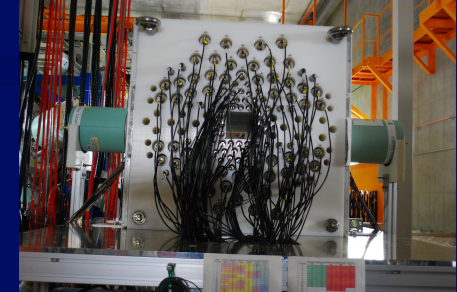


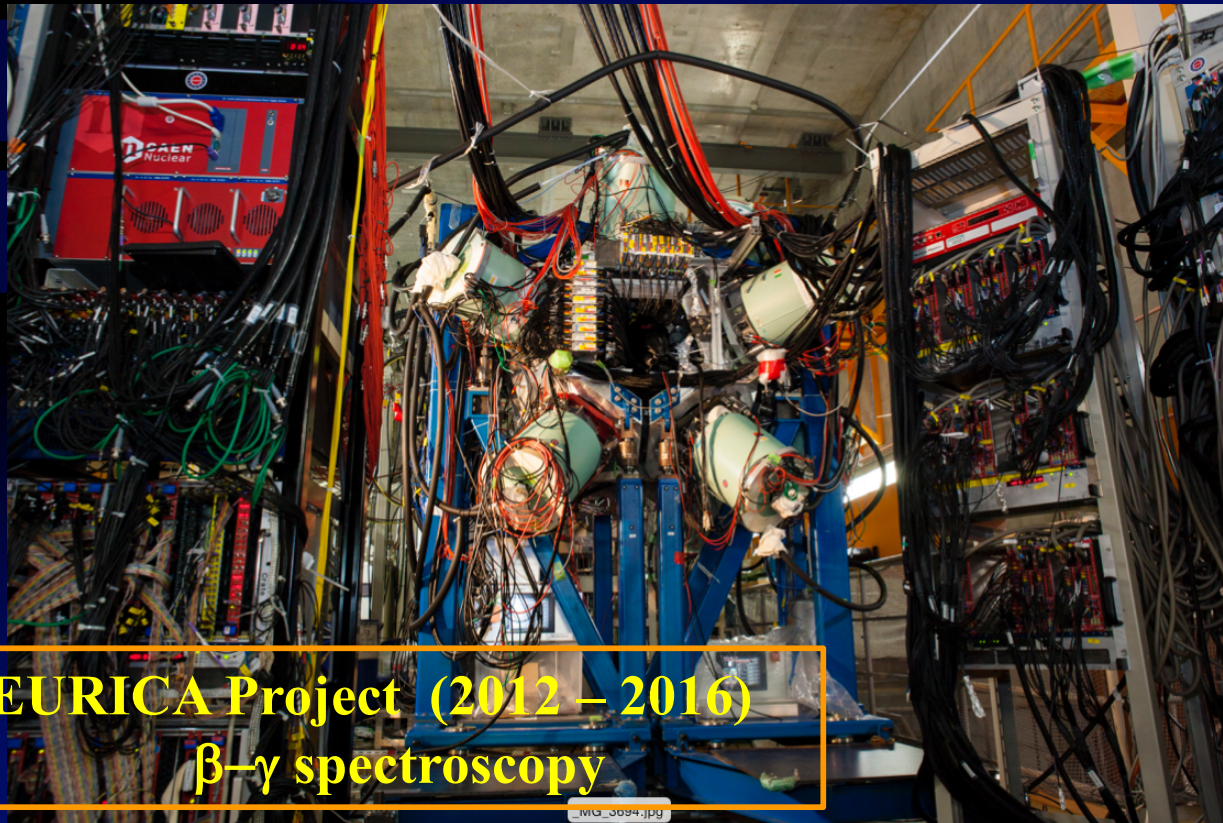
Experimental challenges relevant to r-process nucleosynthesis

Shunji NISHIMURA
RIKEN Nishina Center



- **BRIKEN Project**
(2016 – 2018?)
- β -n- γ spectroscopy

- **Future Projects**
(2019 - ...)
- CAITEN / NiGIRI



EURICA Project (2012 – 2016)
 β - γ spectroscopy

Where is the site of heavy elements ? (r-Process Nucleosynthesis)

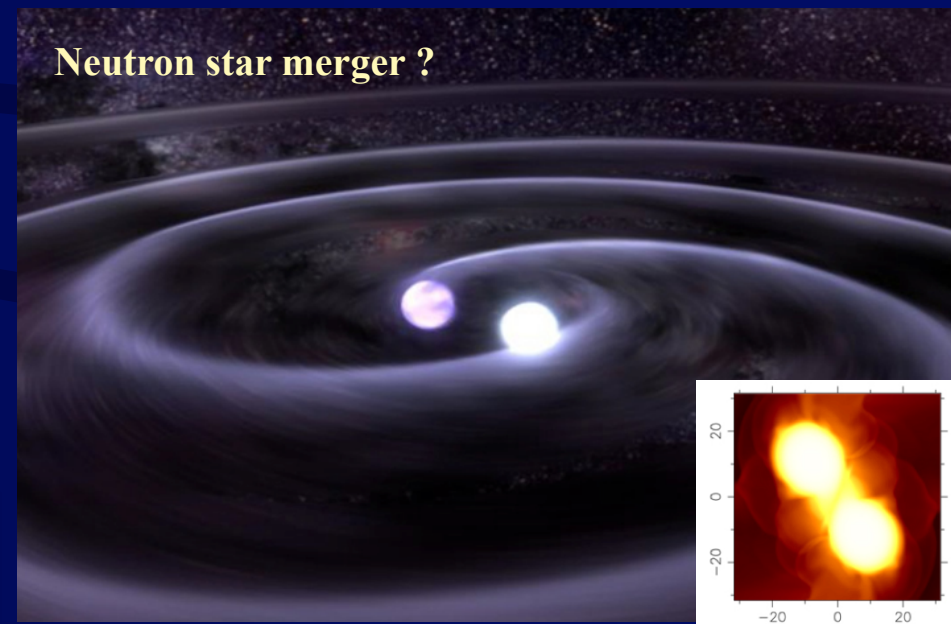


カシオペア座A超新星残骸

NASA/JPL-Caltech/O. Krause
(Steward observatory)

- Mechanism of Explosion.. ?
- Lack of Neutrino, Neutron :
 $Y_e < 0.5$?
- Strong magnetic field ?

N. Nishimura, T. Takiwaki, F.-K. Thielemann
Astrophys. J. 810 109 (2015)



S. Wanajo

- Extremely neutron-rich nuclei
- Very Rare to have two neutron stars
close together.
- Not possible in 1st stars !?

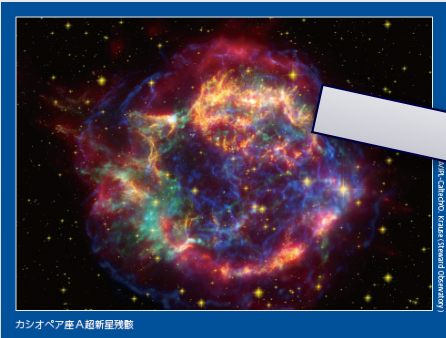
Origin of Heavy Elements Supernovae vs Neutron Star Merger

r-process conditions:
 Y_e , S , τ , EOS, ...

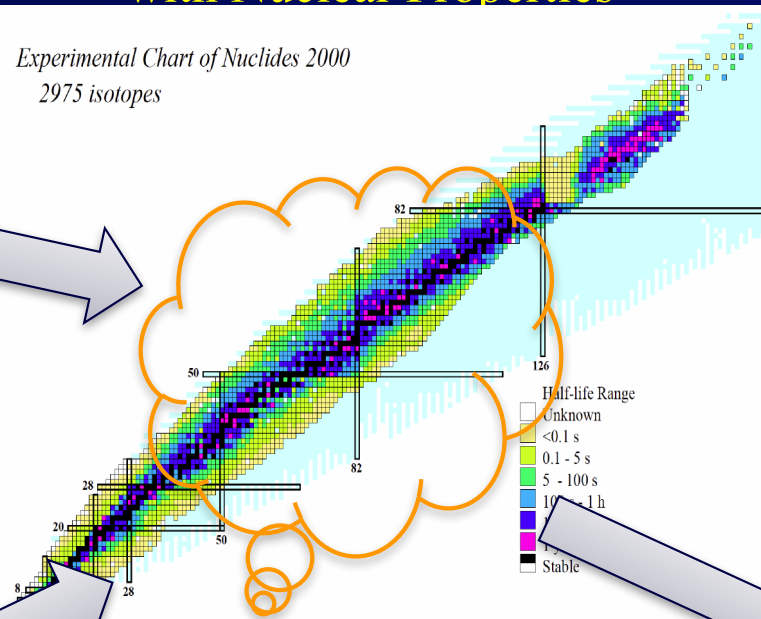
Network calculation
with Nuclear Properties

r-process abundance Z, A
J.J.Cowan C.Snedden, Nature 440 (2006)

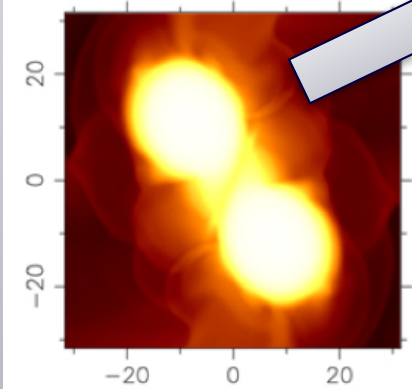
Supernovae explosion?



