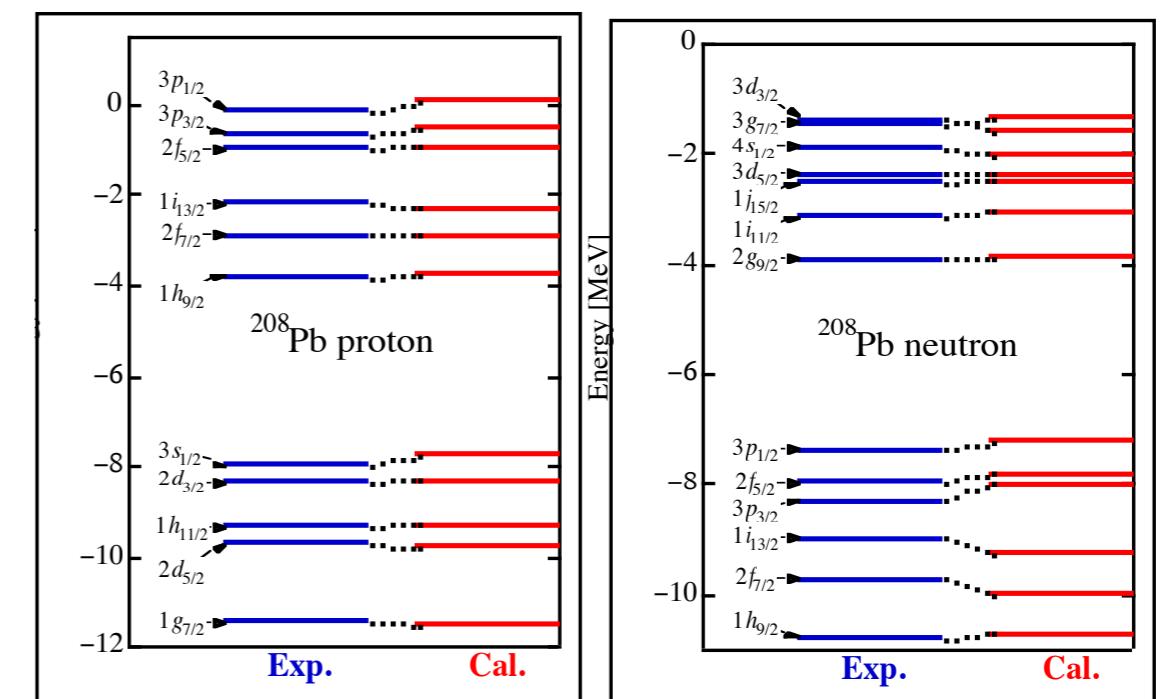
Spherical nuclei

calculated from **Spherical single-particle potential**
for any nuclei (includes the BCS paring, reduction)

$$V_{\text{cen}}(r) = V_0 \frac{1}{\left\{1 + \exp \left[(r - R_v)/a_v\right]\right\}^{a_v/\kappa}} \left\{1 + V_{\text{dp}} \frac{1}{1 + \exp [-(r - R_v)/a_v]}\right\} \quad (\text{Central component})$$



Single-particle levels of ^{208}Pb



Differences between calc. and exp. levels: within 330 keV for $^{132}\text{Sn}, ^{208}\text{Pb}$

Nuclear shell energy with deformation

(Spherical-basis consideration)

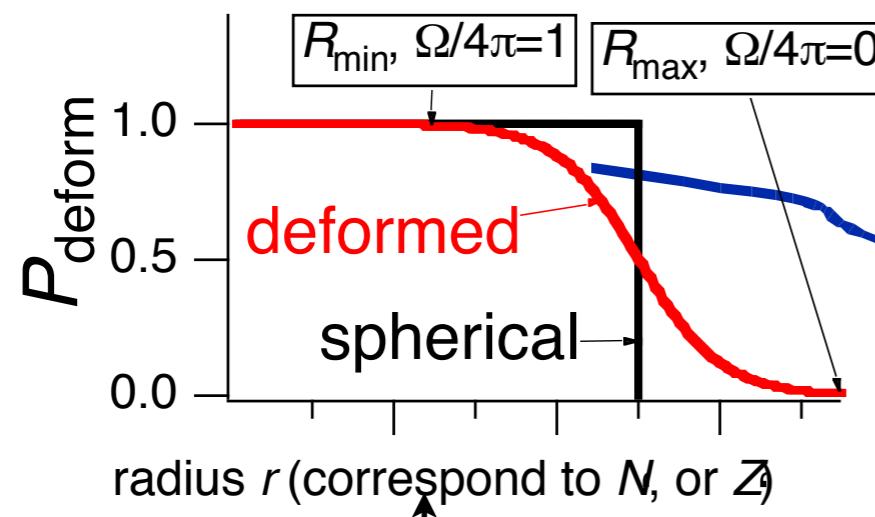
H. K, et al., Nucl. Phys. A 674, 47 (2000)

$$E_{\text{sh}}(Z, N) = \sum_{\text{def.}} \frac{\langle E_{\text{sh}}^{\text{sph}}(Z, N) \rangle_{\text{def.}}}{\text{micro.}} + \frac{\Delta E_S(Z, N) - \Delta E_C(Z, N) - \Delta E_{\text{pro}}(Z, N)}{\text{Liquid-drop}}$$

$$\frac{\langle E_{\text{sh}}^{\text{sph}}(Z, N) \rangle_{\text{def.}}}{\text{micro.}} = \sum_{Z'} W_{\text{pdef}}(Z'; Z, N) E_{\text{psph}}(Z'.N'') + \sum_{N'} W_{\text{ndef}}(N'; Z, N) E_{\text{nsph}}(Z''.N')$$

Spherical shell energy

Mixing weight W_{def}

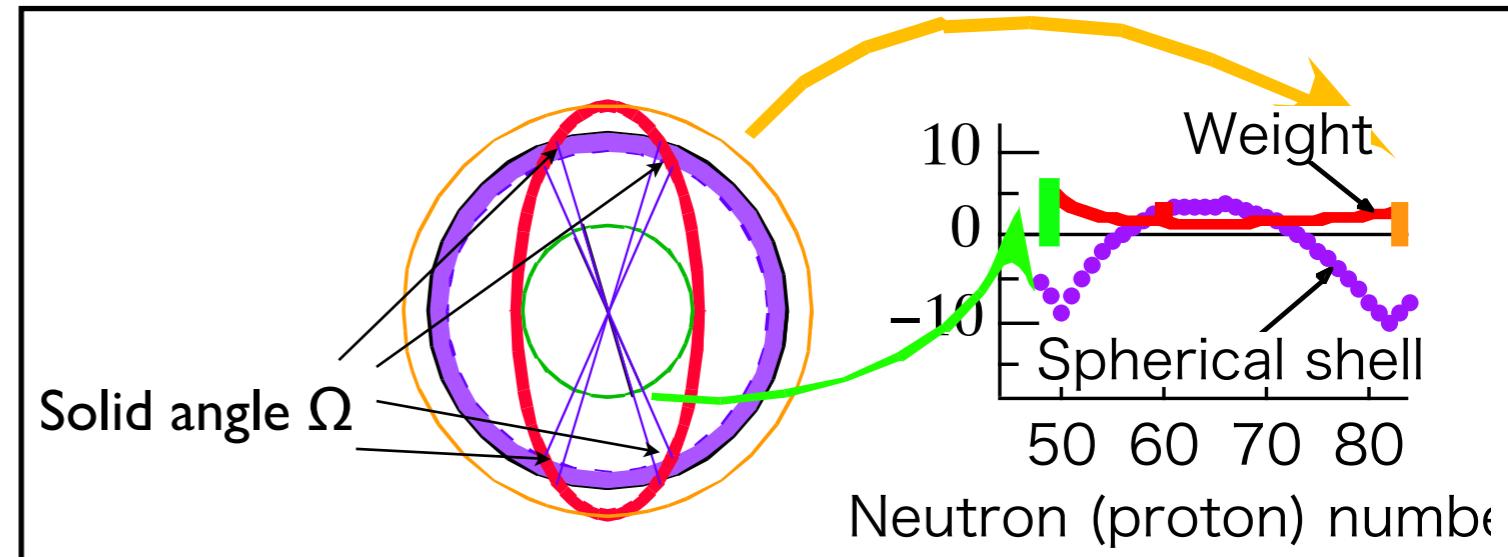


Current calc.:
 $\alpha_{2,4,6}$ deformation
(not asym. shape)

P_{deform} : occupation probability of deformation state on the spherical basis

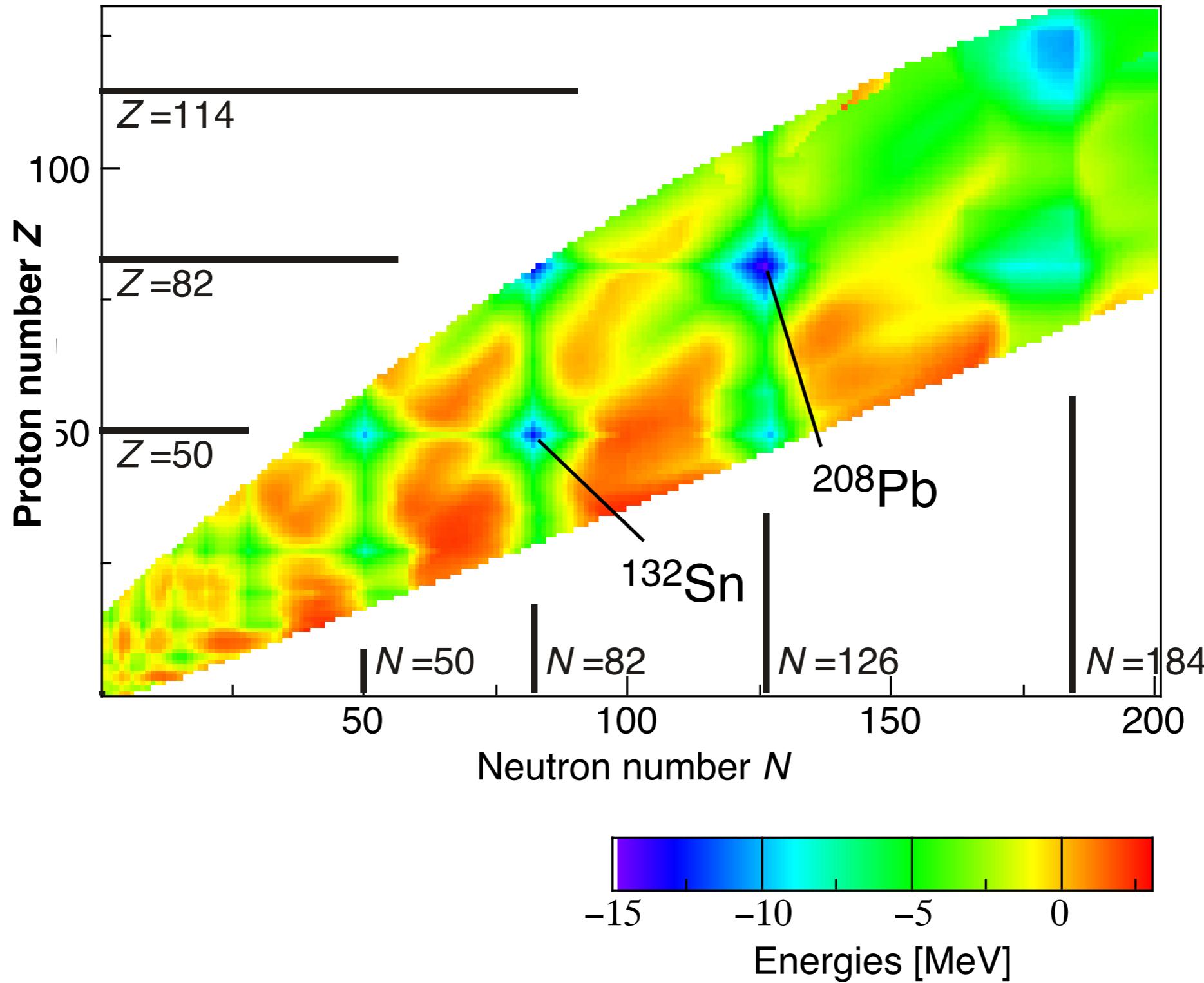
$$W_{\text{ndef}}(N'; Z, N) = \frac{1}{4\pi} \frac{d\Omega(r(N'))}{dN'}$$

Differential



Nuclear shell energy of KUTY

Nuclear shell energies $E_{\text{sh}}(Z, N)$



Other mass models and comparison

Skyrme-Hartree-Fock-Bogoliubov mass formula (2002-)

by S. Goriely et al.

$$E_{\text{tot}} = E_{\text{HFB}} + E_{\text{wigner}}$$

BSk21 force parameter set:

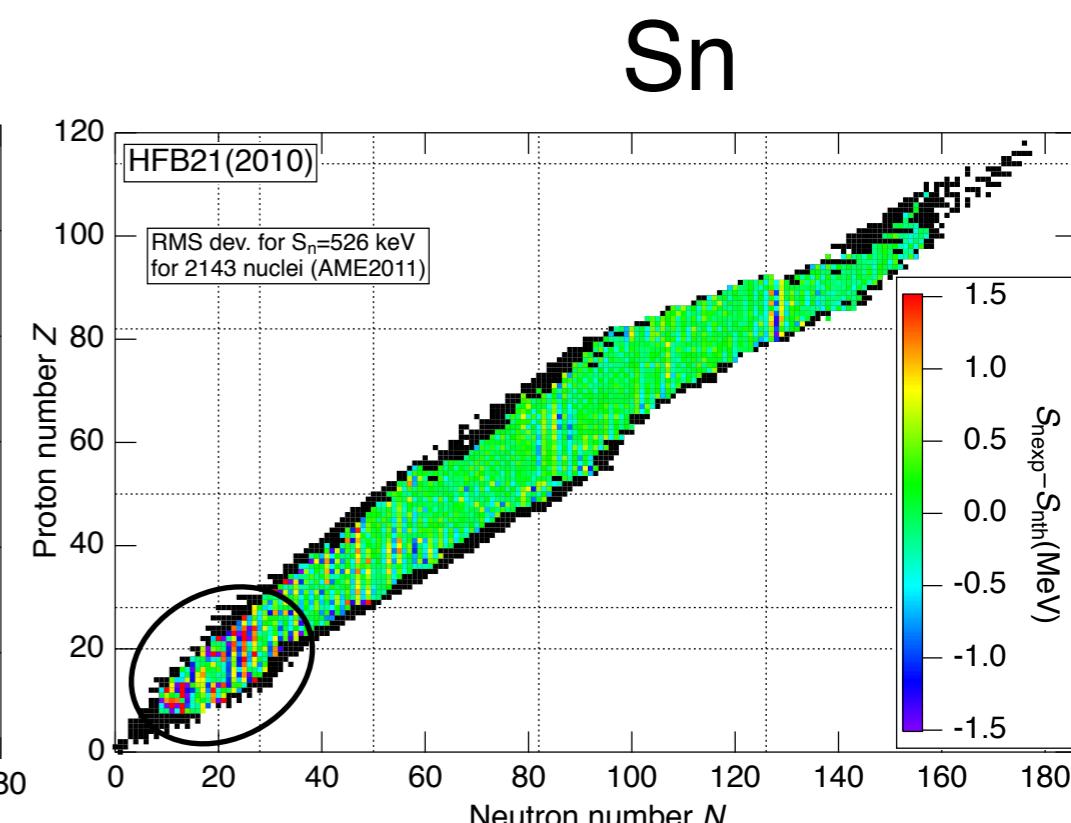
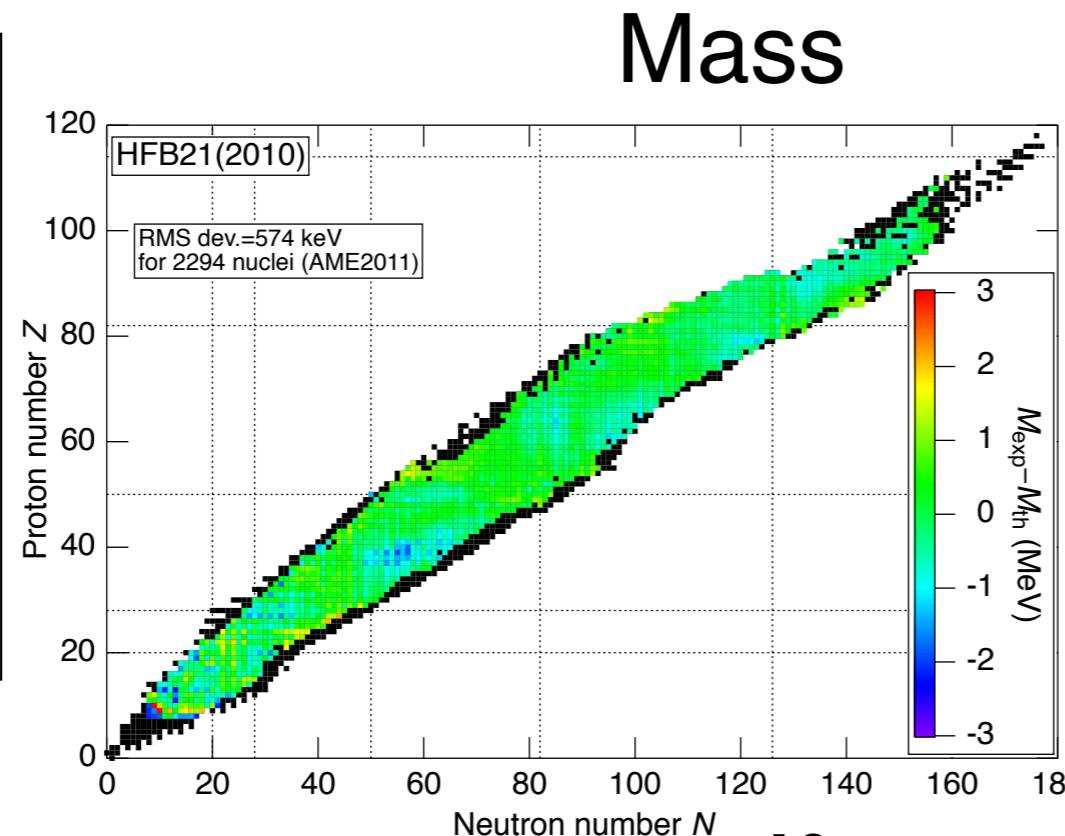
$t_0 = -3961.39 \text{ MeV fm}^3$, $t_1 = 396.131 \text{ MeV fm}^5$
 $t_2 = 0 \text{ MeV fm}^5$, $t_3 = 22588.2 \text{ MeV fm}^{3+3}\alpha$
 $t_4 = -100.000 \text{ MeV fm}^{5+3}\beta$, $t_5 = -150.000 \text{ MeV fm}^{5+3}\gamma$
 $x_0 = 0.885231$, $x_1 = -0.0648452$, $t_2 x_2 = 1390.38 \text{ MeV fm}^5$
 $x_3 = 1.03928$, $x_4 = 2.00000$, $x_5 = -11.0000$
 $W_0 = 109.622 \text{ MeV fm}^5$, $\alpha = 1/12$, $\beta = 1/2$, $\gamma = 1/12$
 $f_n^+ = 1.00$, $f_p^+ = 1.07$, $f_n^- = 1.05$, $f_p^- = 1.13$
 $V_W = -1.80 \text{ MeV}$, $\lambda = 280$, $V'_W = 0.96$, $A_0 = 24$