Experimental Chart of Nuclides 2000
2975 isotopes

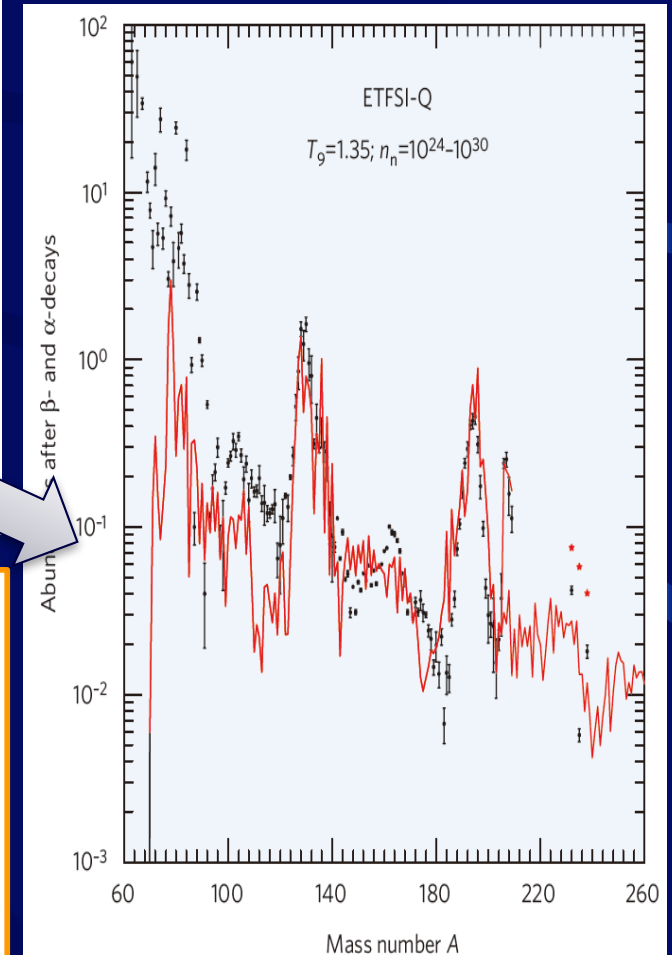


Neutron star merger?



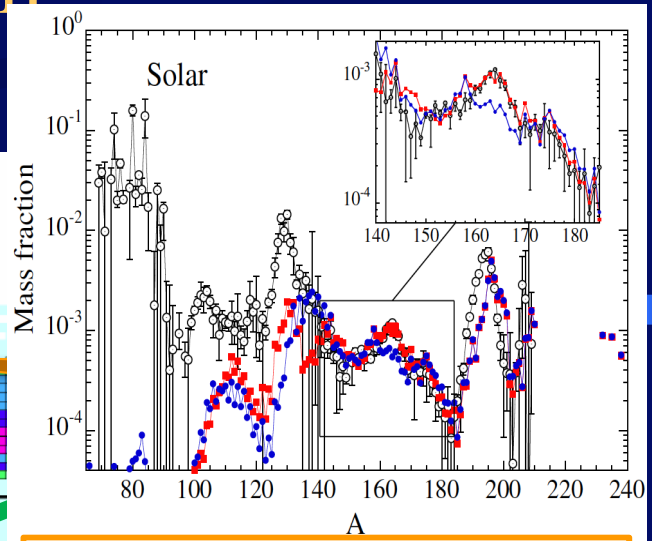
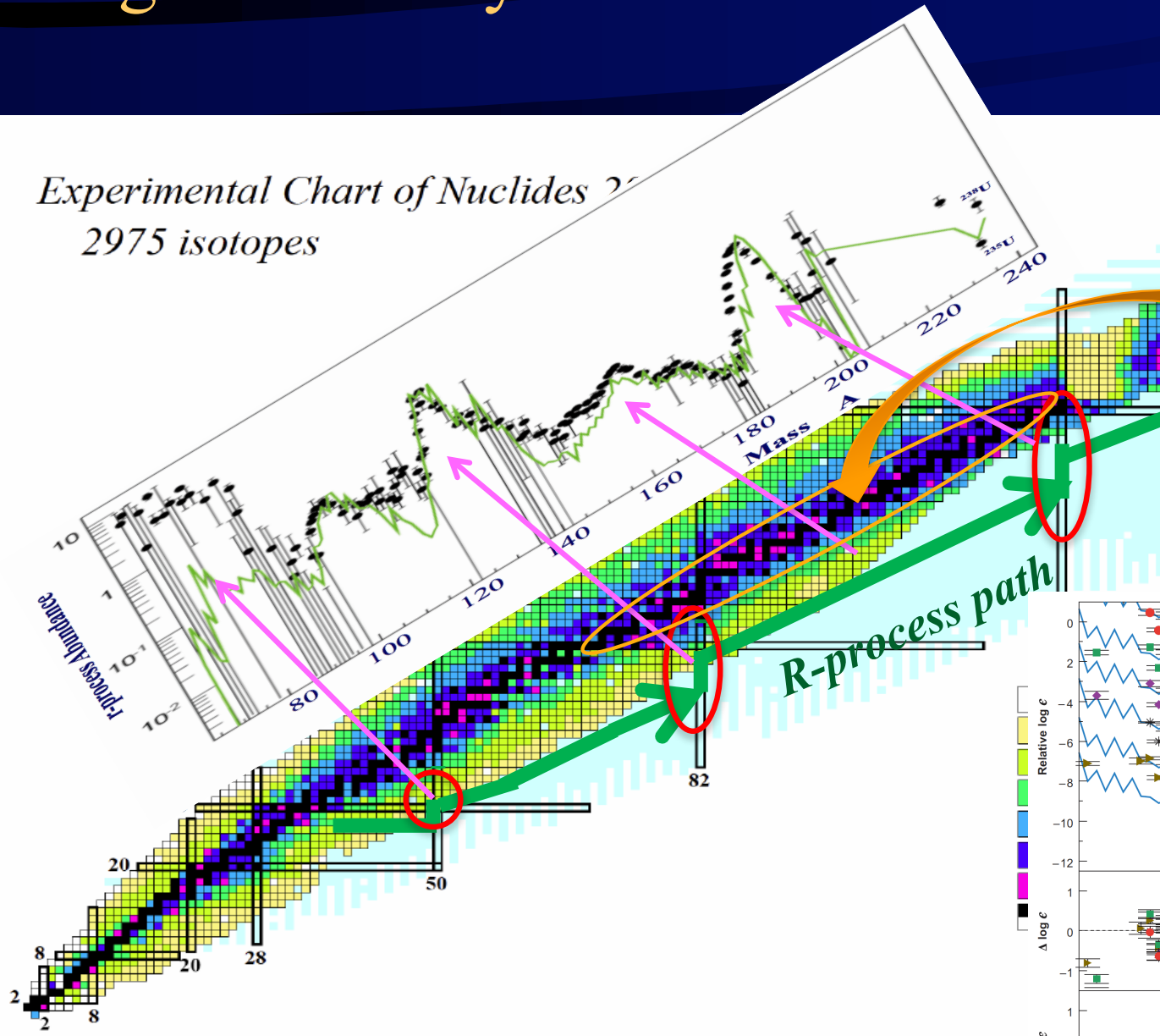
S. Wanajo

Mass ... r-process path $(n, \gamma) \rightleftharpoons (\gamma, n)$
Half-lives ... process speed/abundance
Delayed neutron emission & Fission
... freeze-out path, odd-even
Nuclear Structure
- Neutron Magic ($N=50, 82, 126$),
Deformation, Quenching



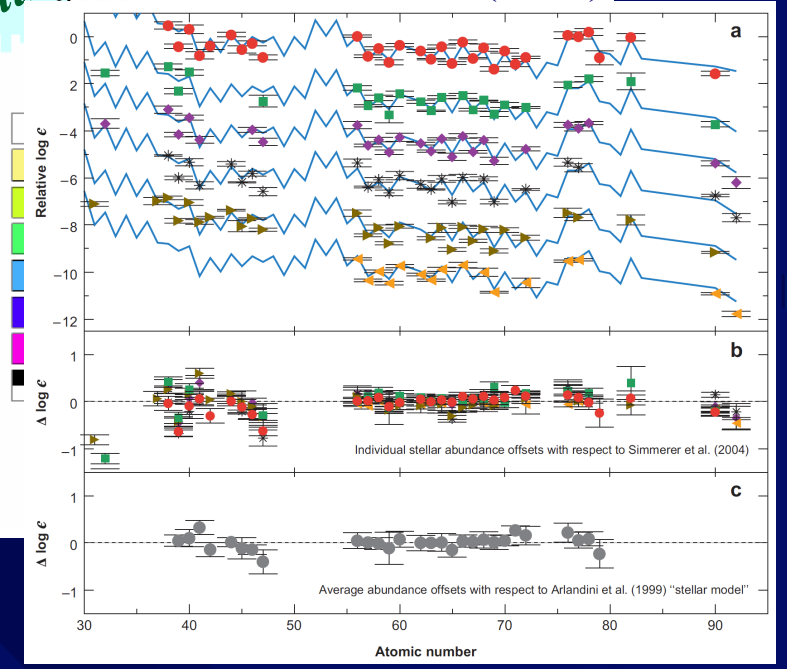
Origin of Heavy Elements: Rare-earth

Experimental Chart of Nuclides 2975 isotopes



S. Goriely et al PRL 111, (2013)

C. Sneden et al. (2008)

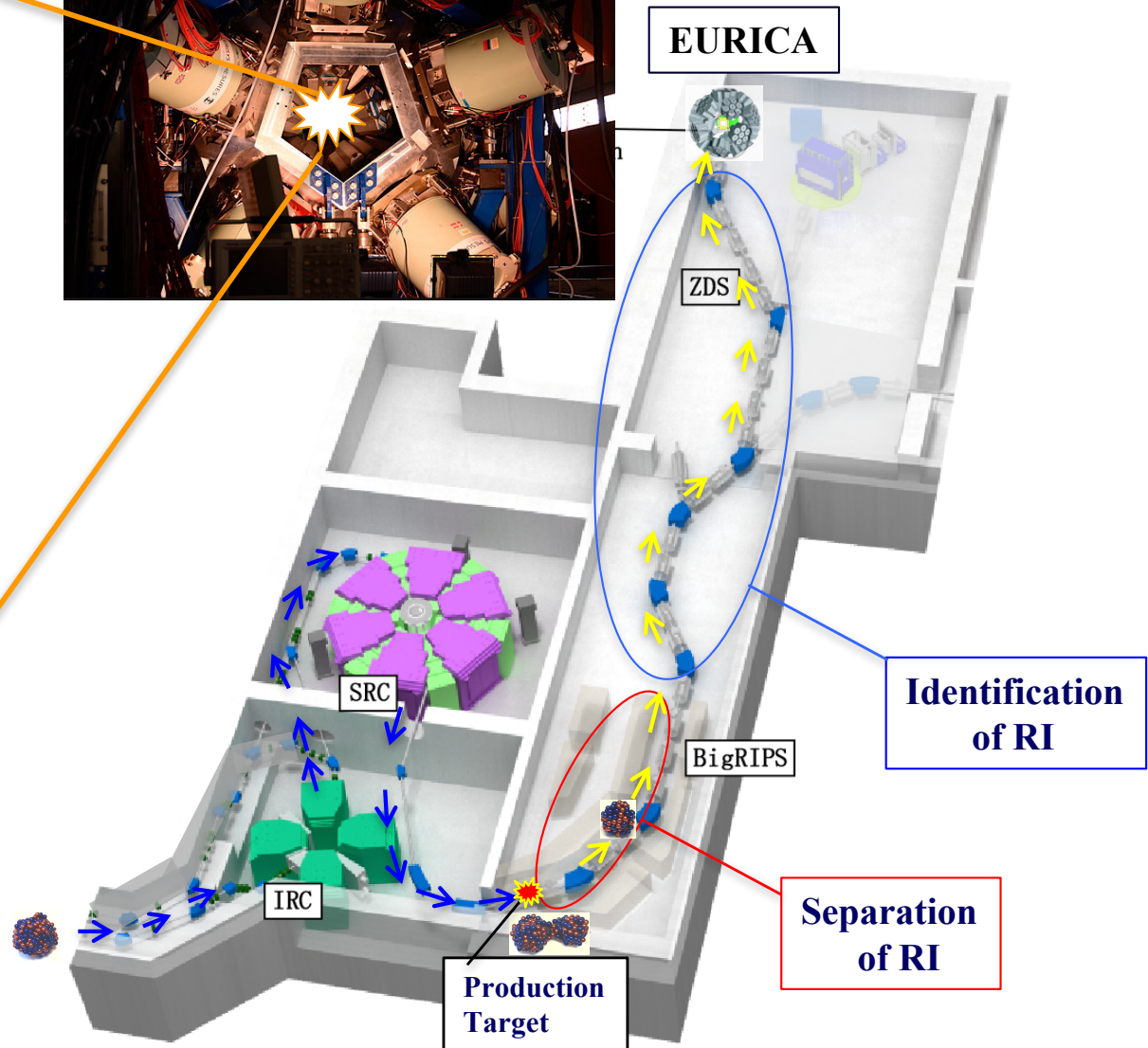
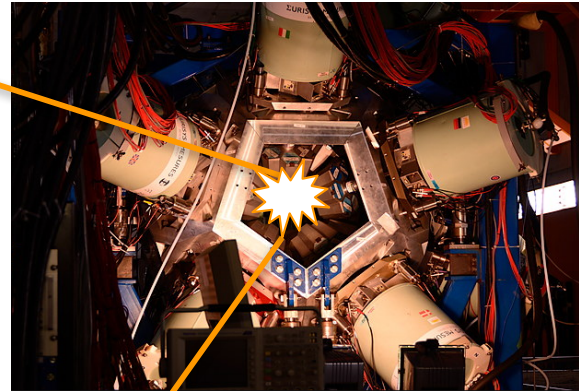
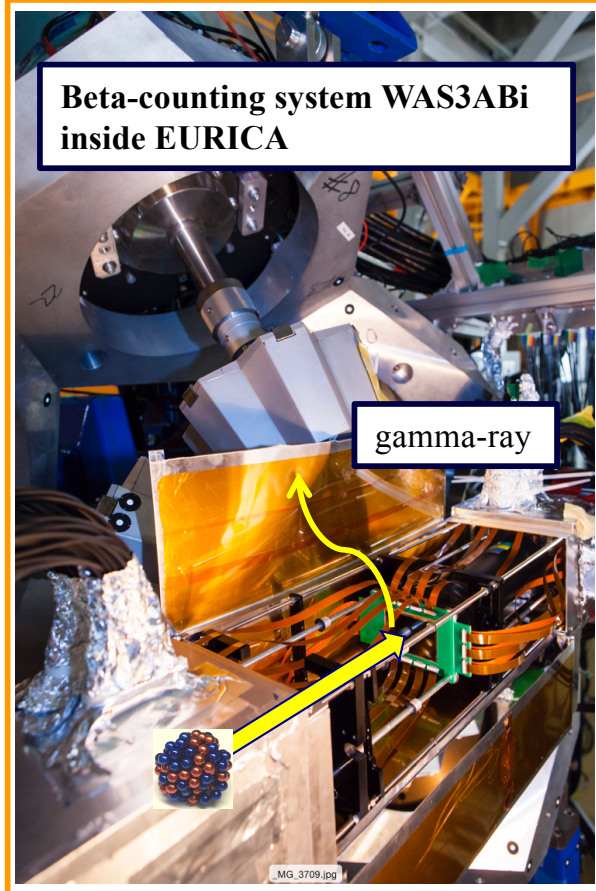