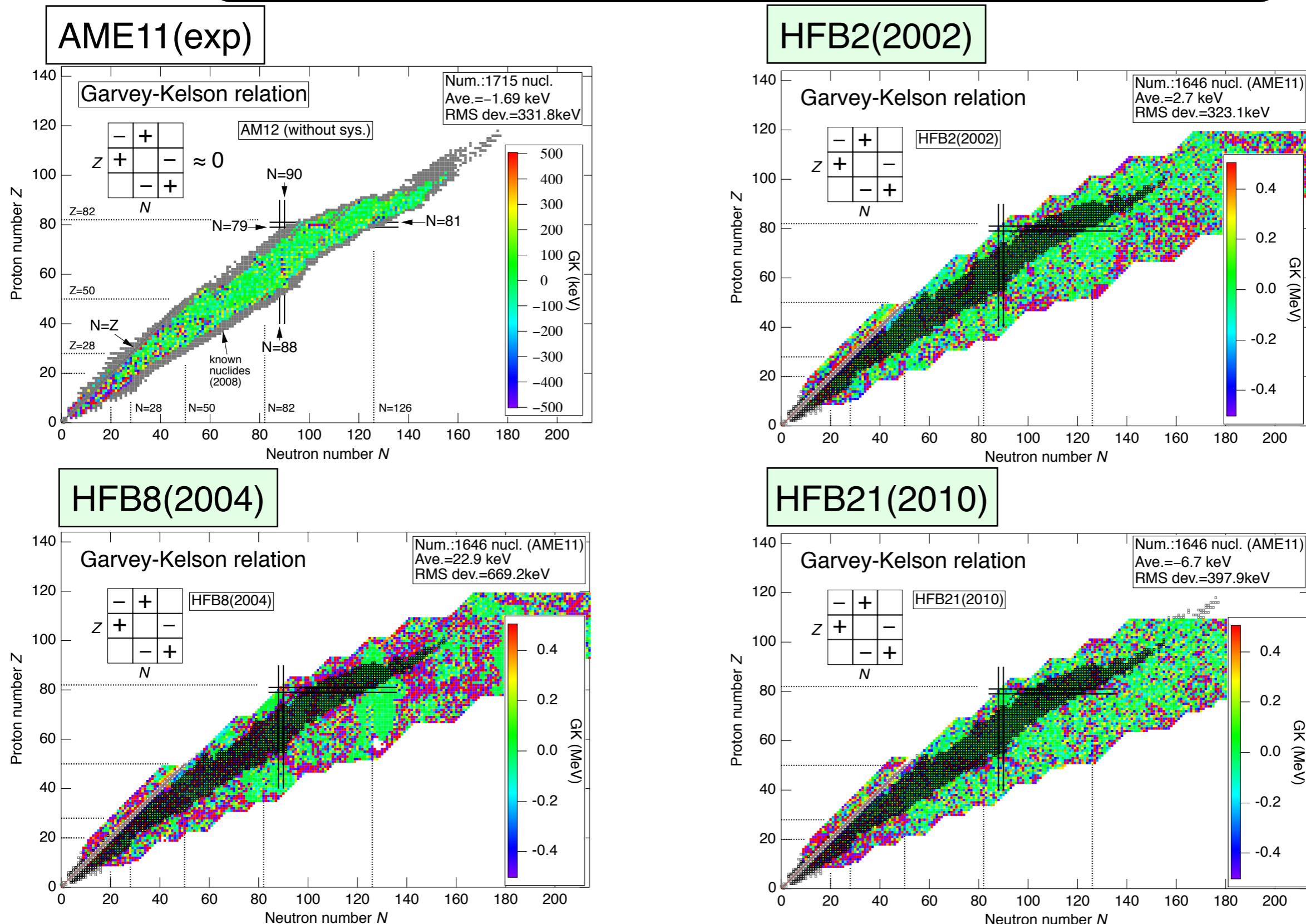
The long road in the HFB mass model development

	HFB-2 :	Possible to fit all 2149 exp masses $Z \geq 8$	659 keV
	HFB-3:	Volume versus surface pairing	635 keV
	HFB-4-5:	Nuclear matter EoS: $M_s^* = 0.92$	660 keV
	HFB-6-7:	Nuclear matter EoS: $M_s^* = 0.80$	657 keV
	HFB-8:	Particle-number projection	635 keV
	HFB-9:	Neutron matter EoS	733 keV
	HFB-10-13:	Low pairing & NLD	717 keV
	HFB-14:	Collective correction and Fission B_f	729 keV
	HFB-15:	Coulomb correlations / CSB	678 keV
	HFB-16:	Pairing constrained to NM	632 keV

HFB21 gives a less than 600 keV of the RMS dev. In the light region there is some discrepancy in derivatives as S_n .

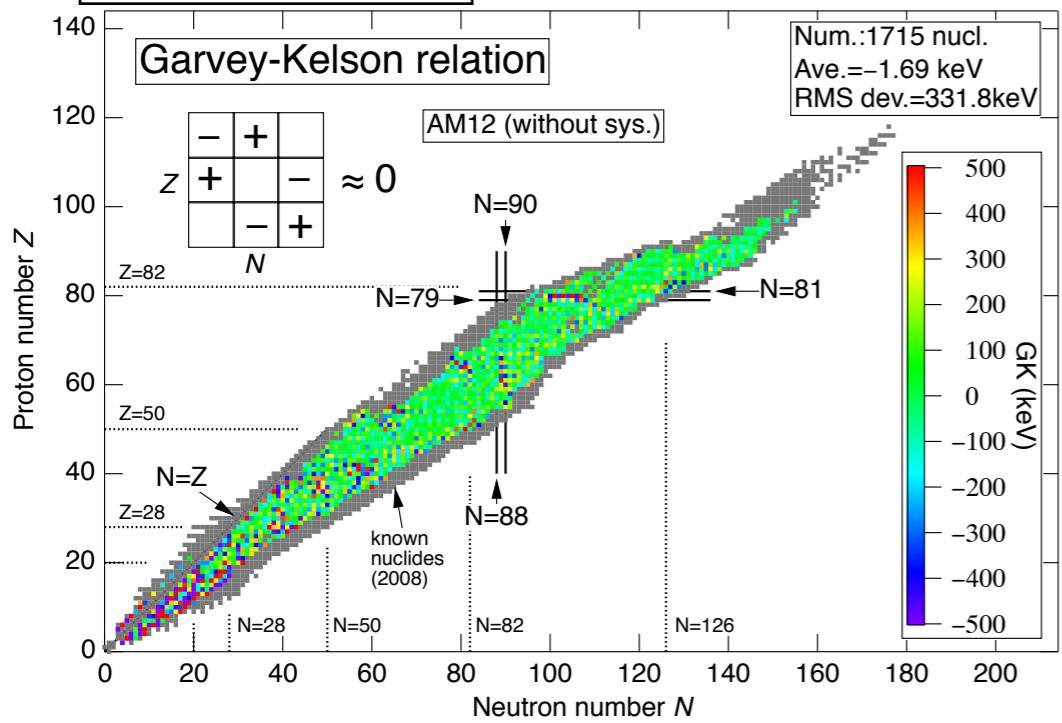
Referred mass data:AME03



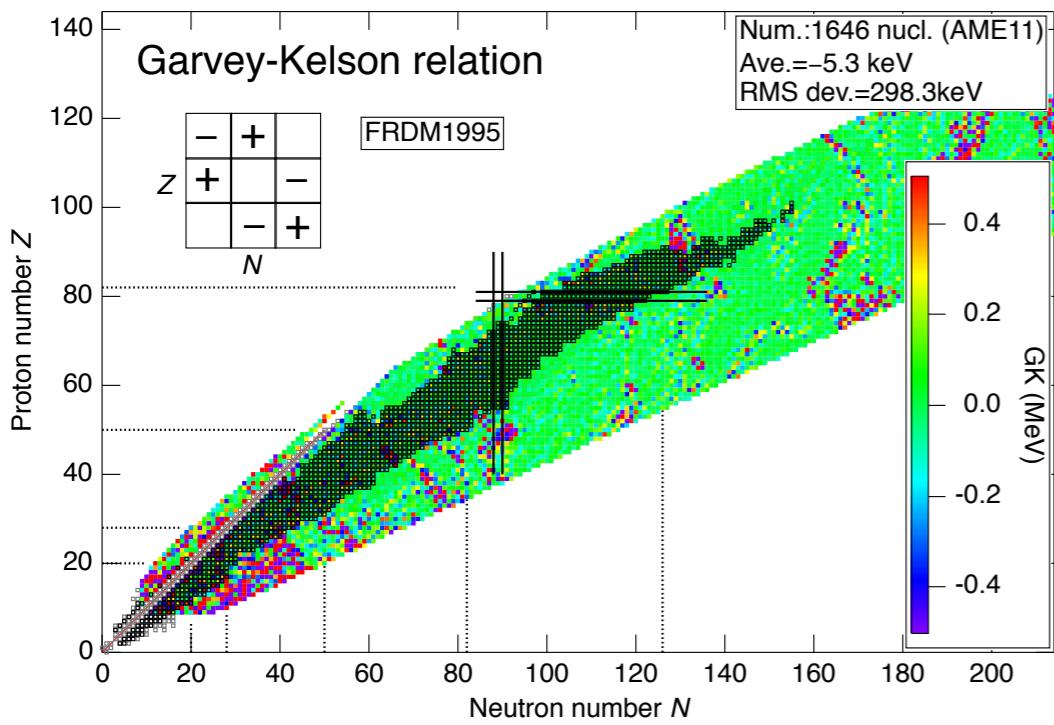


- The HFB calc. seems to violate the GK summation.

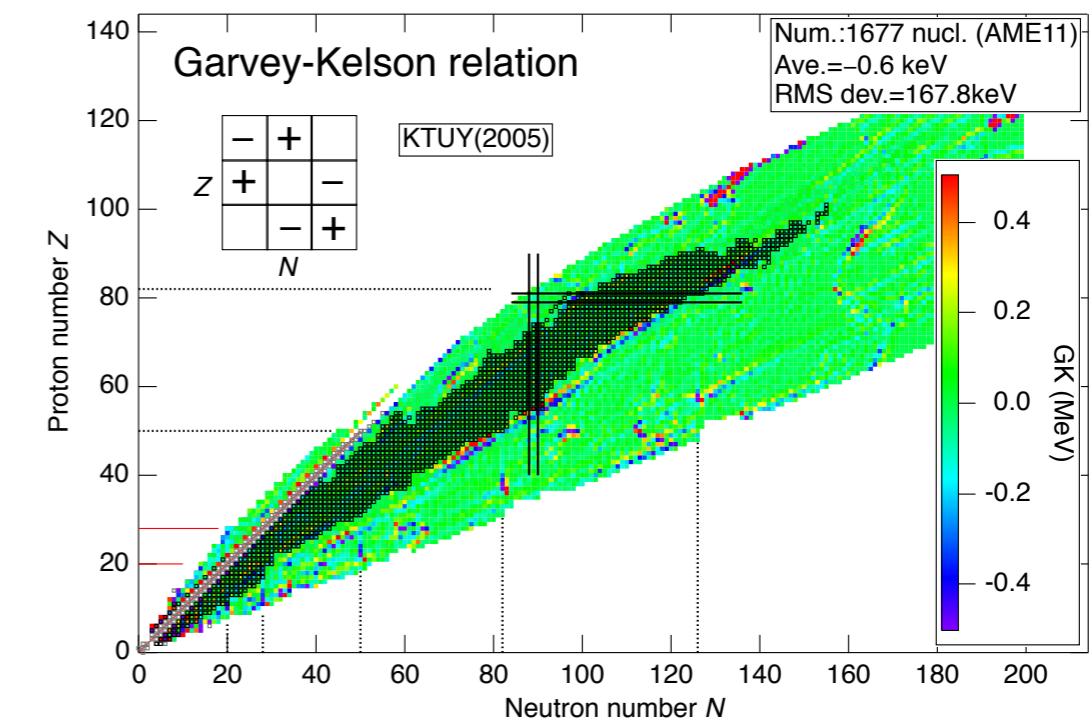
AME11(exp)



FRDM1995 : gives a smooth trends

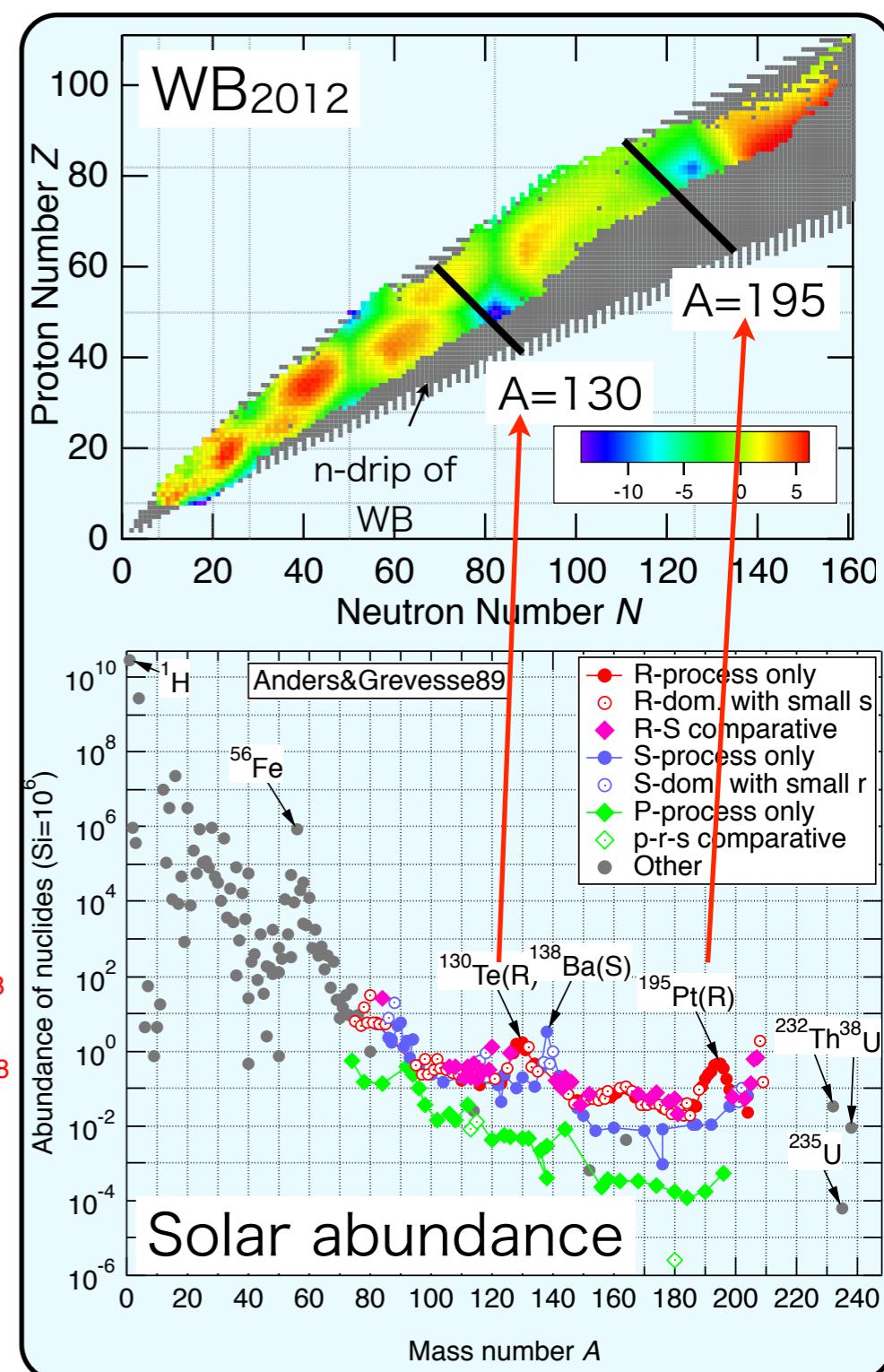
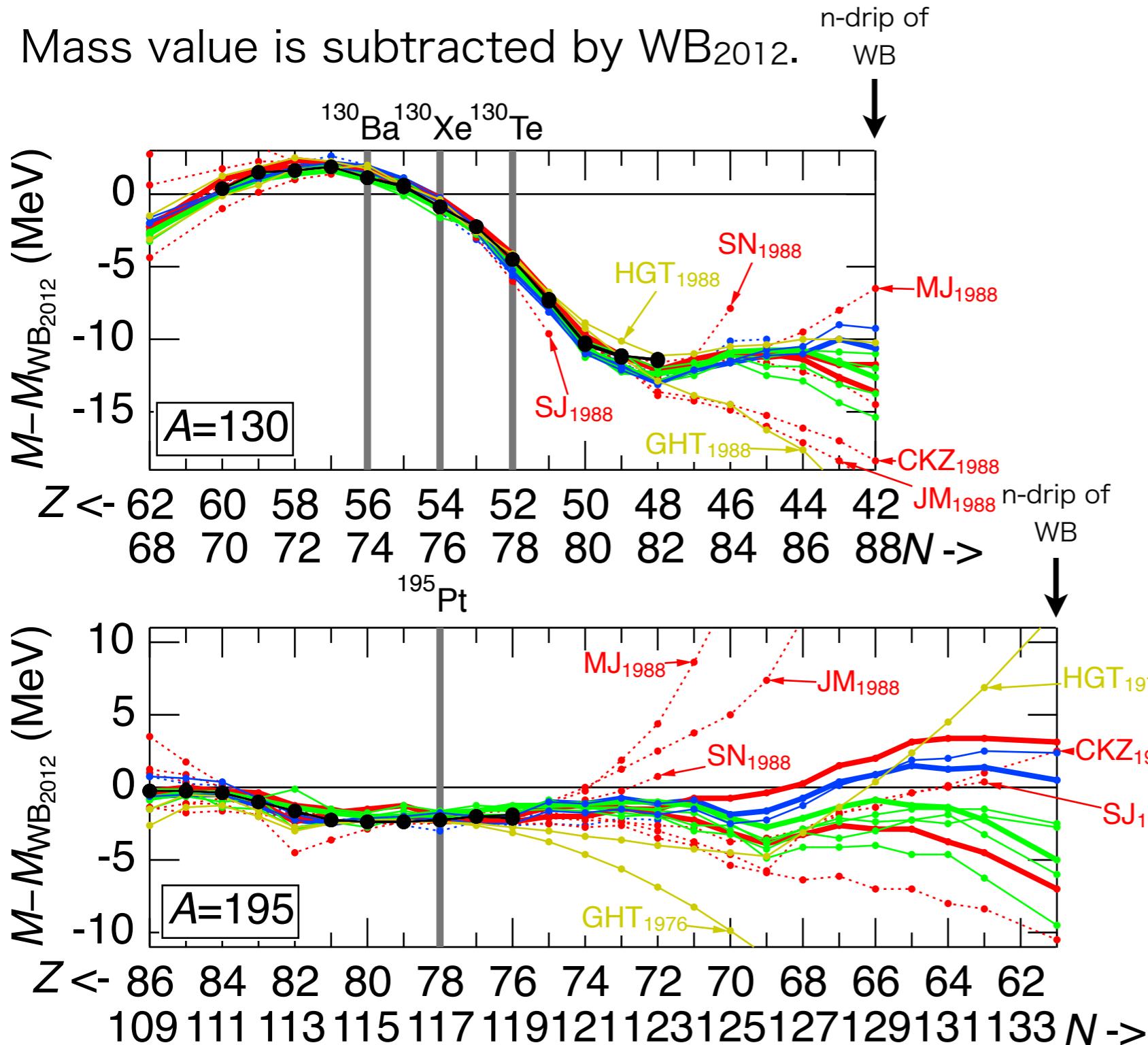


KTUY2005 : gives a smooth trends



- FRDM and KTUY give a smooth trends.