decay rates, masses, fission, b-delayed neutron
 ← → magic, deformation

Decay Spectroscopy at RIBF

EURICA Spectrometer at RIBF

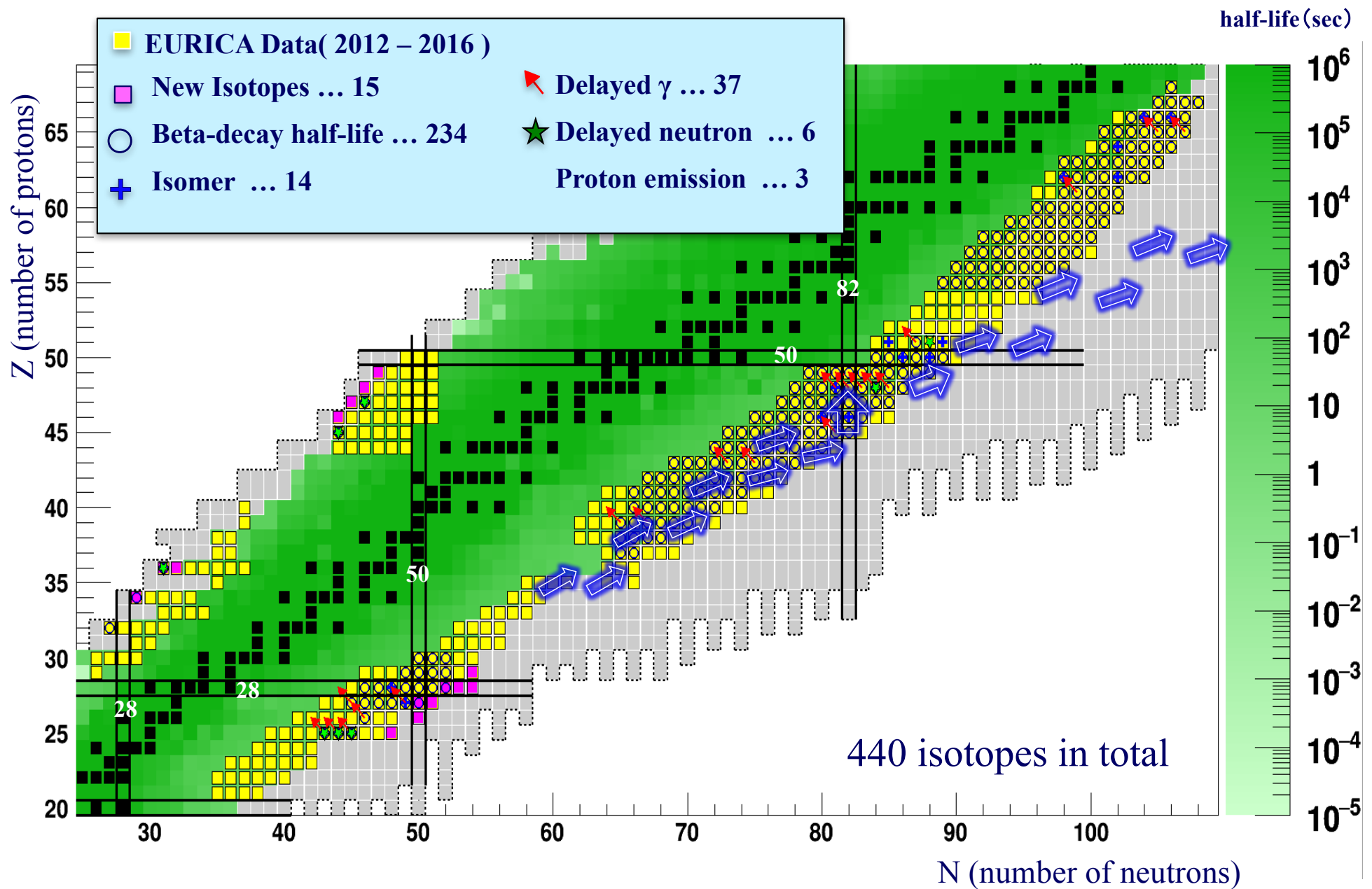
^{238}U ... 345 MeV/u; int. 5 – 12 pA (now 50 pA)



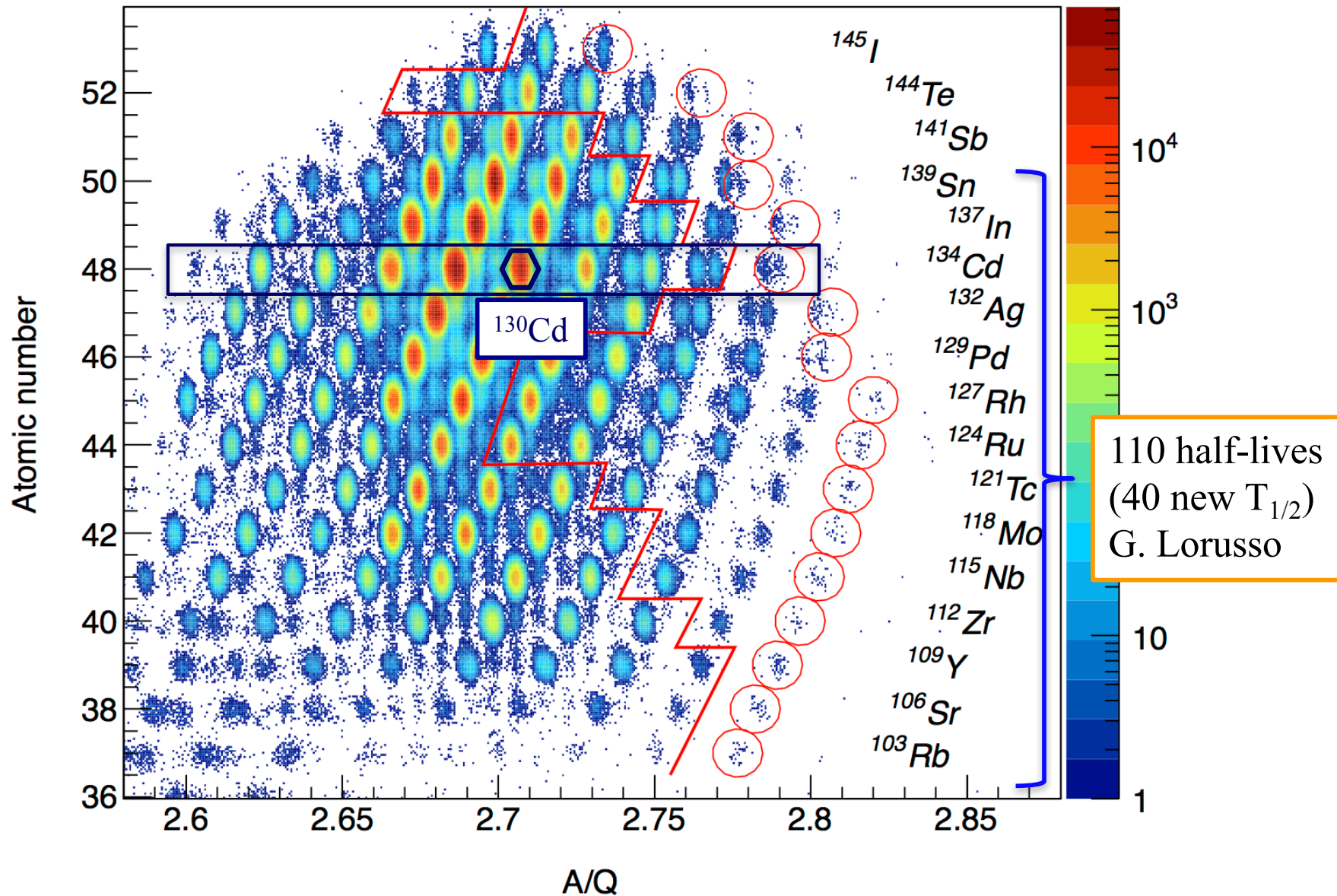
 → Primary Beam
(^{238}U / ^{124}Xe / ^{78}Kr)

 → RI Beam

440 Exotic Isotopes Surveyed by EURICA Spectrometer

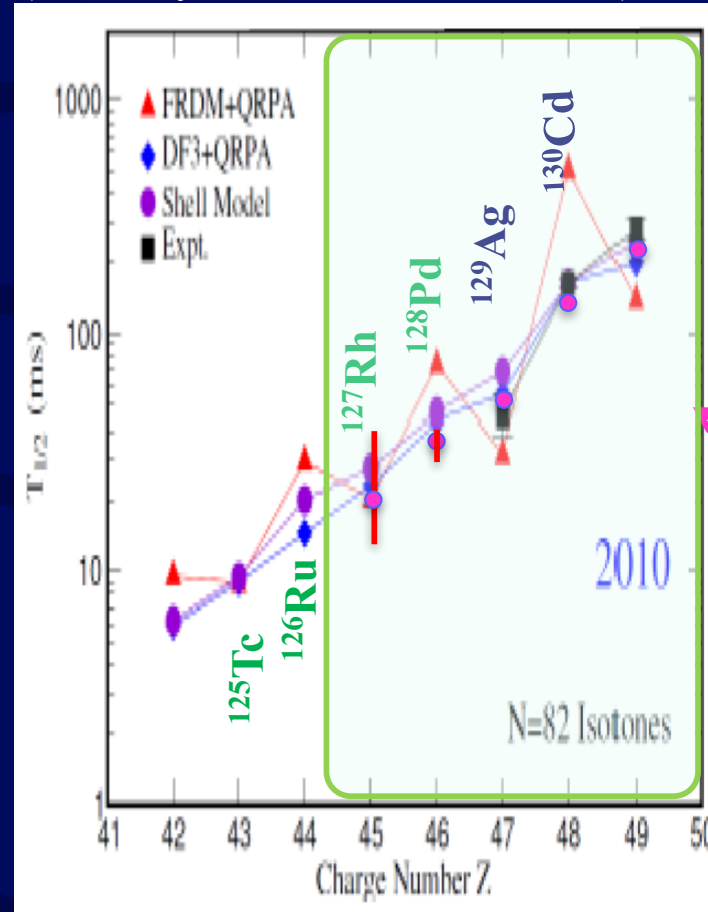
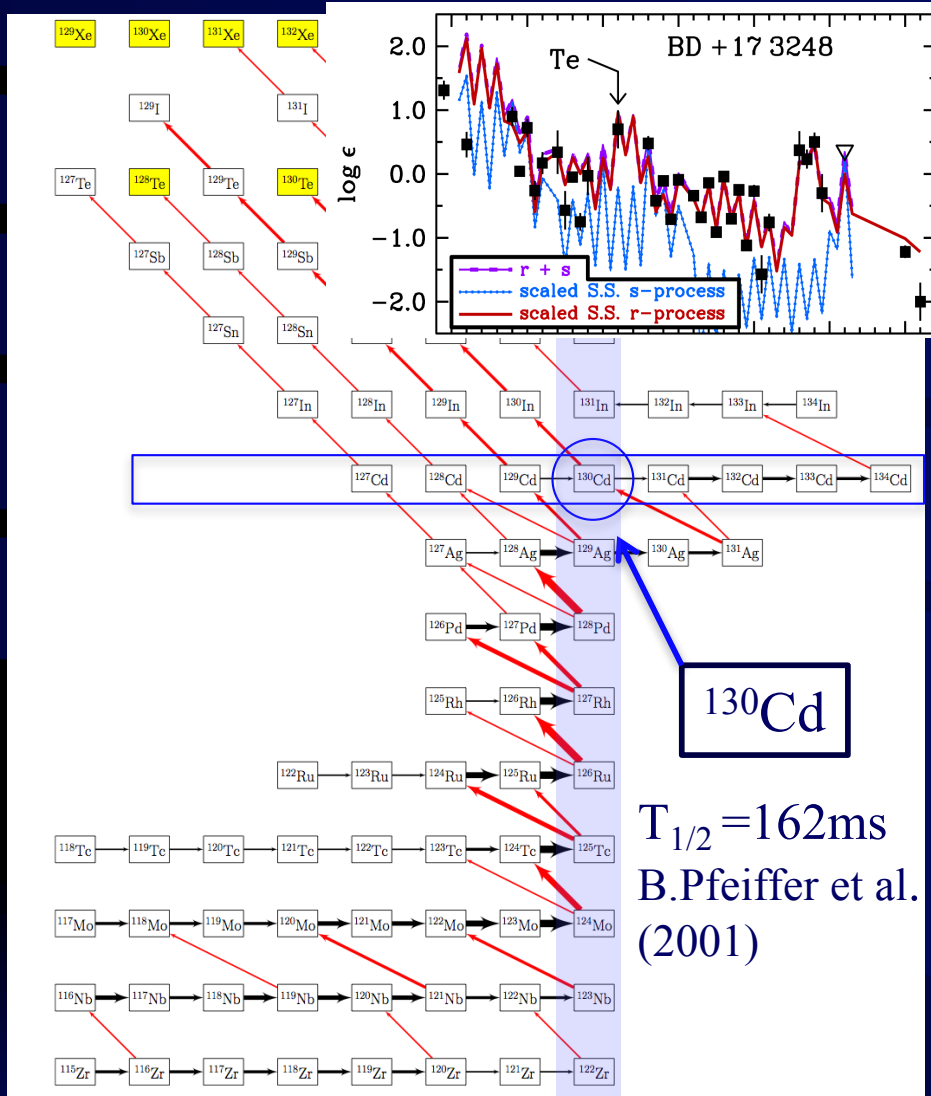


β -decay half-lives on r-process path



Beta-decay Half-lives $N = 82 \rightarrow$ Feedback to the Theory

K.Langanke Phys. Scr. T152 (2013) 014011
(Courtesy of G. Martinez-Pinedo)



So call r-process waiting point nuclei ($N=82$)

- r-process path
- residual r-matter flow in freeze-out

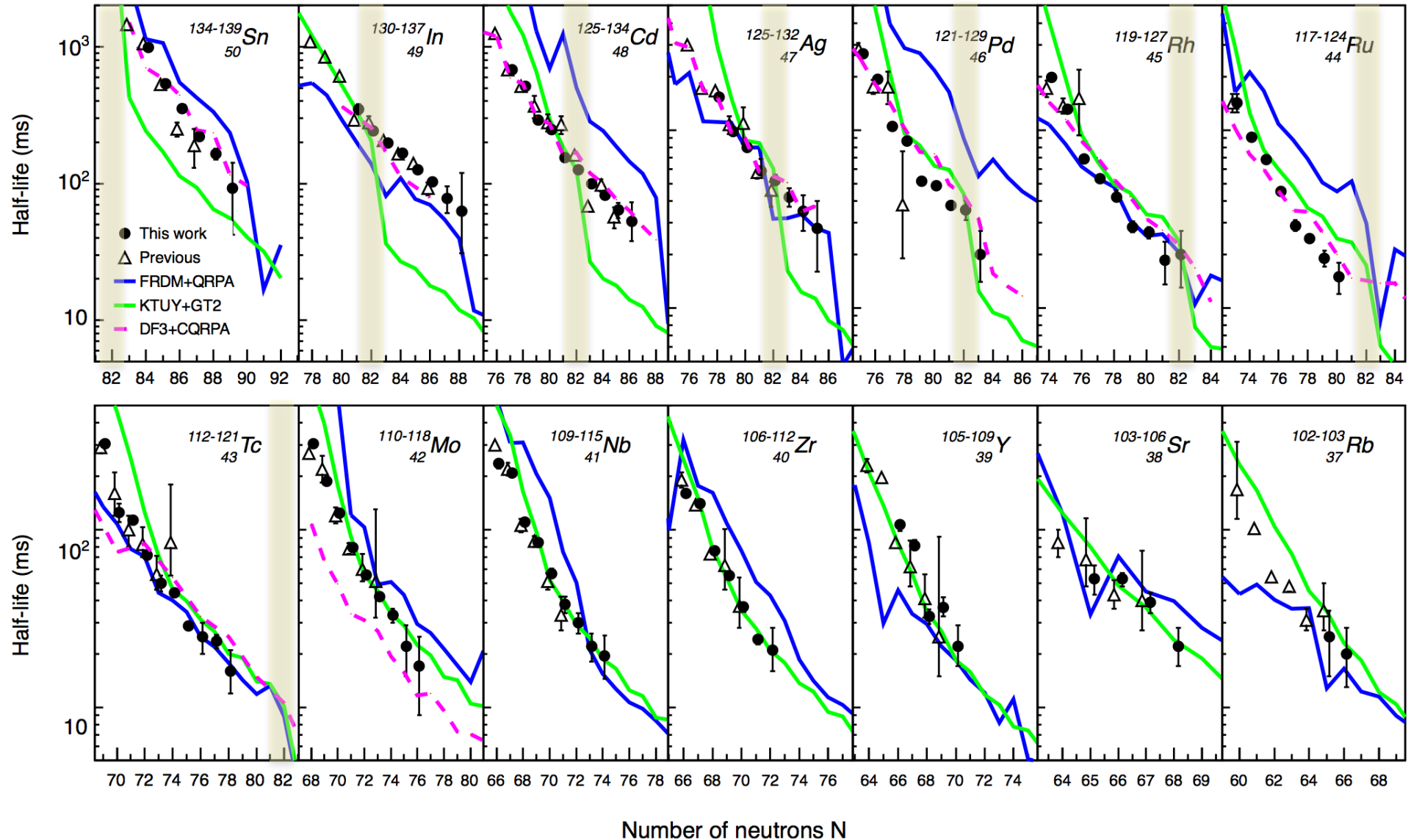
Shell model $\leftarrow \sim 25\%$ shorter $T_{1/2}$

Confirmed by R. Dunlop (2016, TRIUMF)
 $^{130}\text{Cd} \dots T_{1/2} = 126(4)\text{ms}$