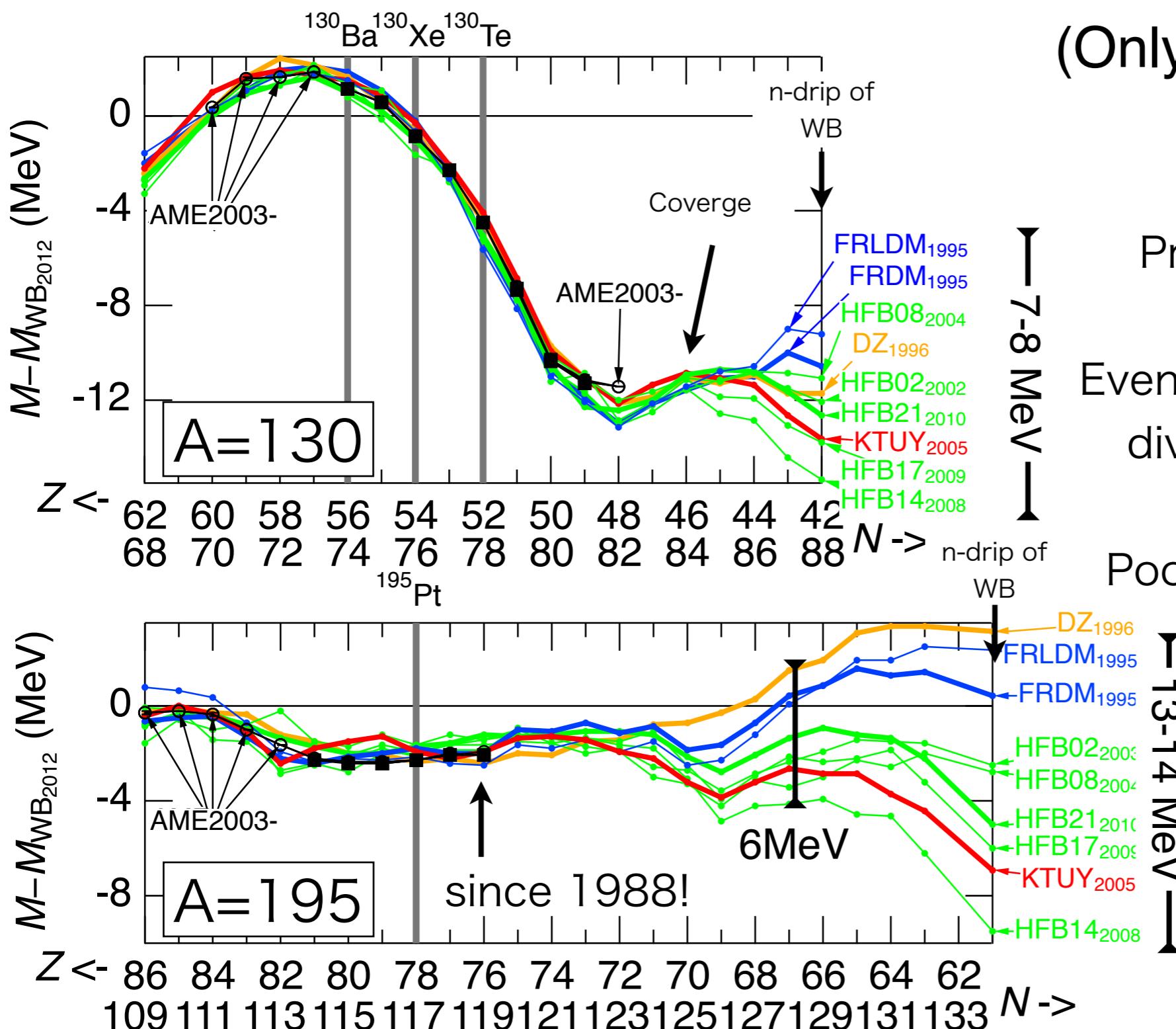
Extrapolation to the n-rich nuclei



- In old-type mass formulae (-1988), mass values extremely diverge in the very neutron-rich region

Extrapolation to the n-rich nuclei

(Only since 1990-)



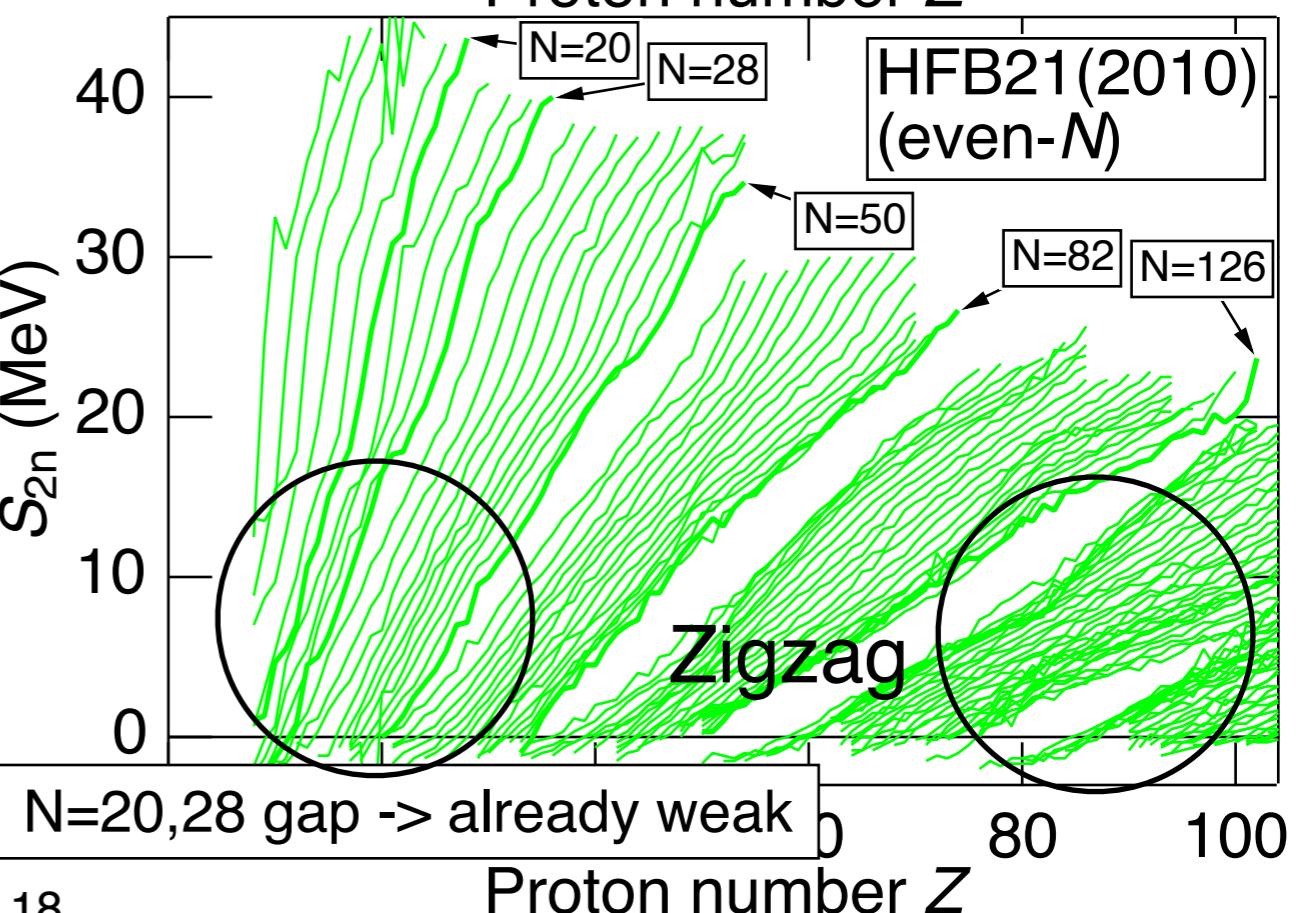
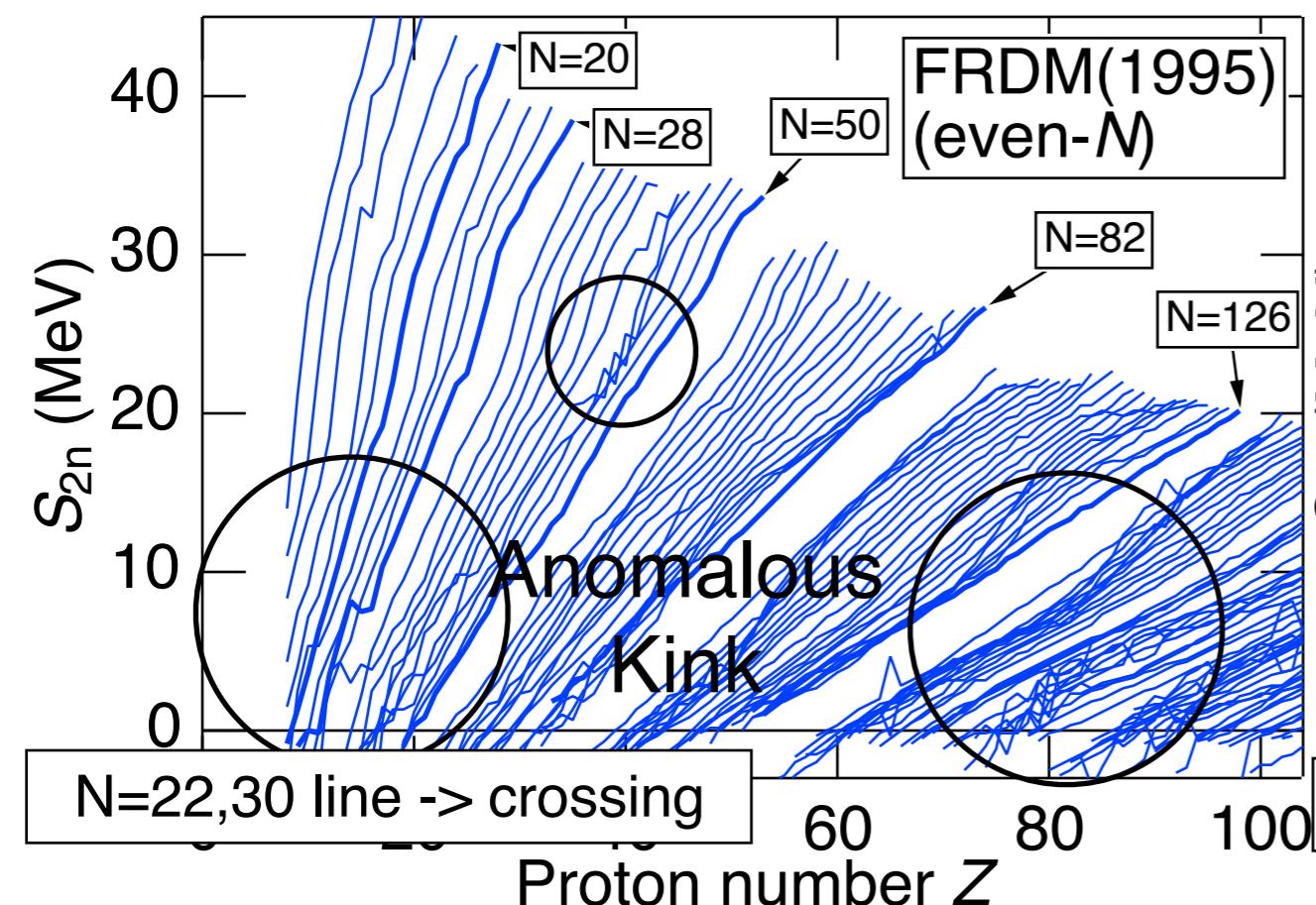
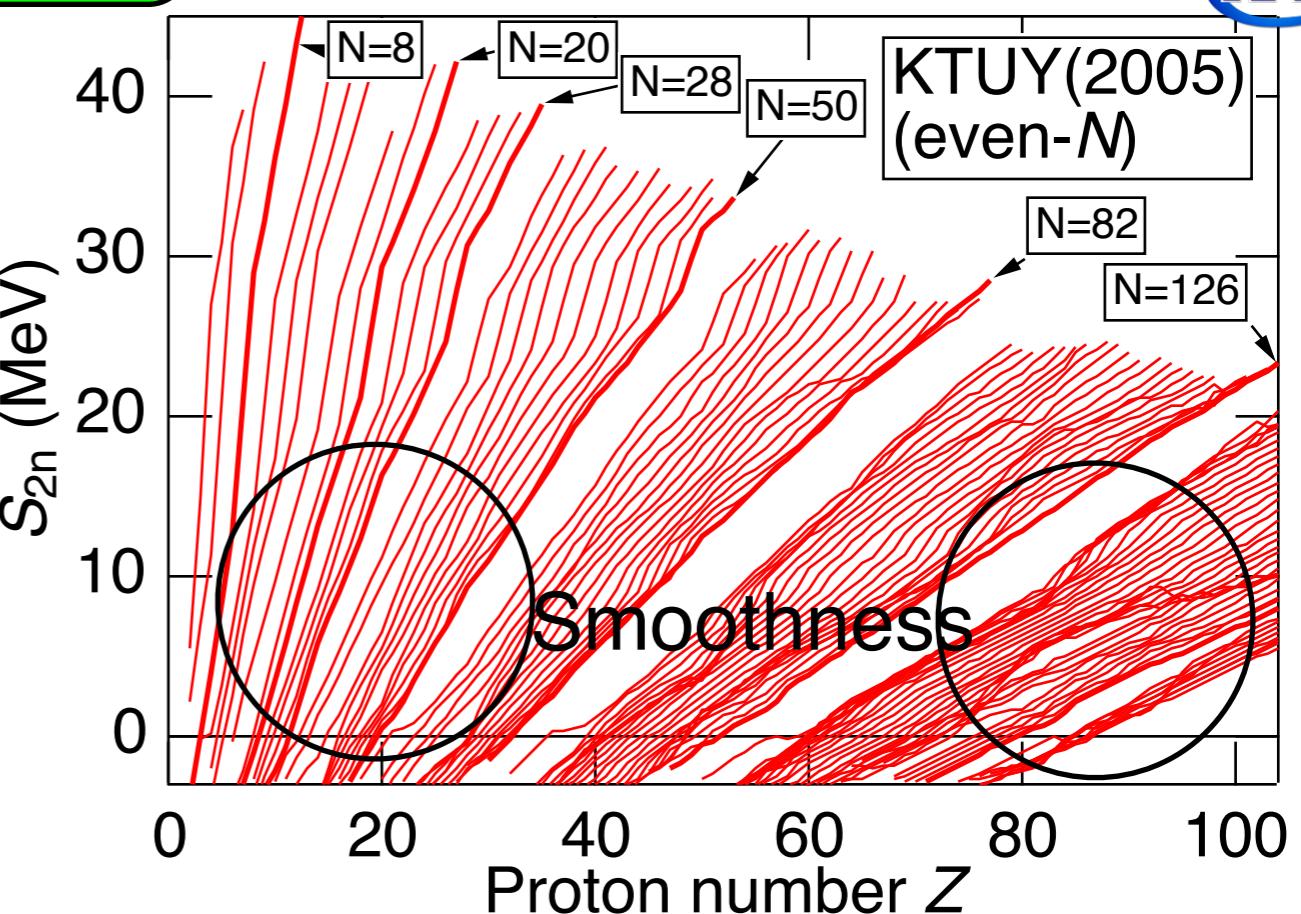
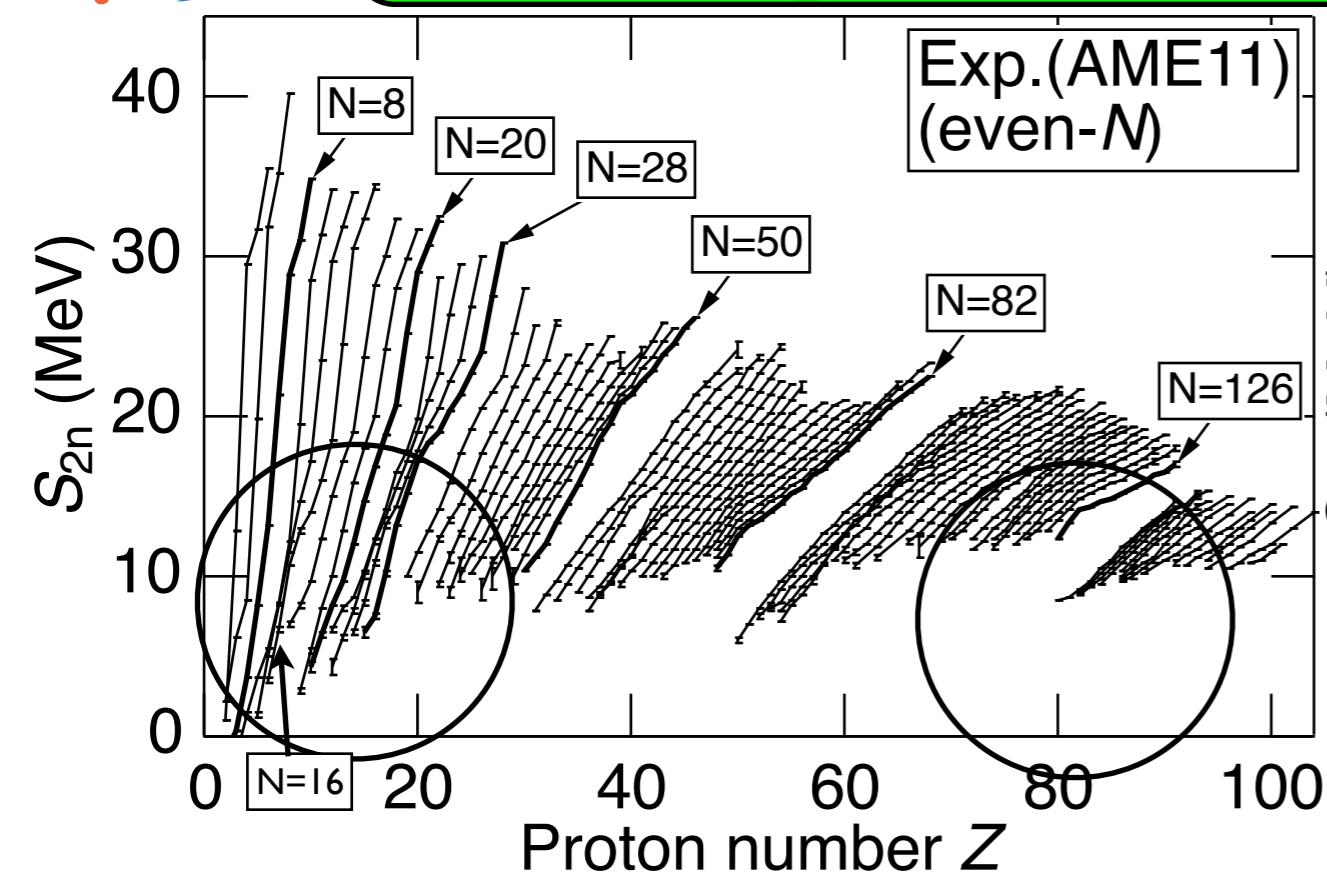
Predicted mass values still diverge.