110 Half-lives of Very Neutron-Rich Rb to Sn

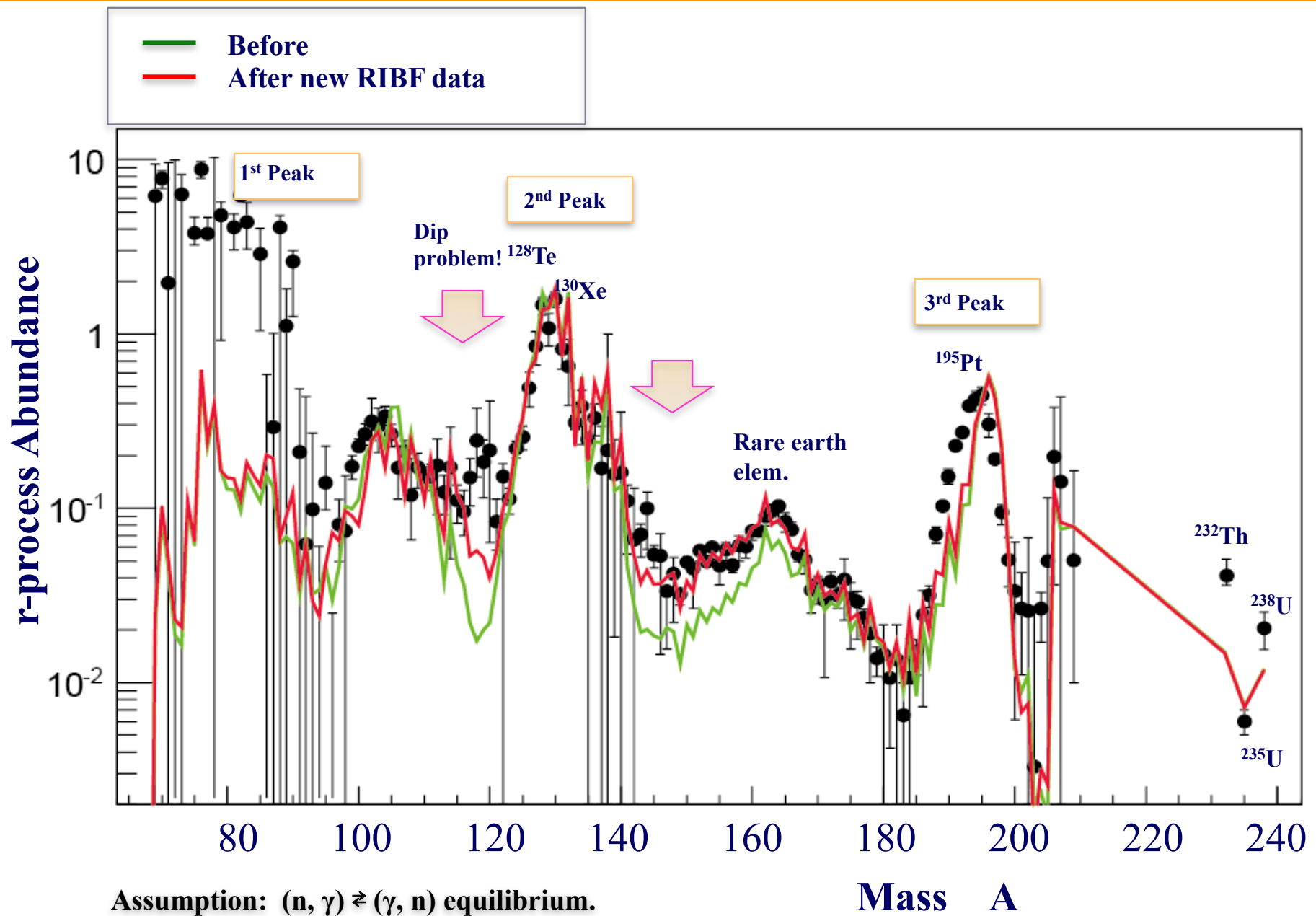
40 new half-lives ! G.Lorusso et al.,
PRL 114, 192501 (2015)

SN
18 new half-lives ! PRL 106, 052502 (2011)

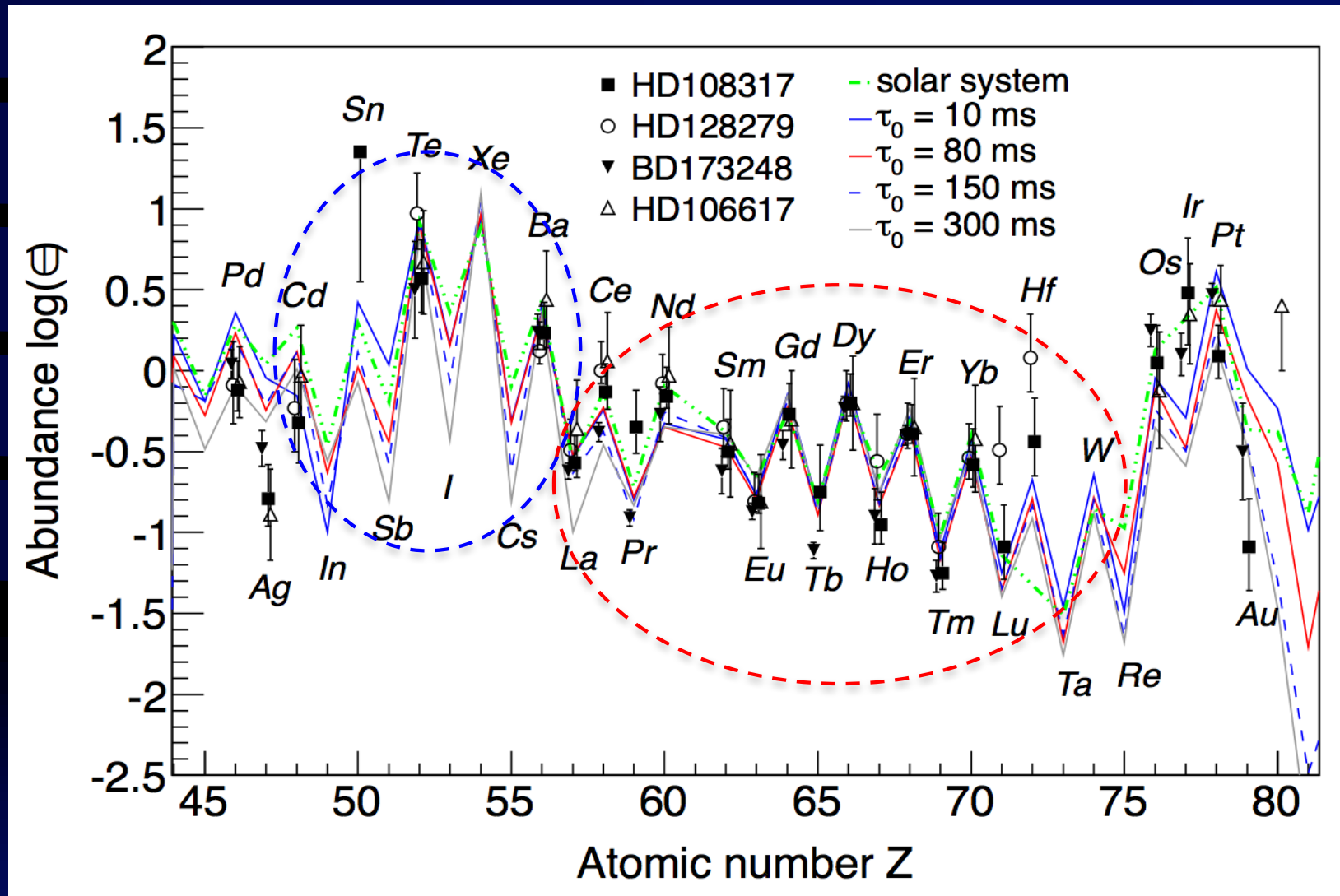


r-process Abundance with New $T_{1/2}$ (RIBF)

G.Lorusso et al., PRL (2015)



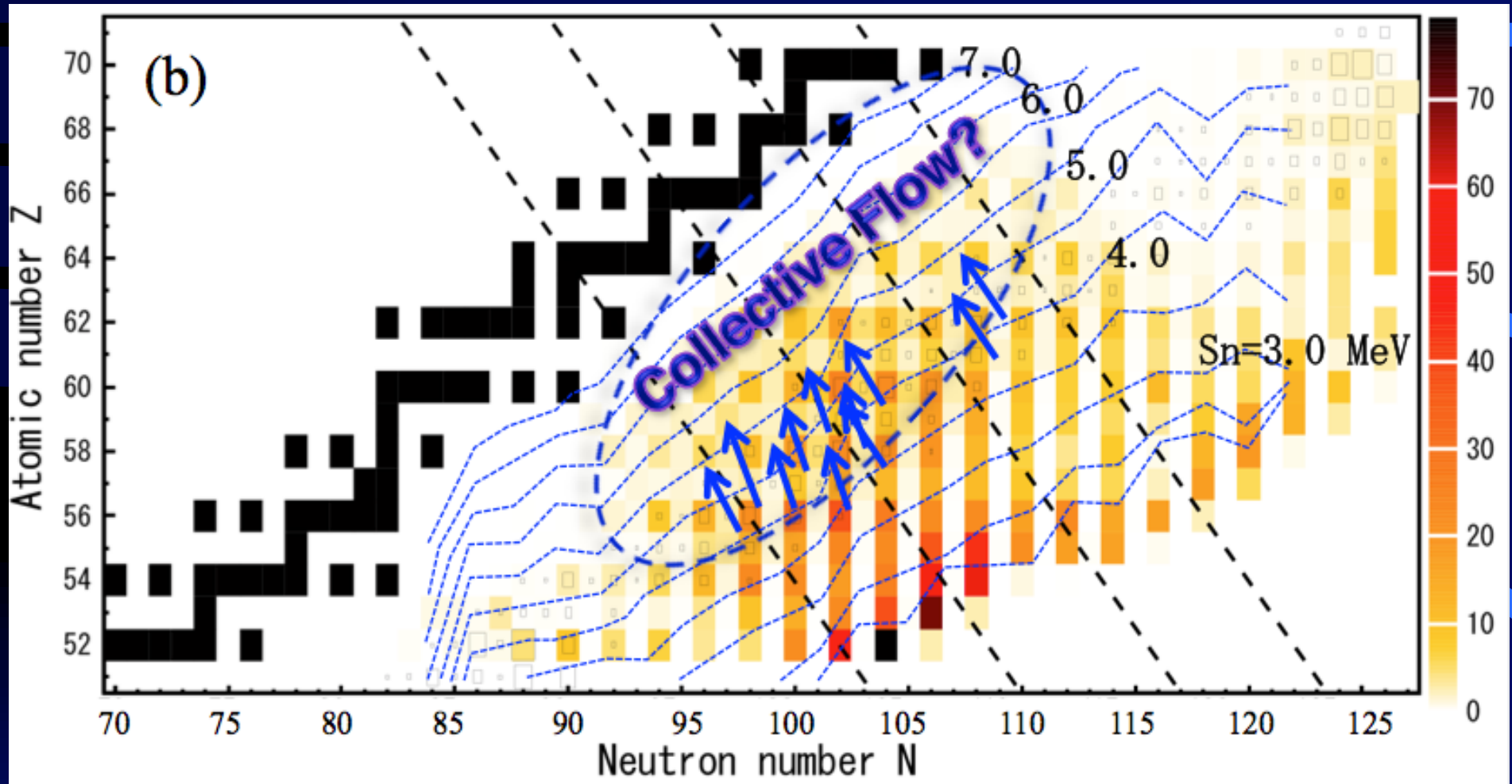
Universality of r-process elements ($Z \geq 56$)



Decay Spectroscopy around mass $A=160$ was performed !

Rare-Earth Peak Formation at Freeze-Out Time

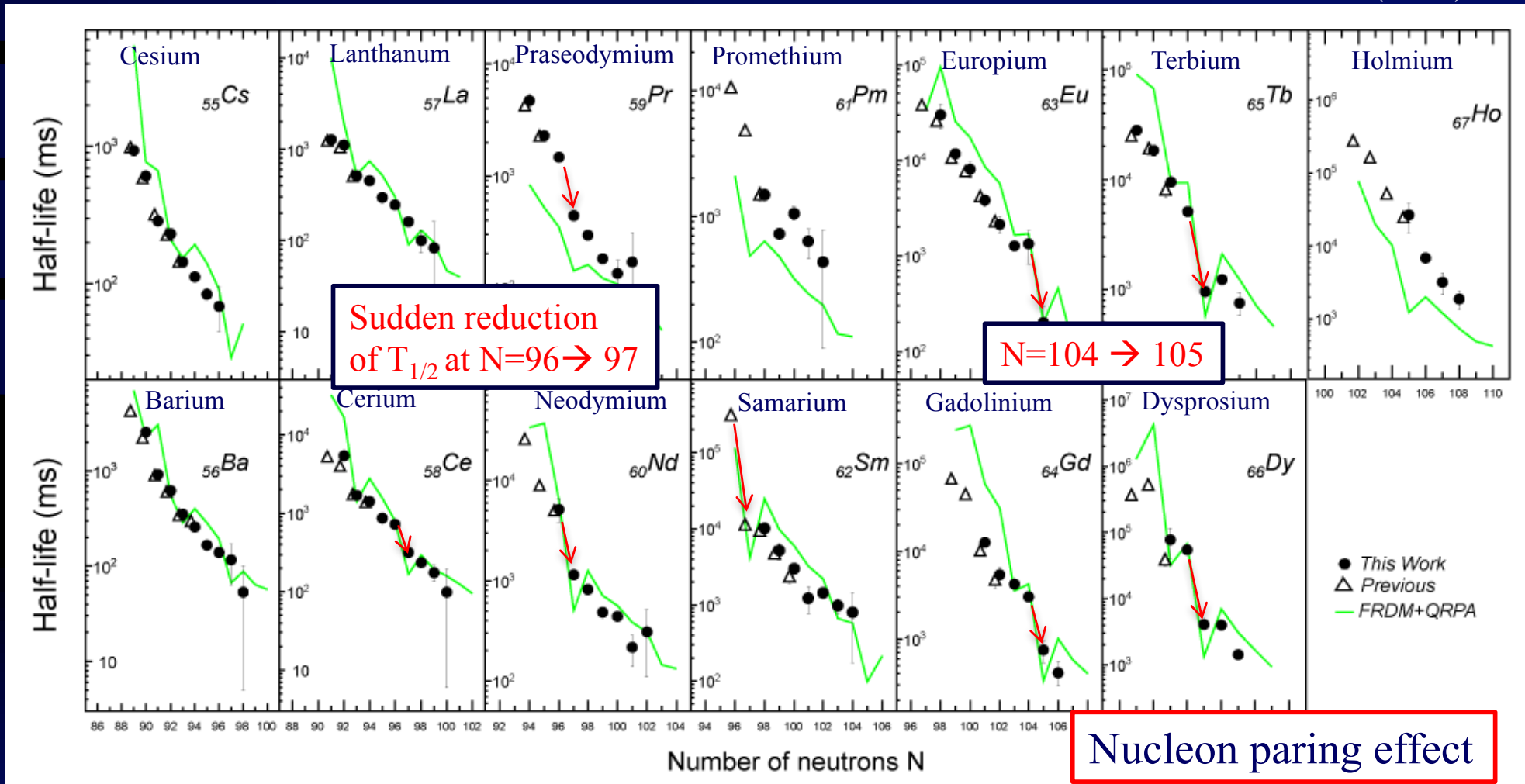
R. Surman et al. PRL 79 (1997)



Test r-Process calculation with new RIBF data

92 β -Decay Half-lives (Mass A = 144 – 175) vs FRDM+QRPA

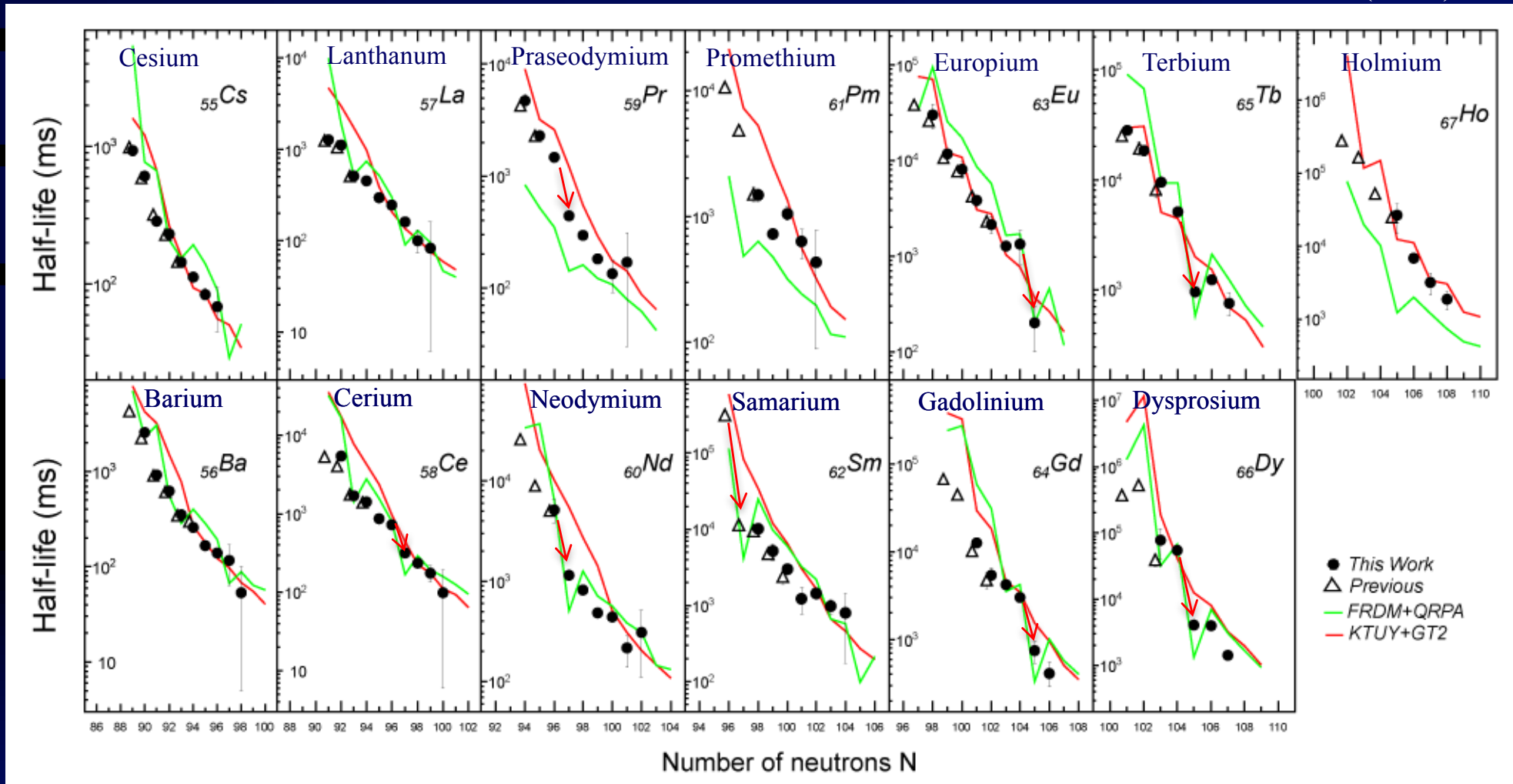
J. Wu, PRL (2017)



The β -decay half-lives of 92 neutron-rich $^{144-151}\text{Cs}$, $^{146-154}\text{Ba}$, $^{148-156}\text{La}$, $^{150-158}\text{Ce}$, $^{153-160}\text{Pr}$, $^{156-162}\text{Nd}$, $^{159-163}\text{Pm}$, $^{160-166}\text{Sm}$, $^{161-168}\text{Eu}$, $^{165-170}\text{Gd}$, $^{166-172}\text{Tb}$, $^{169-173}\text{Dy}$, and $^{172-175}\text{Ho}$ were measured at the Radioactive Isotope Beam Factory (RIBF).