Even among HFB's, mass values diverge in the n-rich region.
(several MeV)

Poor experimental mass data.

S_{2n} systematics

shell gap and smoothness



IV. Application to r-process

Mass-model dependency

-Check the mass formulae as astrophysical data-

- Canonical model

Steady flow + Waiting point Approximation

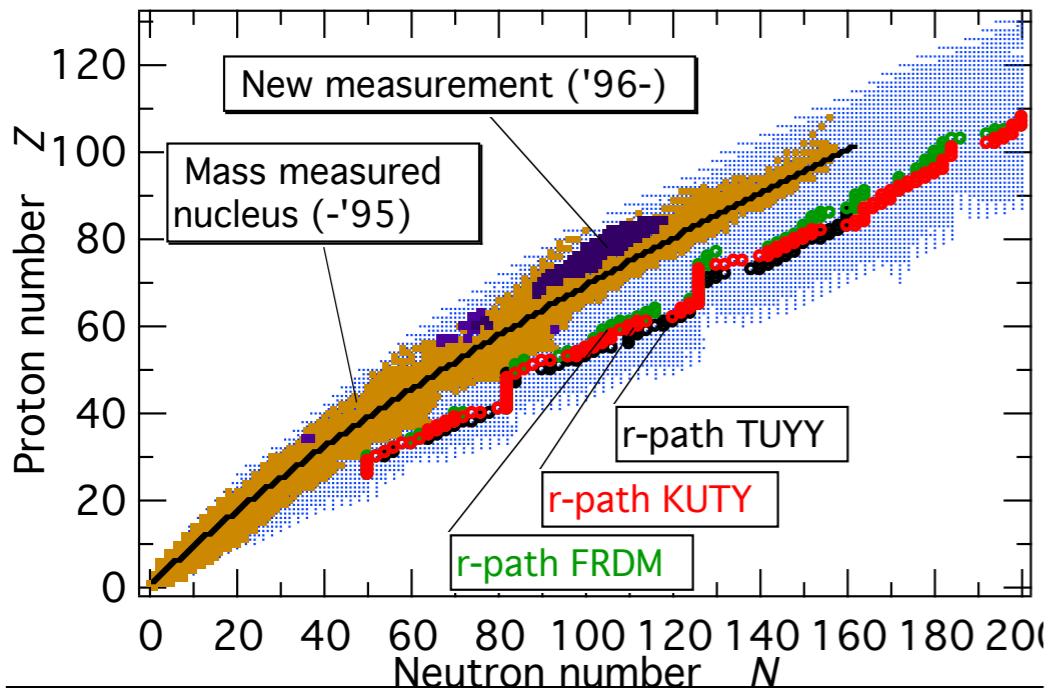
Neutron-number density (N_n) and temperature (T_9) are constants

(n,γ) - (γ,n) equilibrium is established over an irradiation time τ

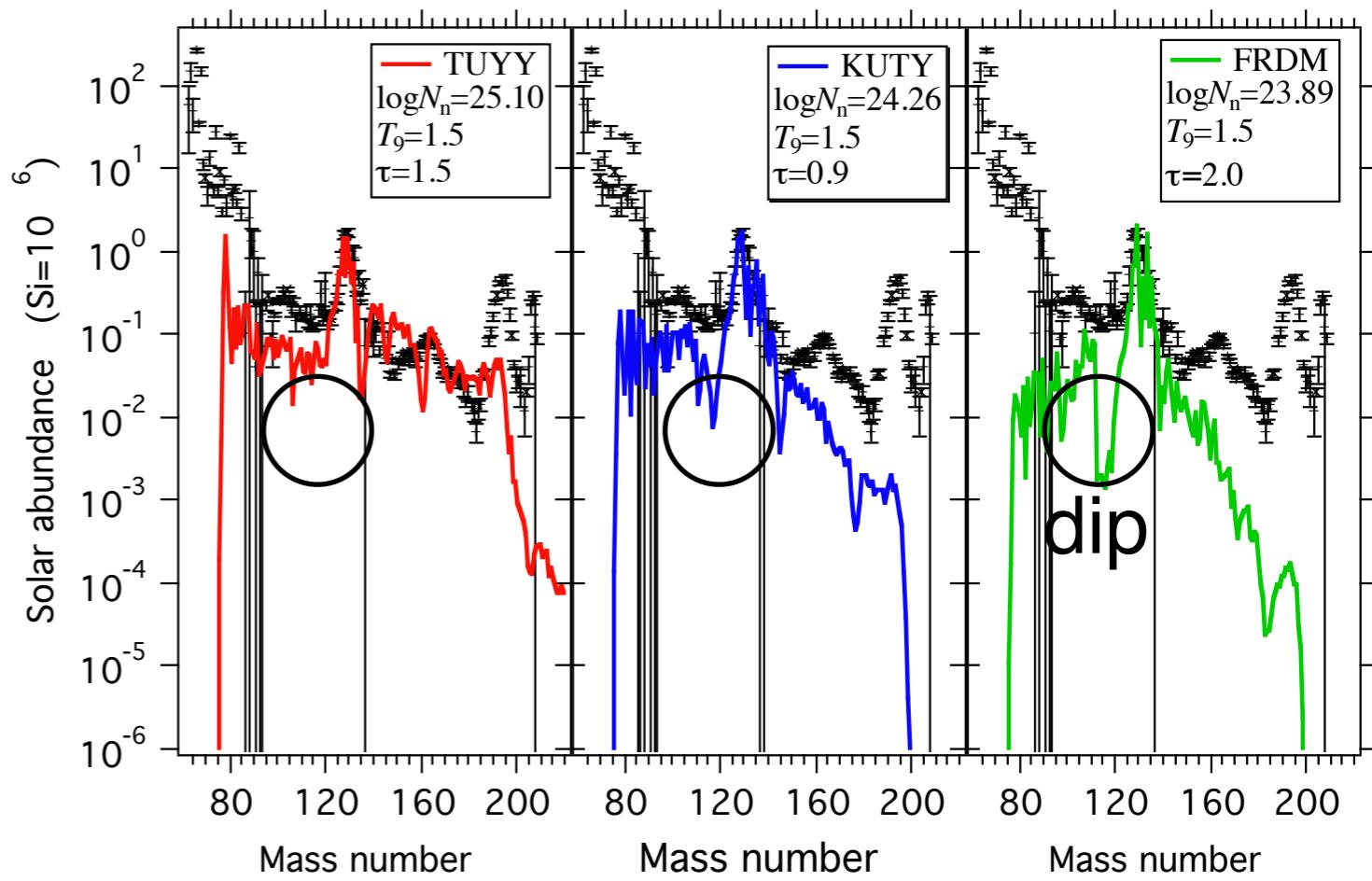
N_n, T_9, τ : chosen to reproduce the abundance peak at $A=130$ (obs.)

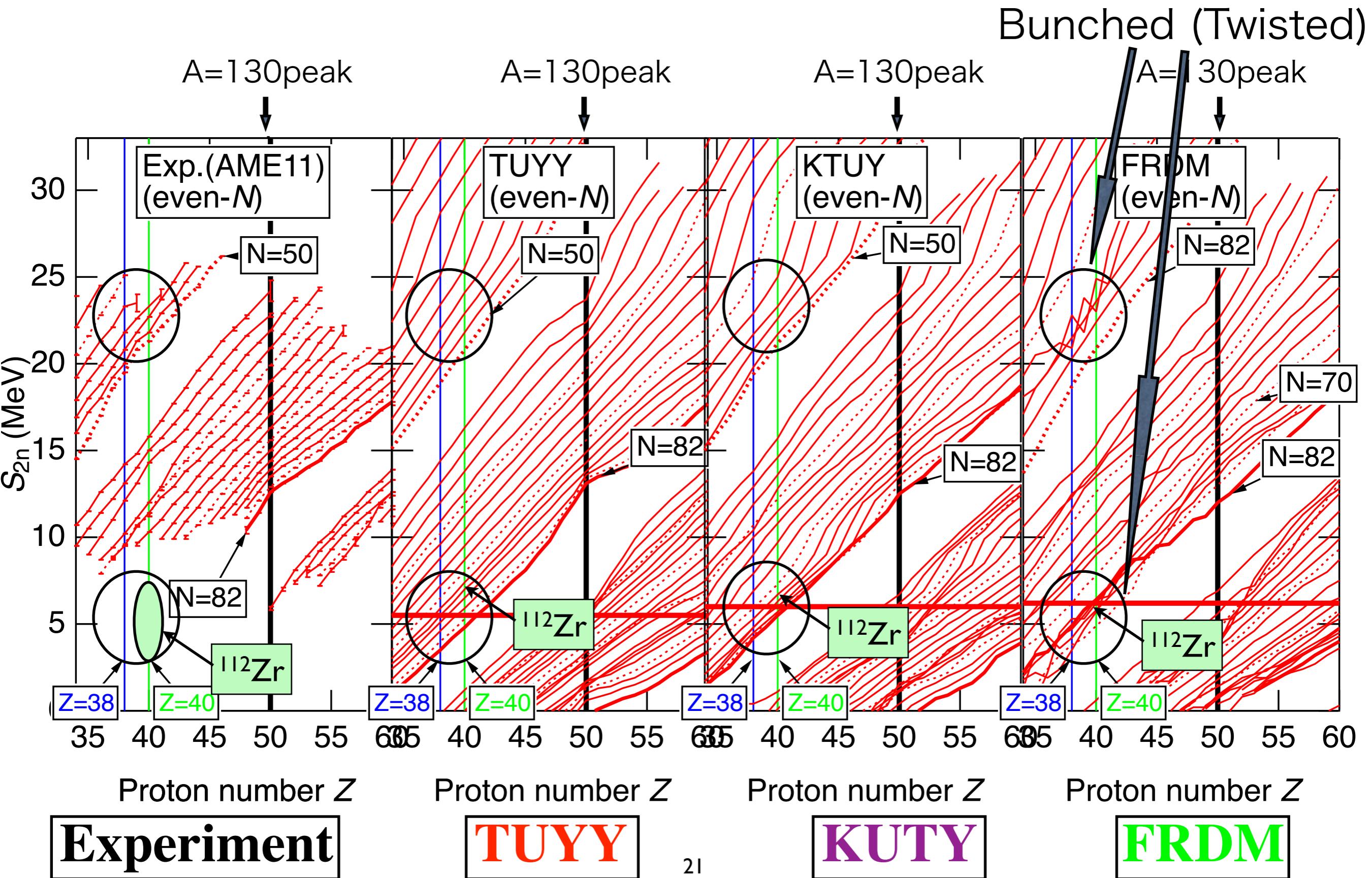
S_{2n} for equilibrium eq. (determine the path) and Q_β for λ_β :
estimated from mass formulae (TUYY, KUTY, FRDM)

+ $N_r = N(\text{Solar abund.}) - N_s$
 × N_r r-only nuclei

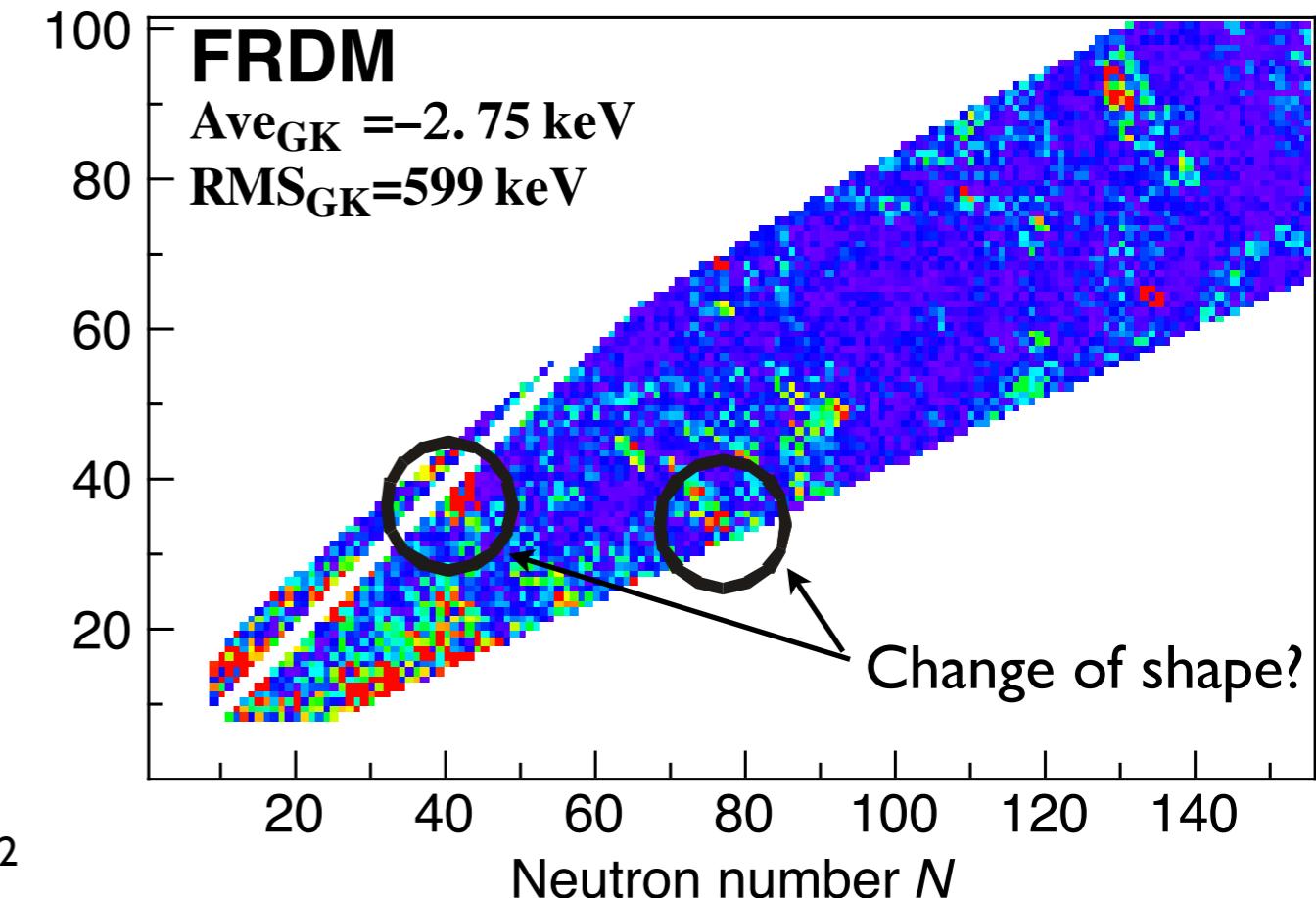
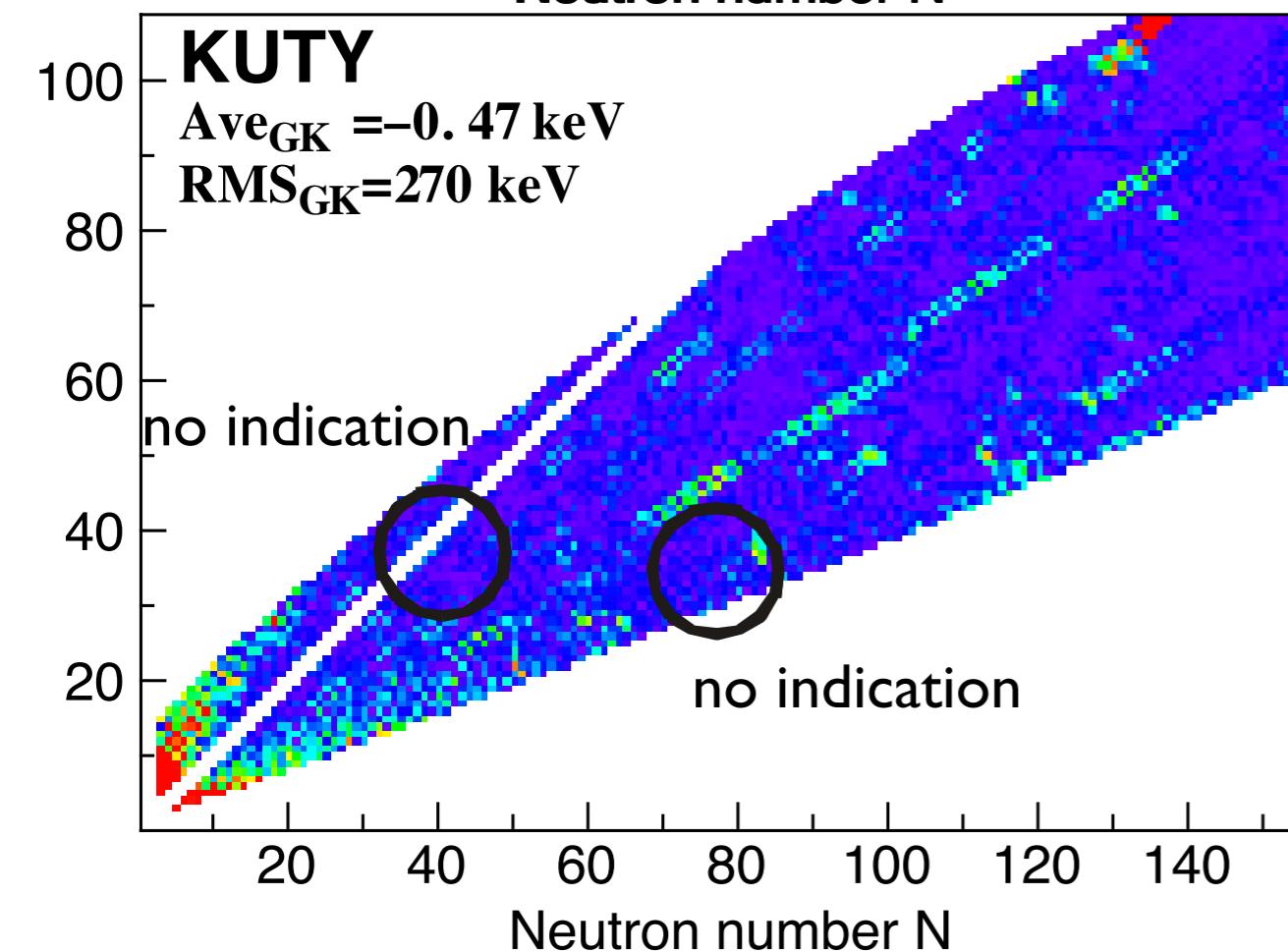
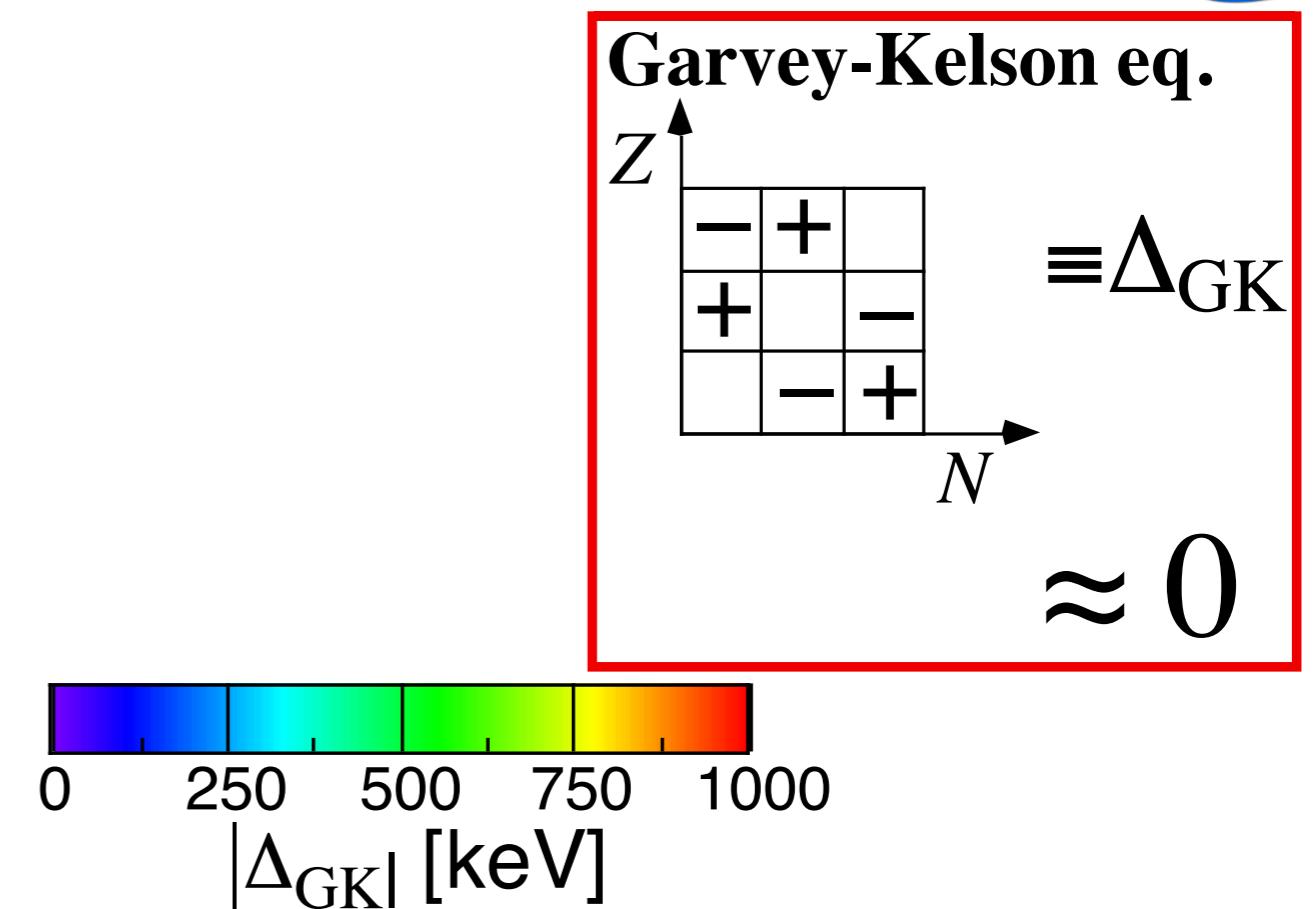
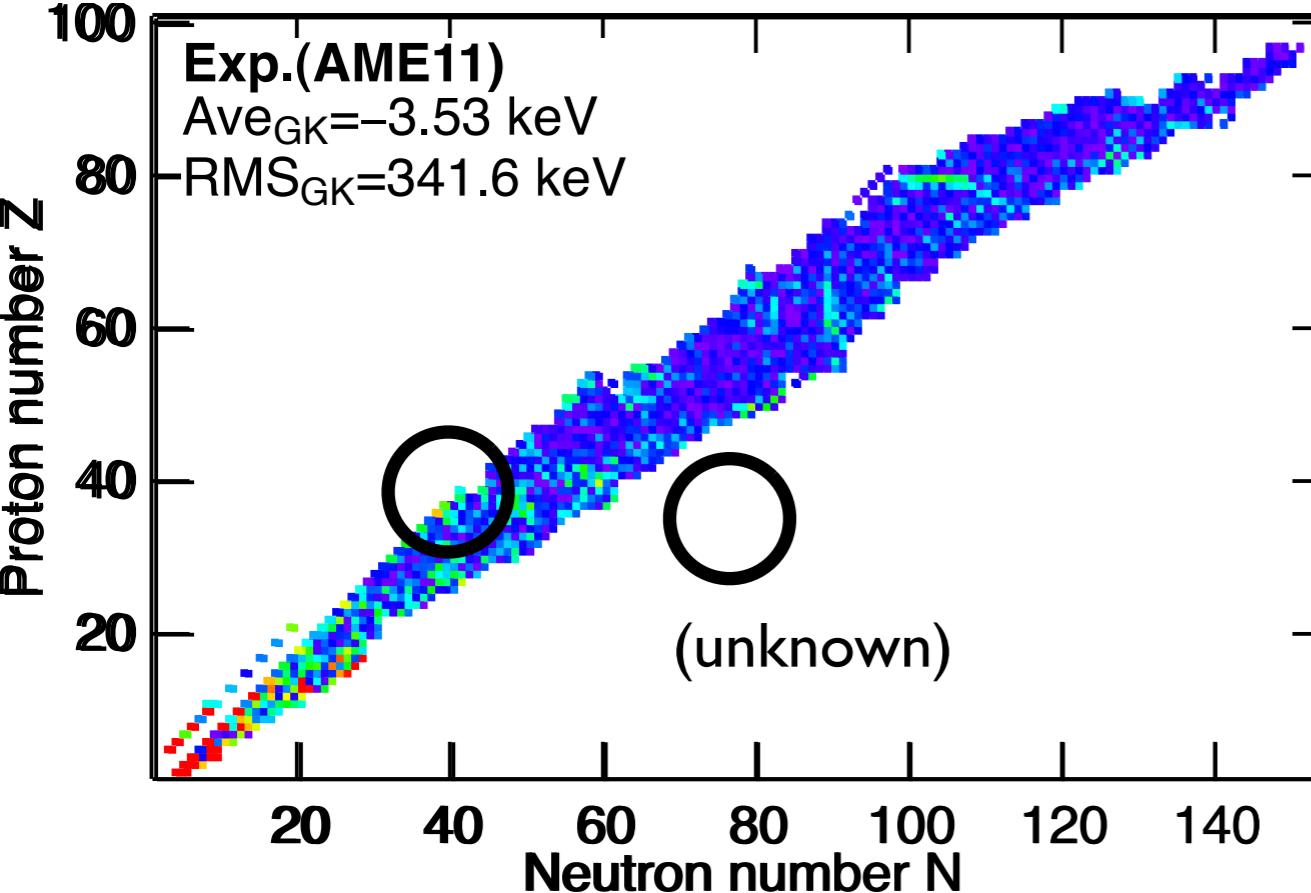


- TUYY: gross term (WB-like with higher expansion) + empirical shell term.
- KTUY: TUYY gross term + deformed shell with a modified Woods-Saxon pot.
- FRDM: Macroscopic Droplet + microscopic deformed shell with a folded Yukawa pot.



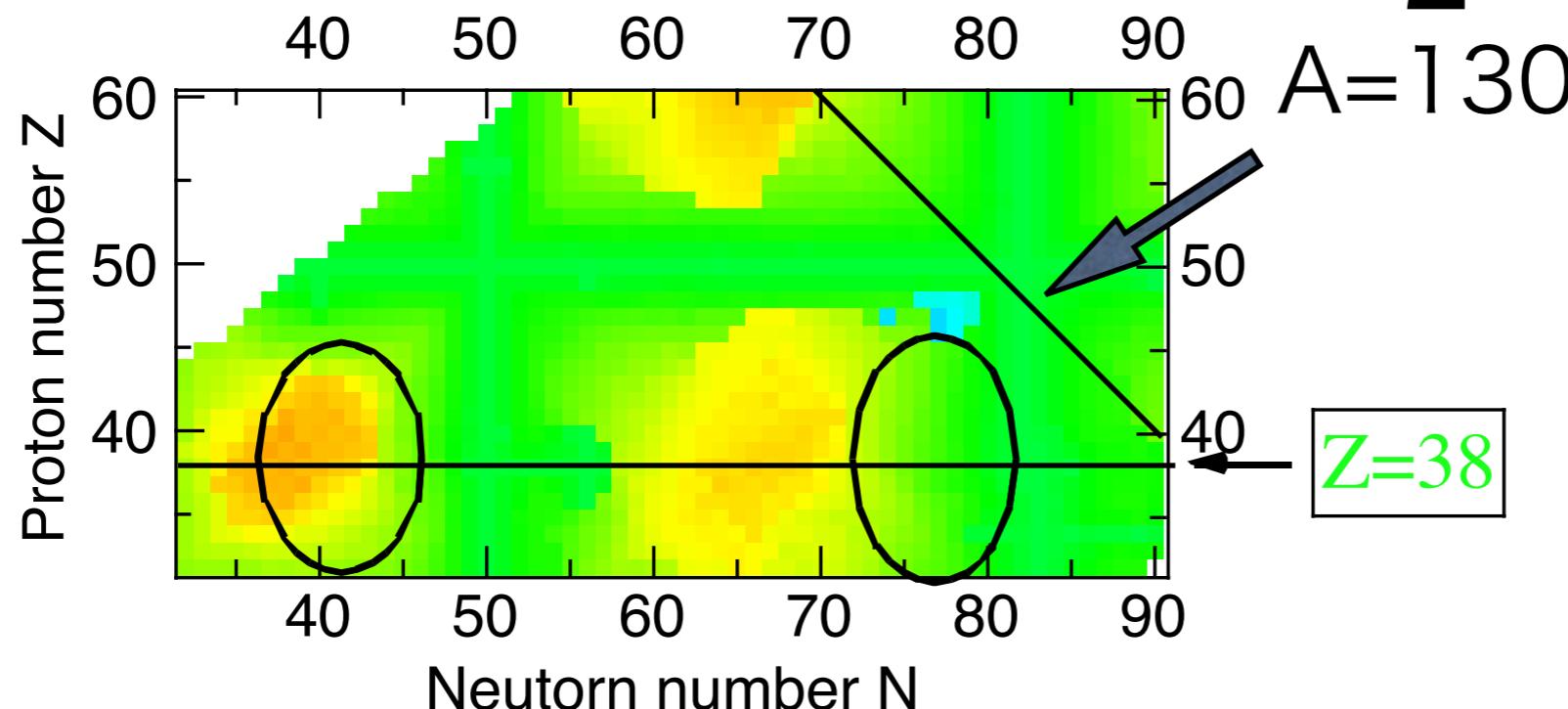


Garvey-Kelson mass relationship

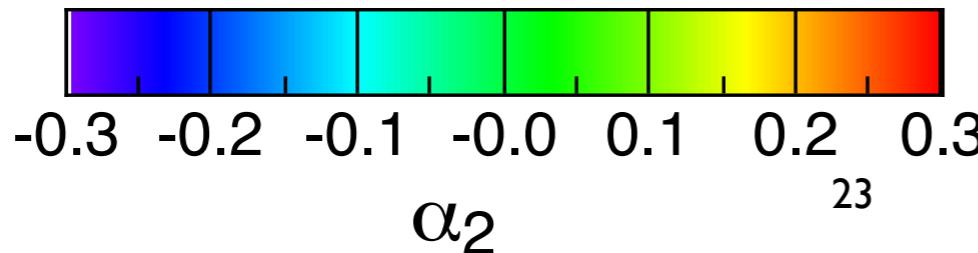
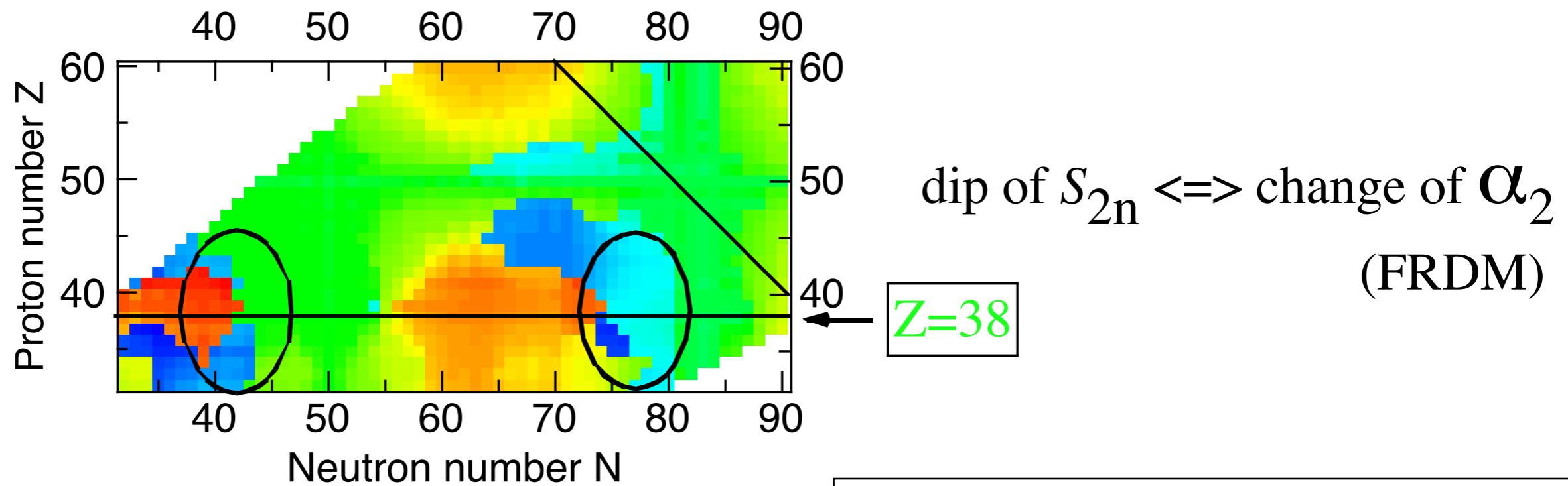


Deformation parameter α_2

KUTY



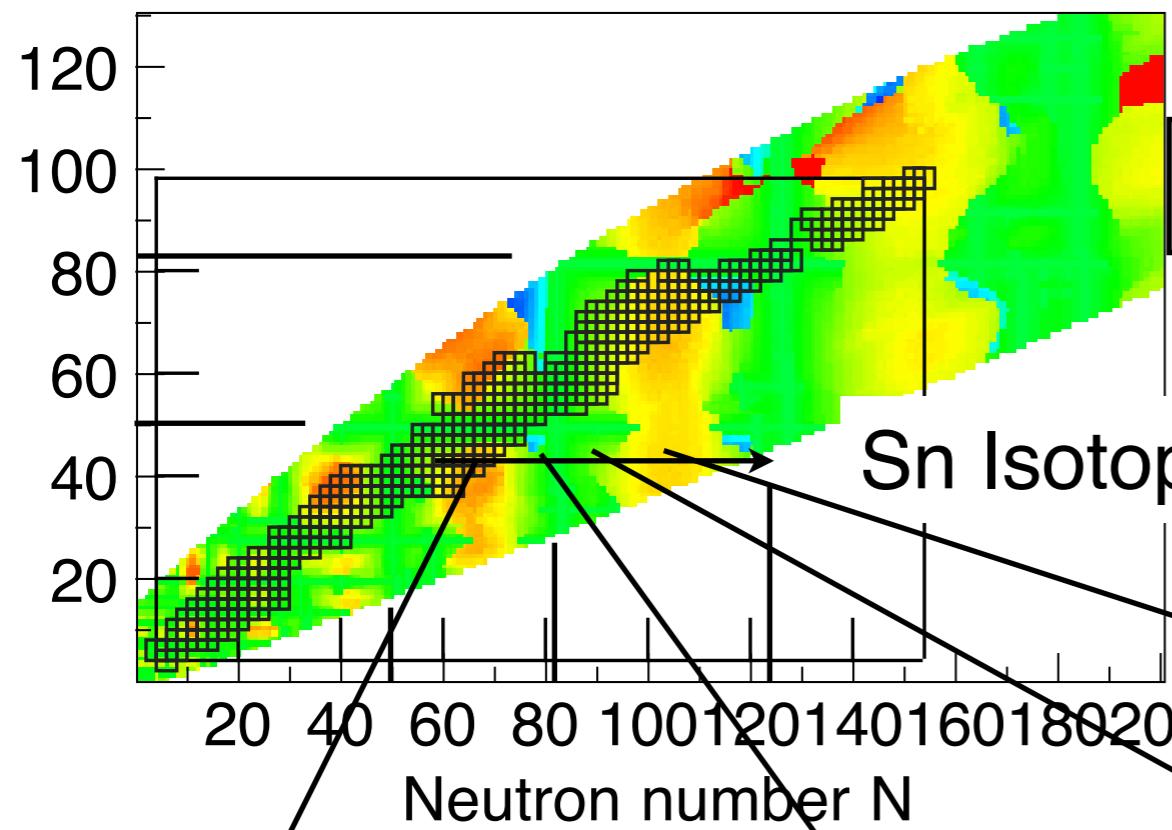
FRDM



Discontinuous change of shape
would give kink of Sn. (FRDM case)
Theoretical (numerical) problem?

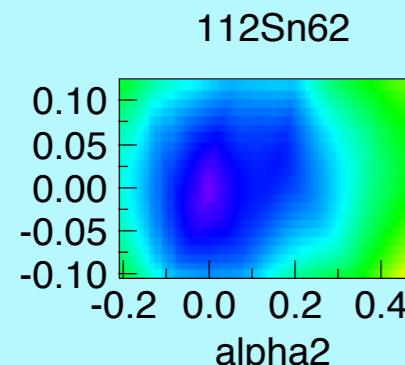
Deformation parameter alpha2 of KUTY

Potential energy surface

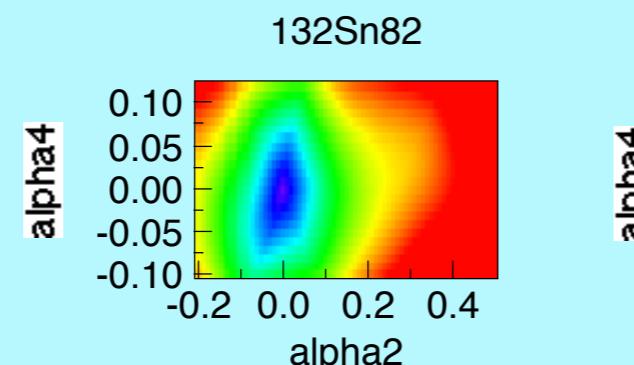


$$E_{sh}(Z, N) = \sum_{\text{def.}} (\langle E_{sh}^{\text{sph}}(Z, N) \rangle_{\text{def.}} + \Delta E_S(Z, N) - \Delta E_C(Z, N) - \Delta E_{\text{pro}}(Z, N))$$

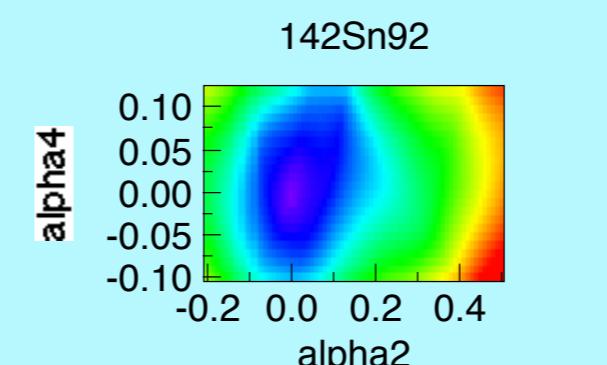
sph. min



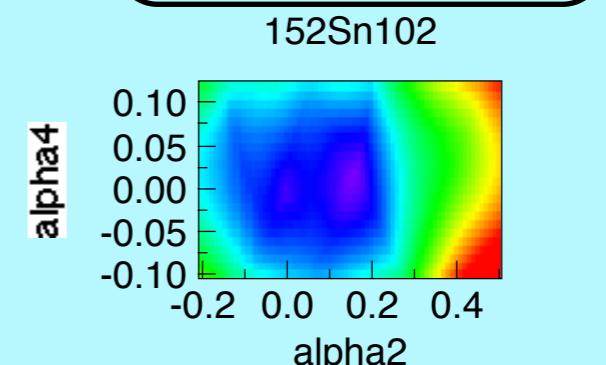
sph. min



sph. min



prolate min



112Sn₆₂

132Sn₈₂

142Sn₉₂

152Sn₁₀₂

Systematical Study of Tin deformation (Shape coexistence)

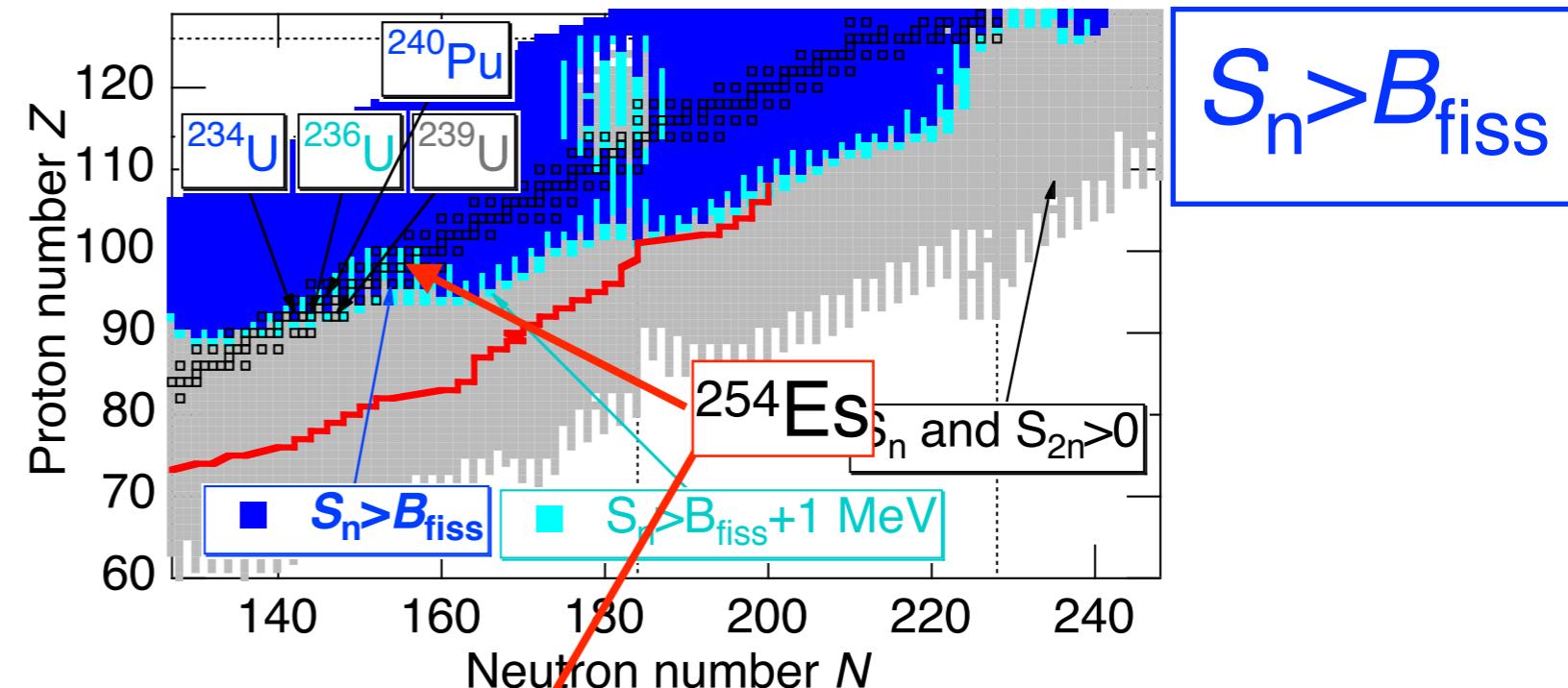
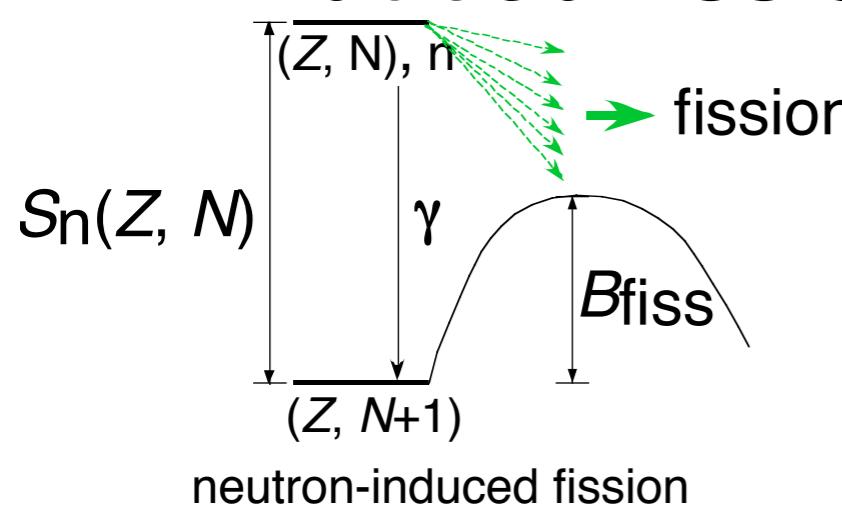
IV. Application to r-process

Effect of fission

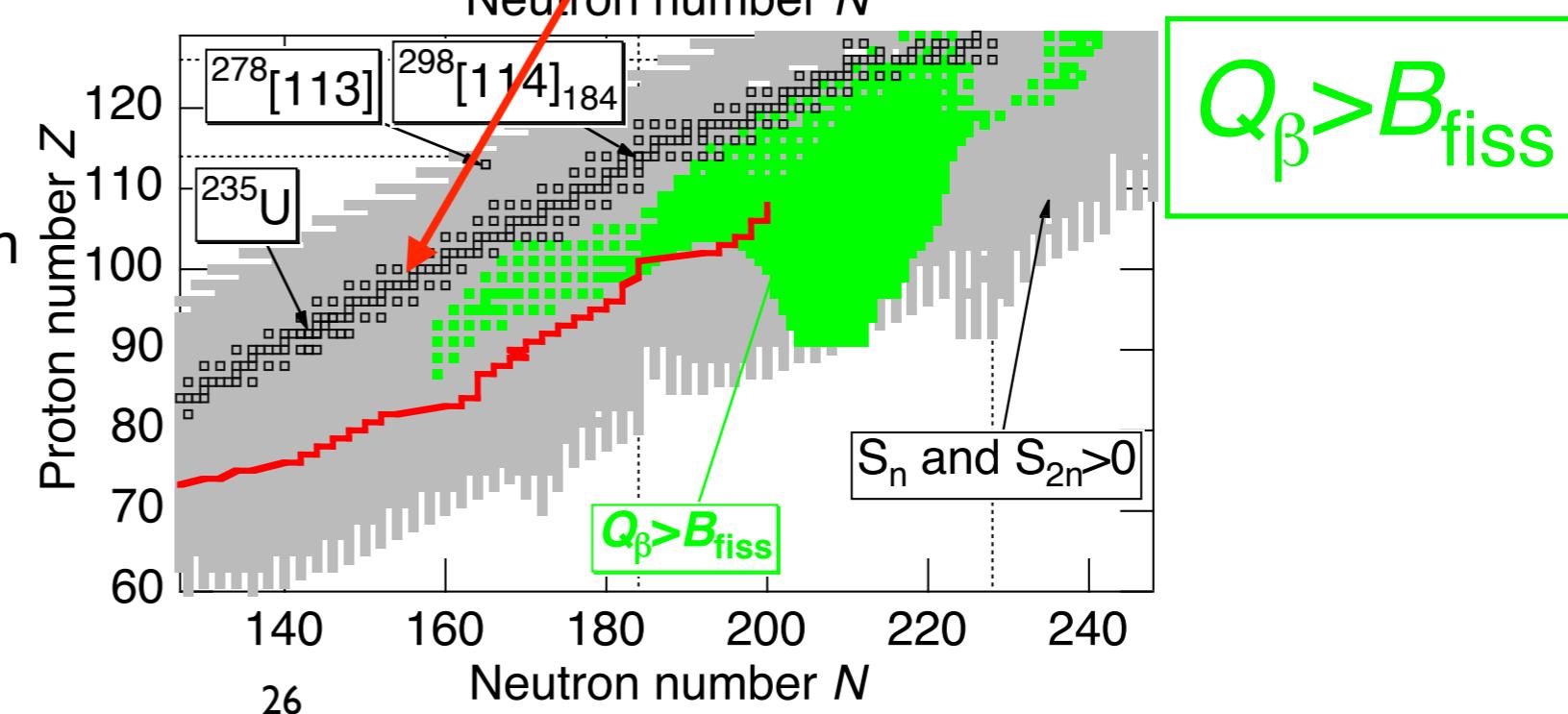
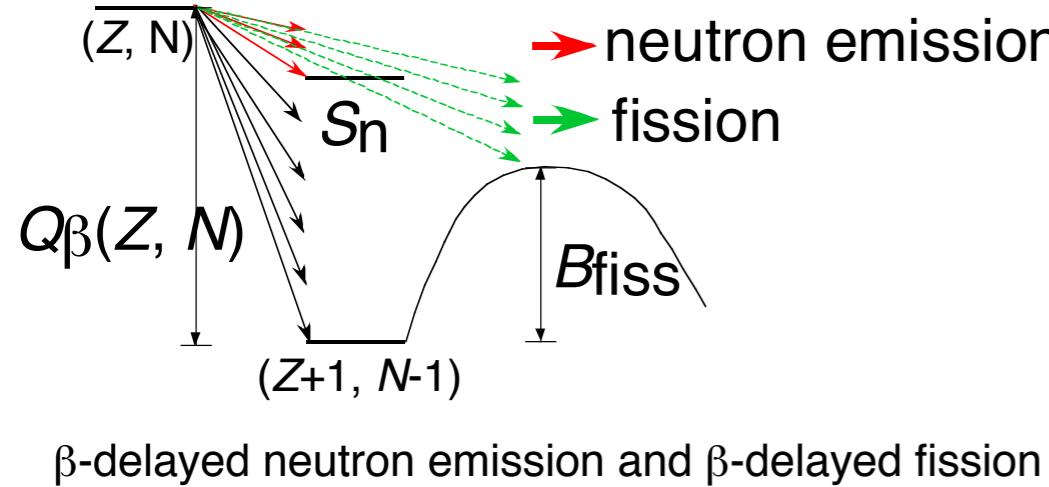
Neutron-induced fission and β -delayed fission

Nuclear masses and fission barrier:
KTUY (Koura-Tachibana-Uno-Yamada) mass formula

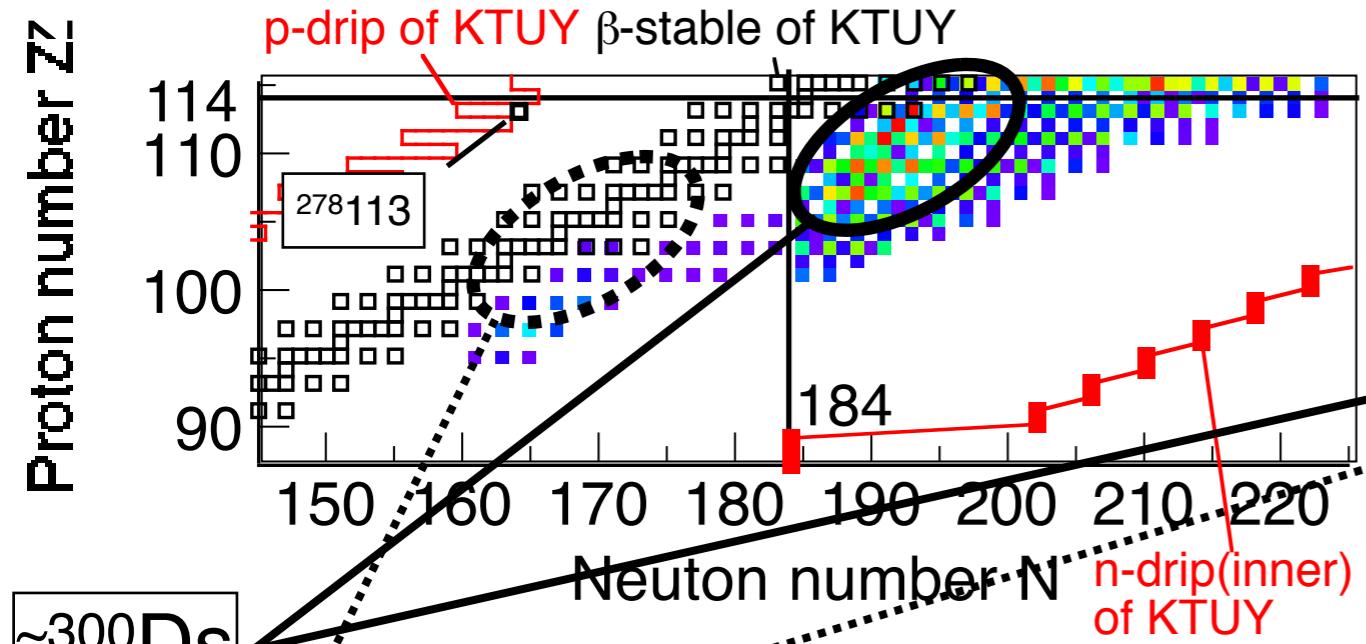
- **n -induced fission**



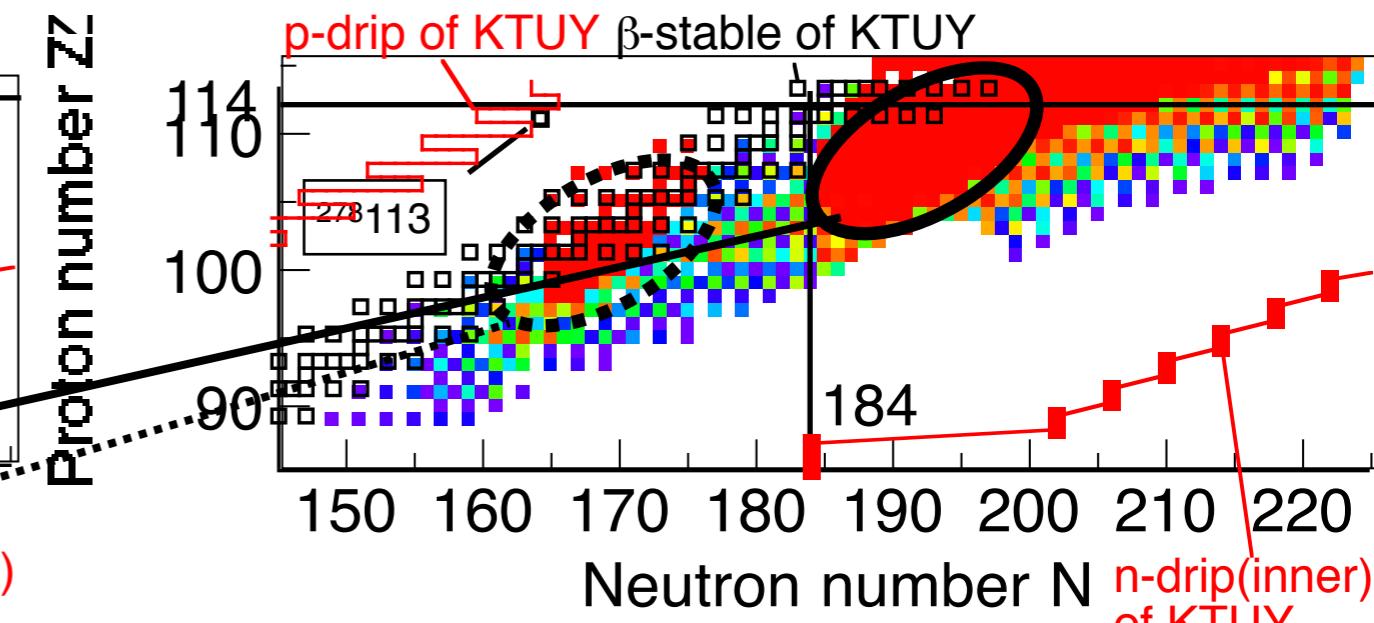
- **β -delayed fission**



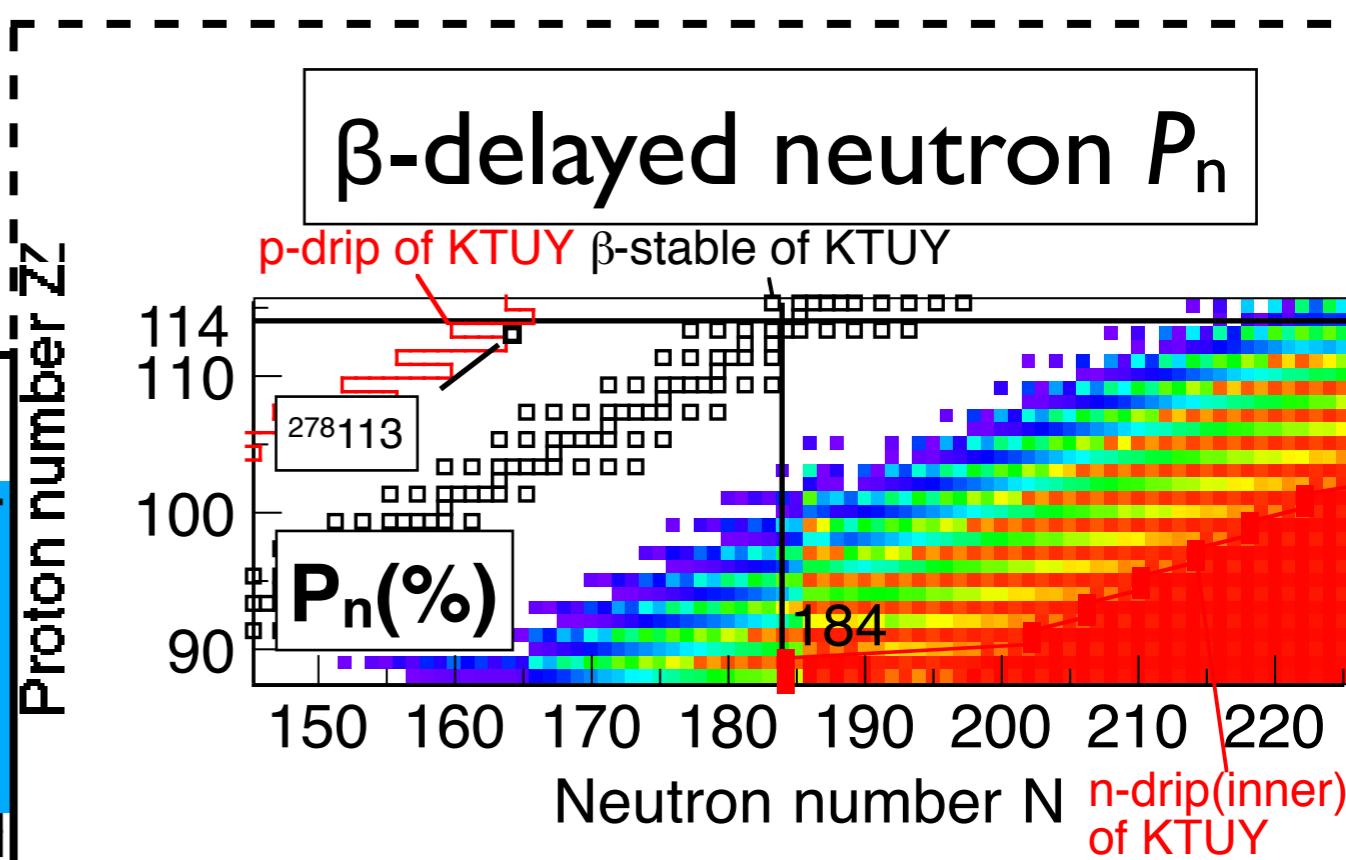
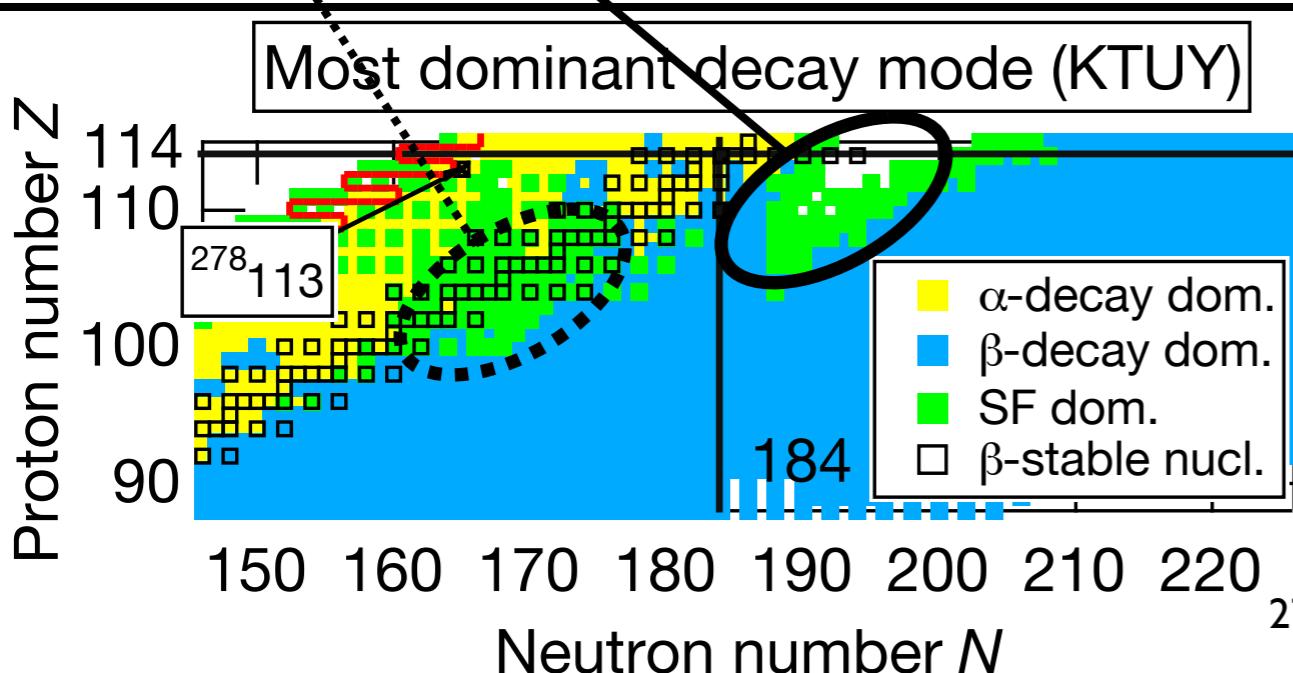
P_f (no correction)



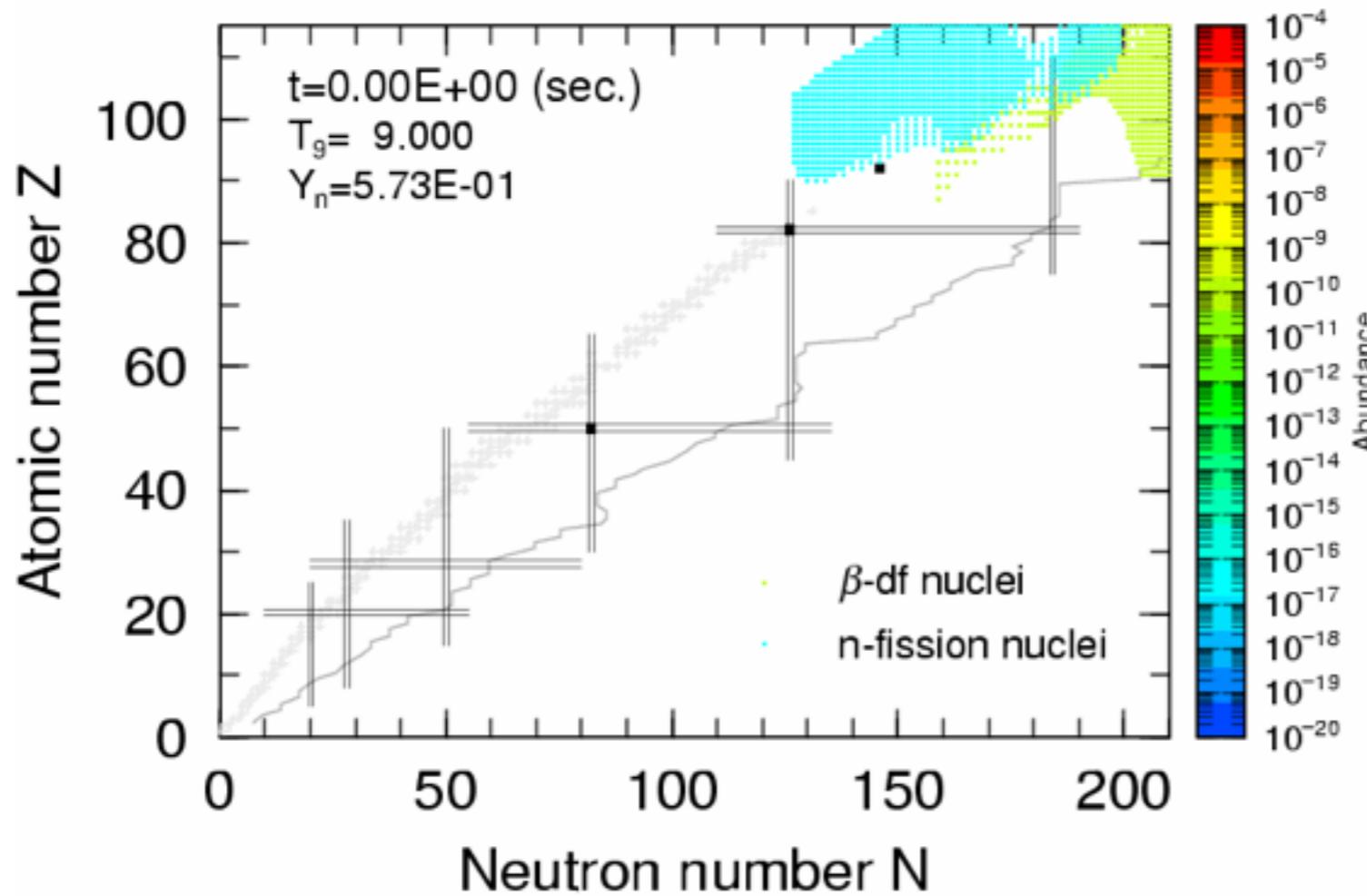
P_f (B_f is 3MeV reduced)



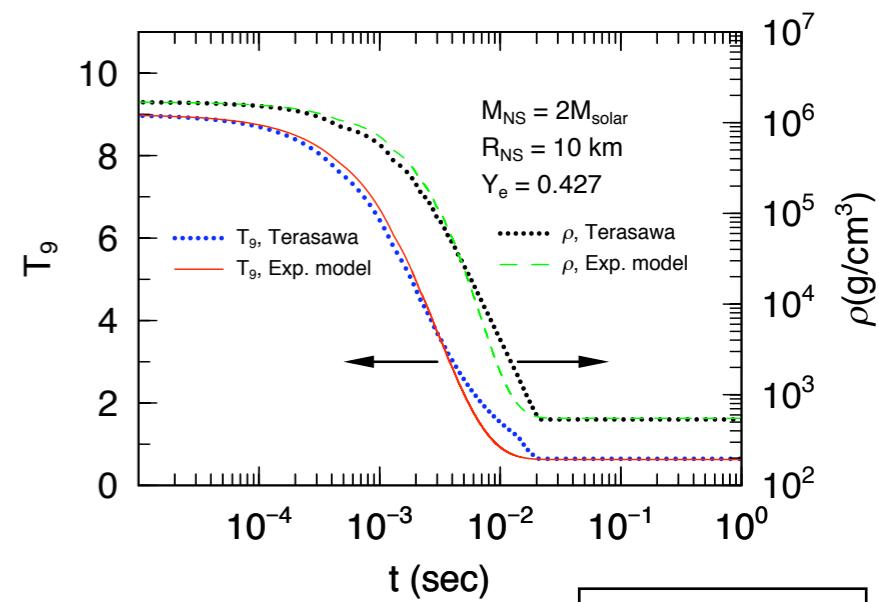
β -delayed neutron P_n



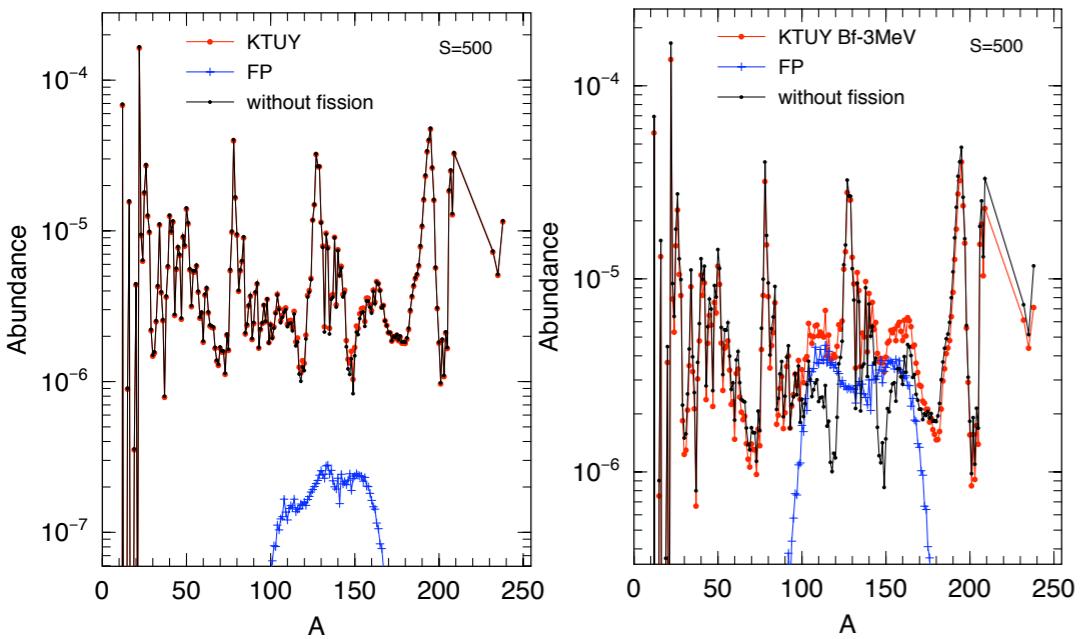
$S=200, B_f \rightarrow B_{fKTUY}-3\text{MeV}$



Trajectory of supernova matter



Sumiyoshi, Kajino



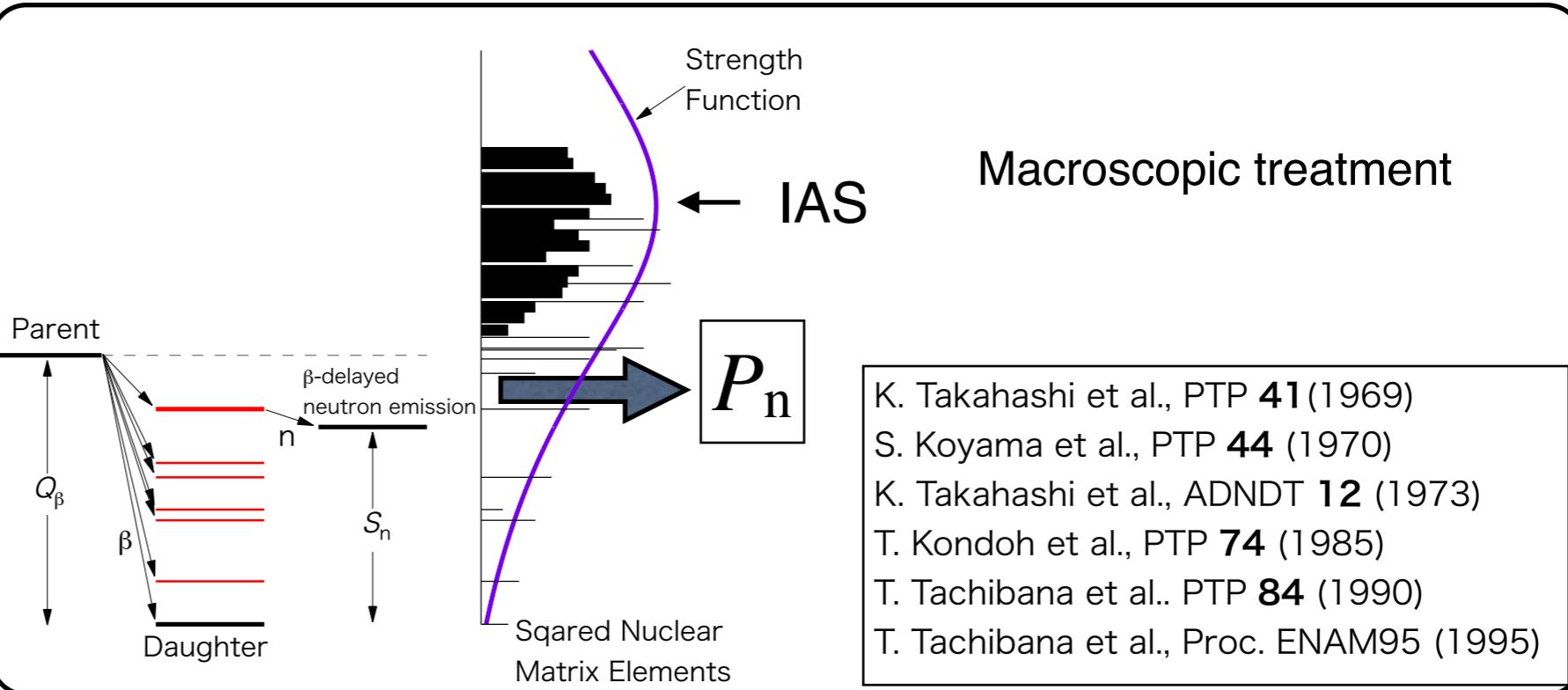
Fission cycle from superheavy region affects the r-process abundance.



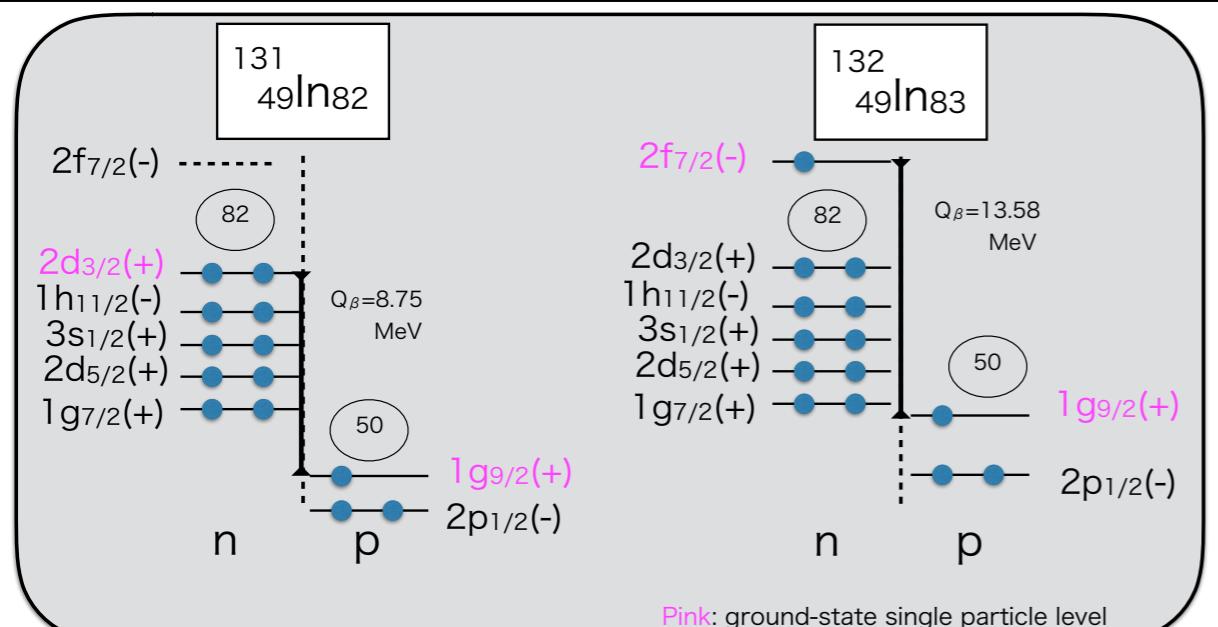
V. Recent study of beta-decay



Gross theory



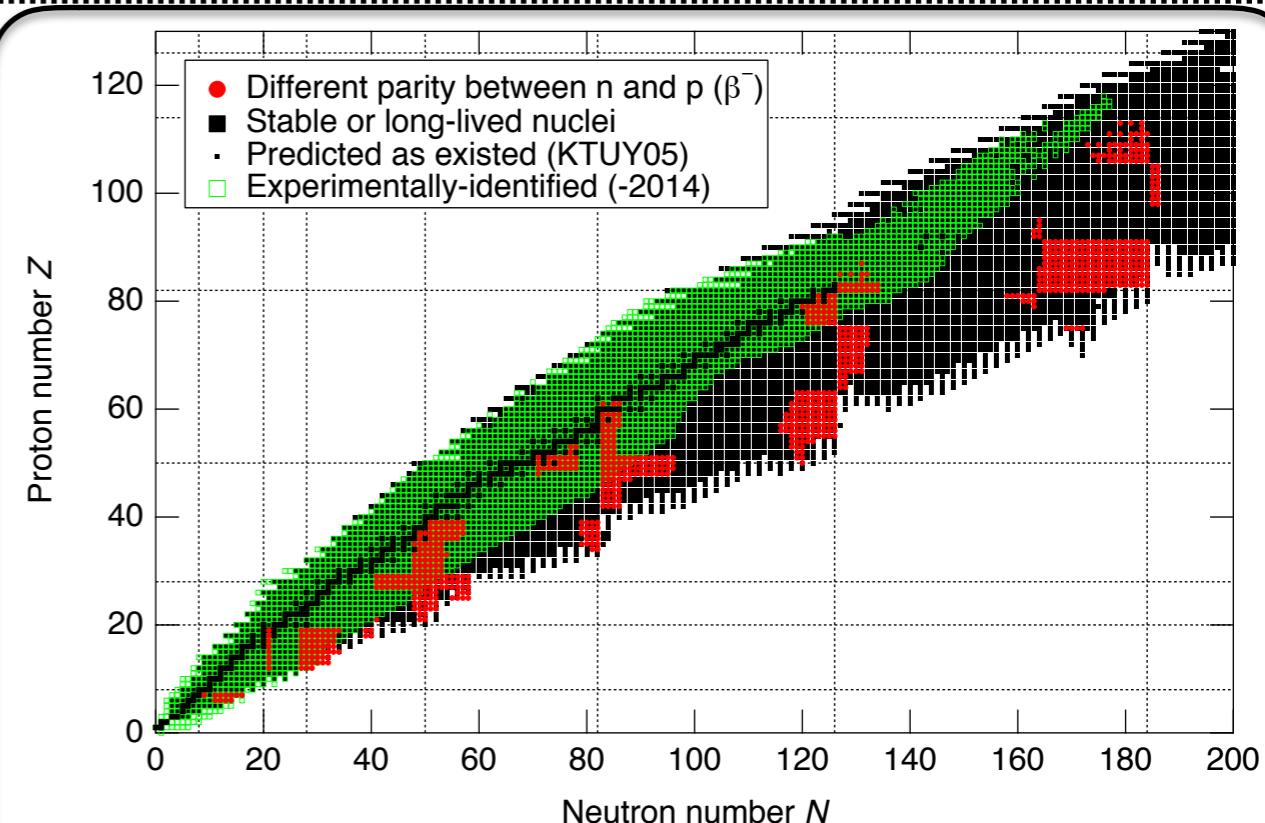
Improvement of the gross theory : Hindrance factor is introduced for allowed transition on change in parity obtained from calculated SPL.



No parity change

Parity change

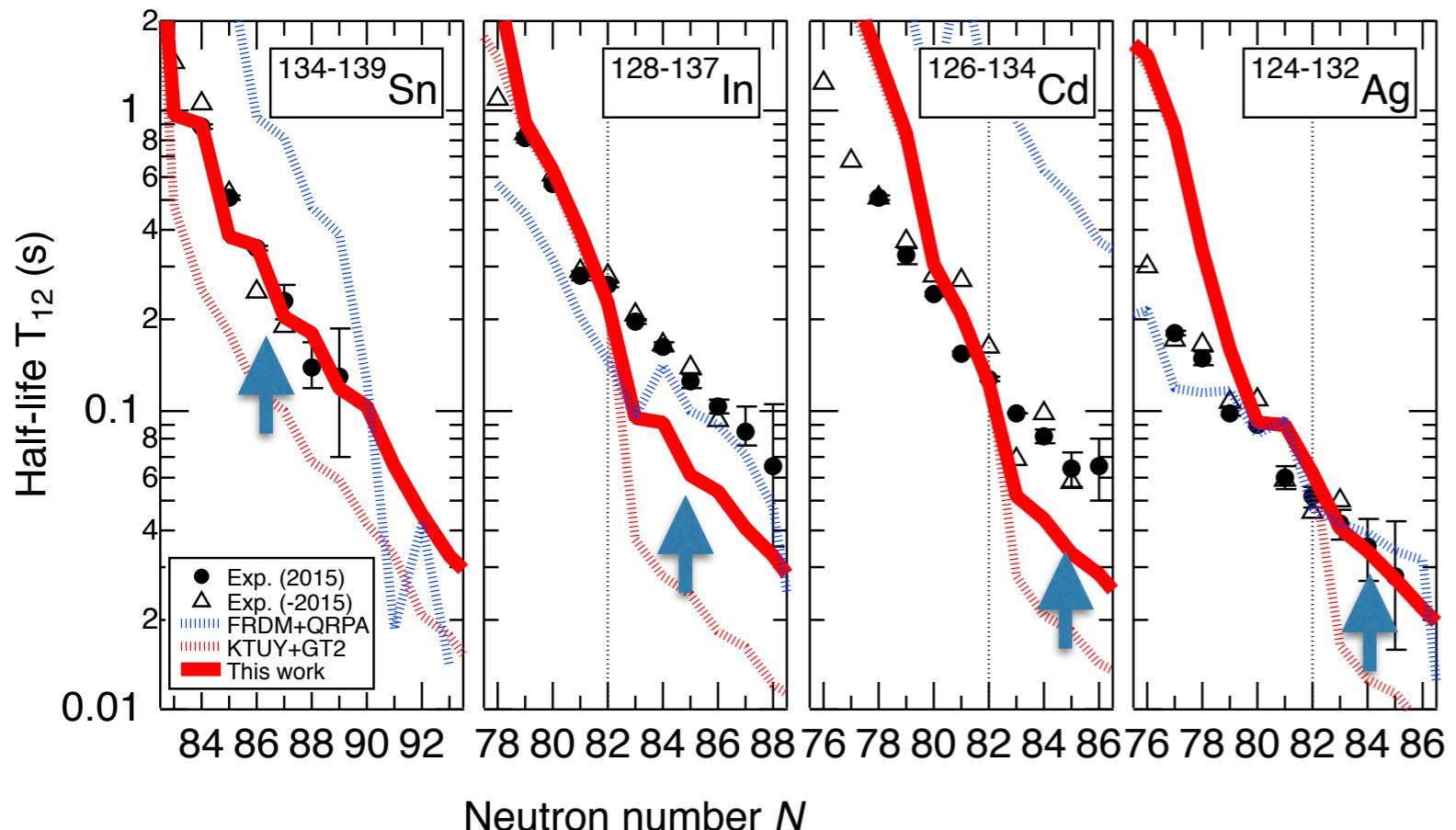
30



Estimated region of parity-mismatching by a modified Woods-Saxon pot.
(H.K. and M. Yamada, NPA **671** 2000)

Recent study: Improvement of the gross theory

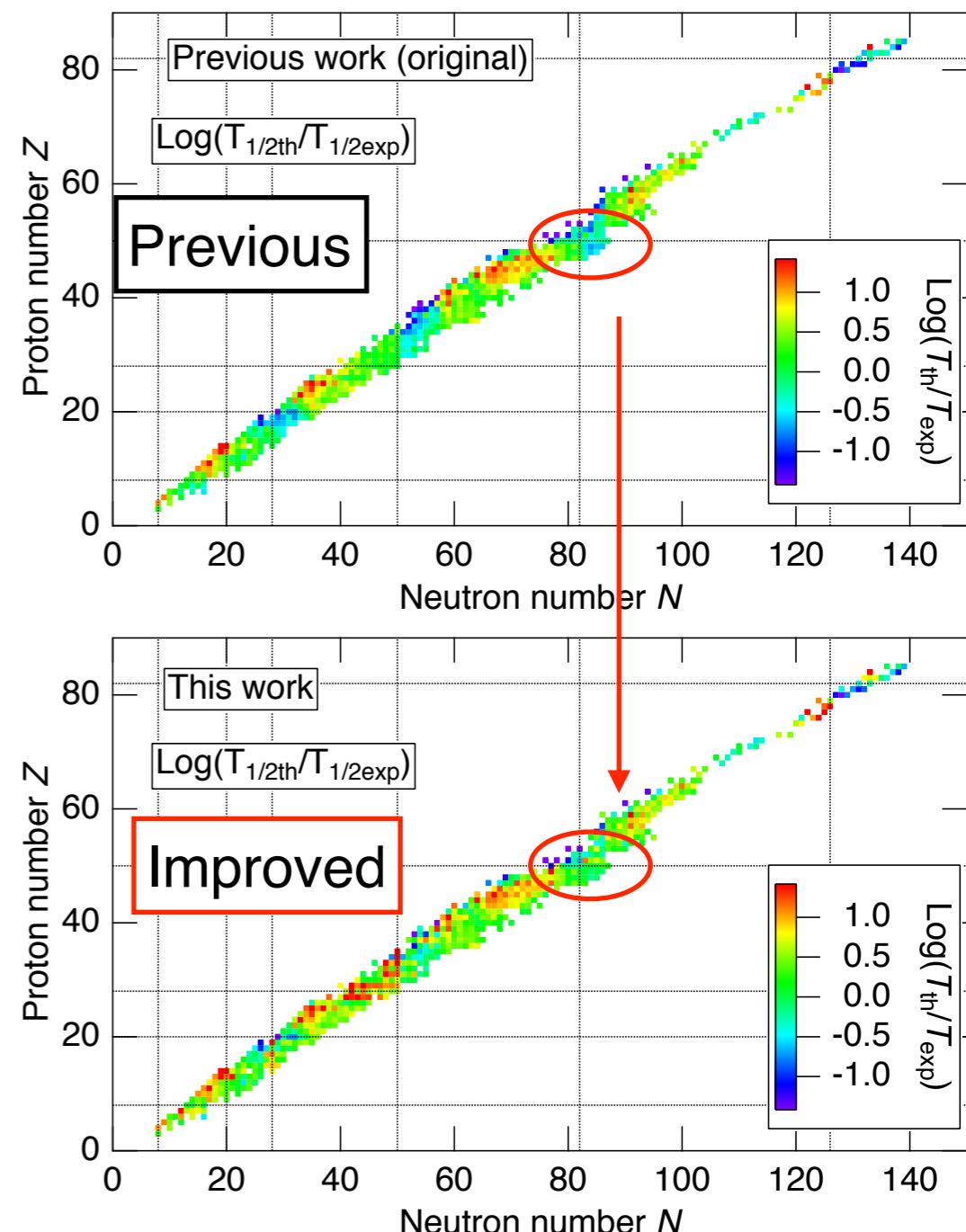
Results



KTUY+improved gross theory

Steep changes at $N=82$ (closed shell) disappear.

H. K. and S. Chiba, PRC 95 064304 (2017)



- A short review of systematical properties of experimental nuclear masses is given.
 - Mass-systematics like G-K is a good tool to check mass values.
- Some mass formulae are reviewed and compared:
 - Old-parametrized mass formulae (in 1976, 1988) generally fail to extrapolation.
 - HFB-type mass formulae give anomaly on GK-sys.
- A mass-model dependency to the r-process is discussed.
- Fission cycle would affect the r-process abundance.