92 β -Decay Half-lives (Mass A = 144 – 175) vs KTUY+GT2

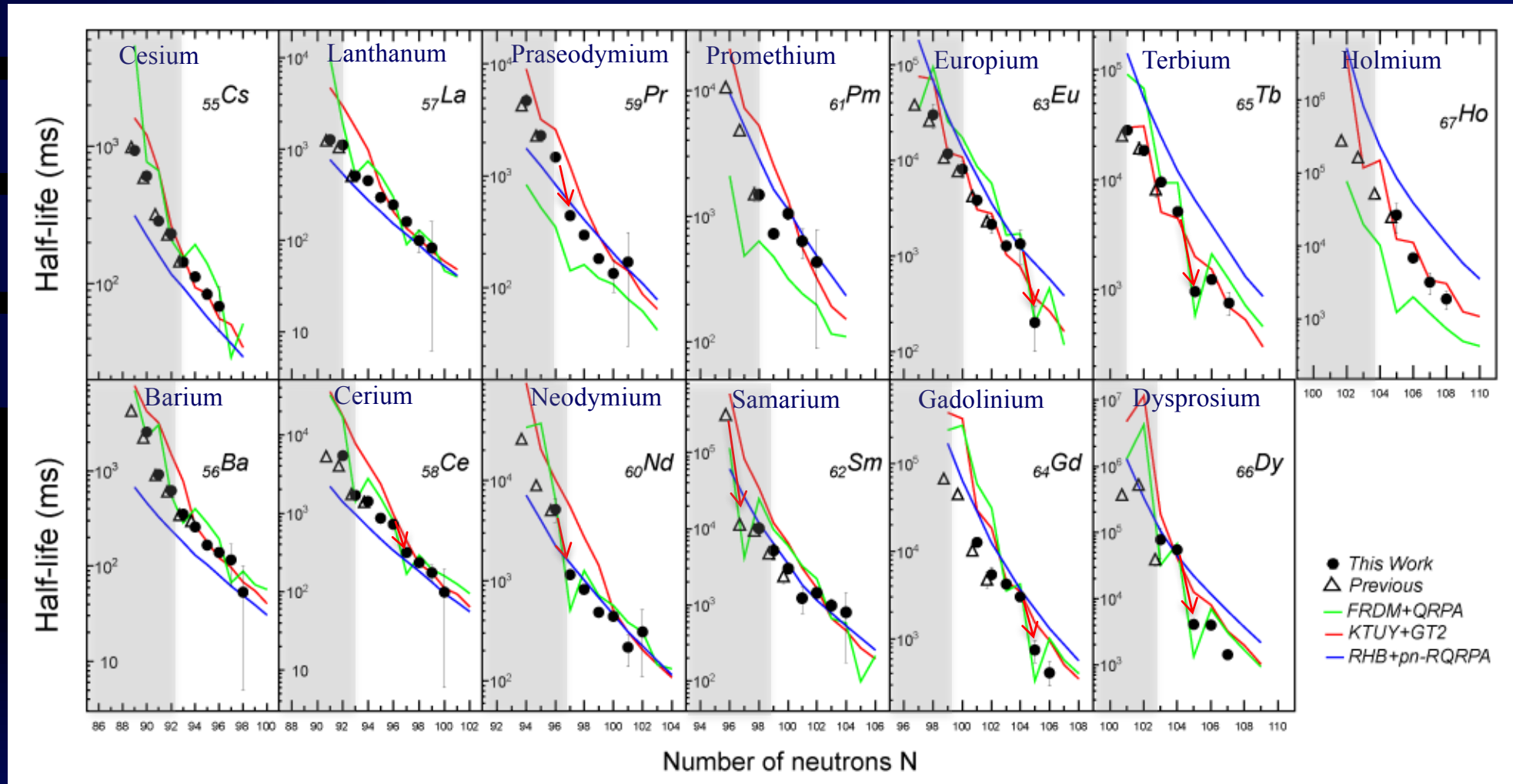
J. Wu, PRL (2017)



KTUY+GT2 shows systematic trends similar with exp. values in mid-shell.
 However, it can not reproduce even-odd effects.

92 β -Decay Half-lives (Mass A = 144 – 175) vs RHB+pn-RQRPA

J. Wu, PRL (2017)

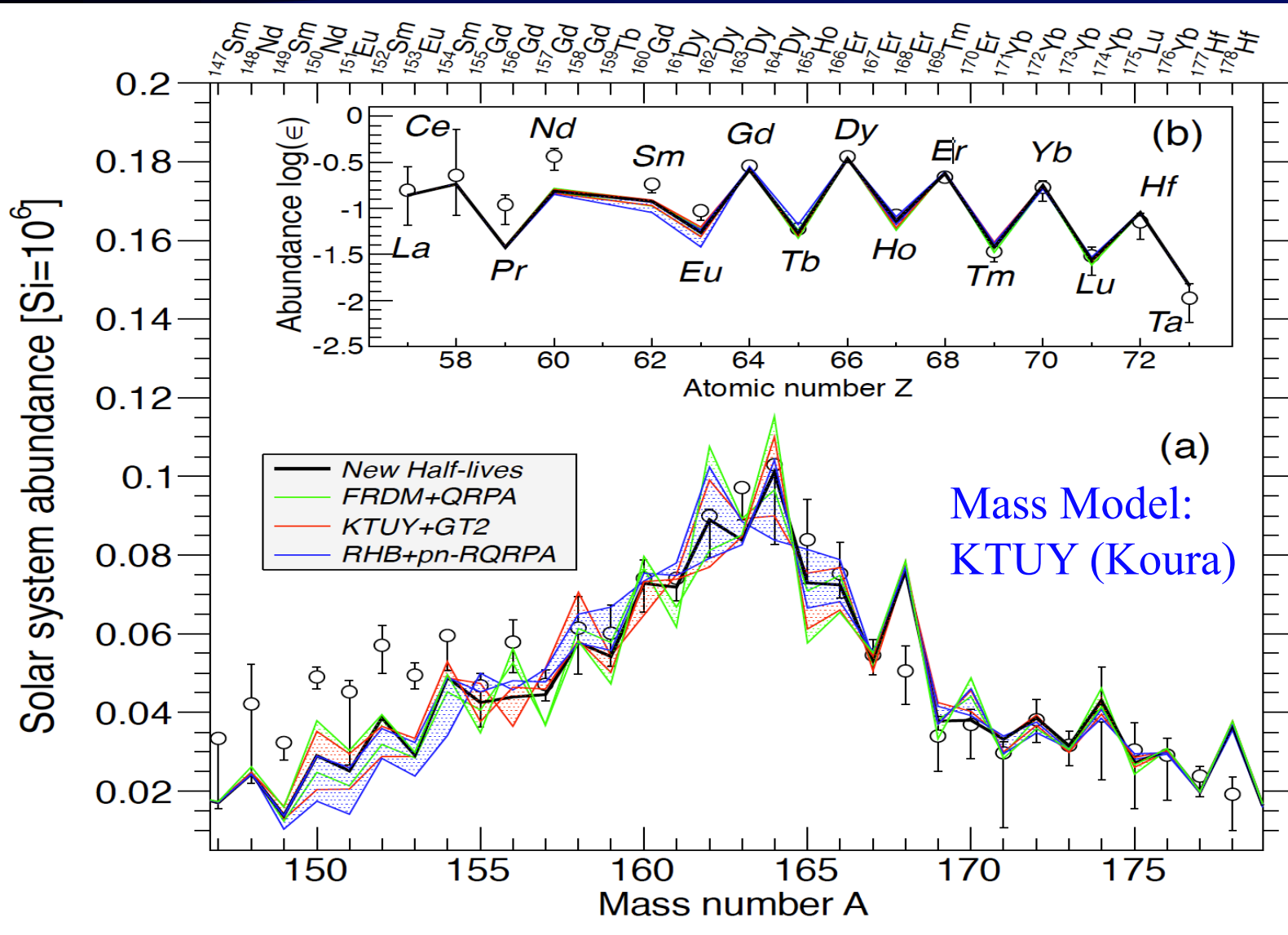


RHB+pn-RQRPA underestimates the $T_{1/2}$ for $Z=55-56$ (Cs, Ba) and overestimates the $T_{1/2}$ for $Z=63-67$ (Eu, Gd, Tb, Dy, Ho).
No reproduce of even-odd effects.

r-Process Elemental Abundance ($A = 150-179$)

Uncertainty of β -decay half-lives \rightarrow r-Process Rare-Earth Peak

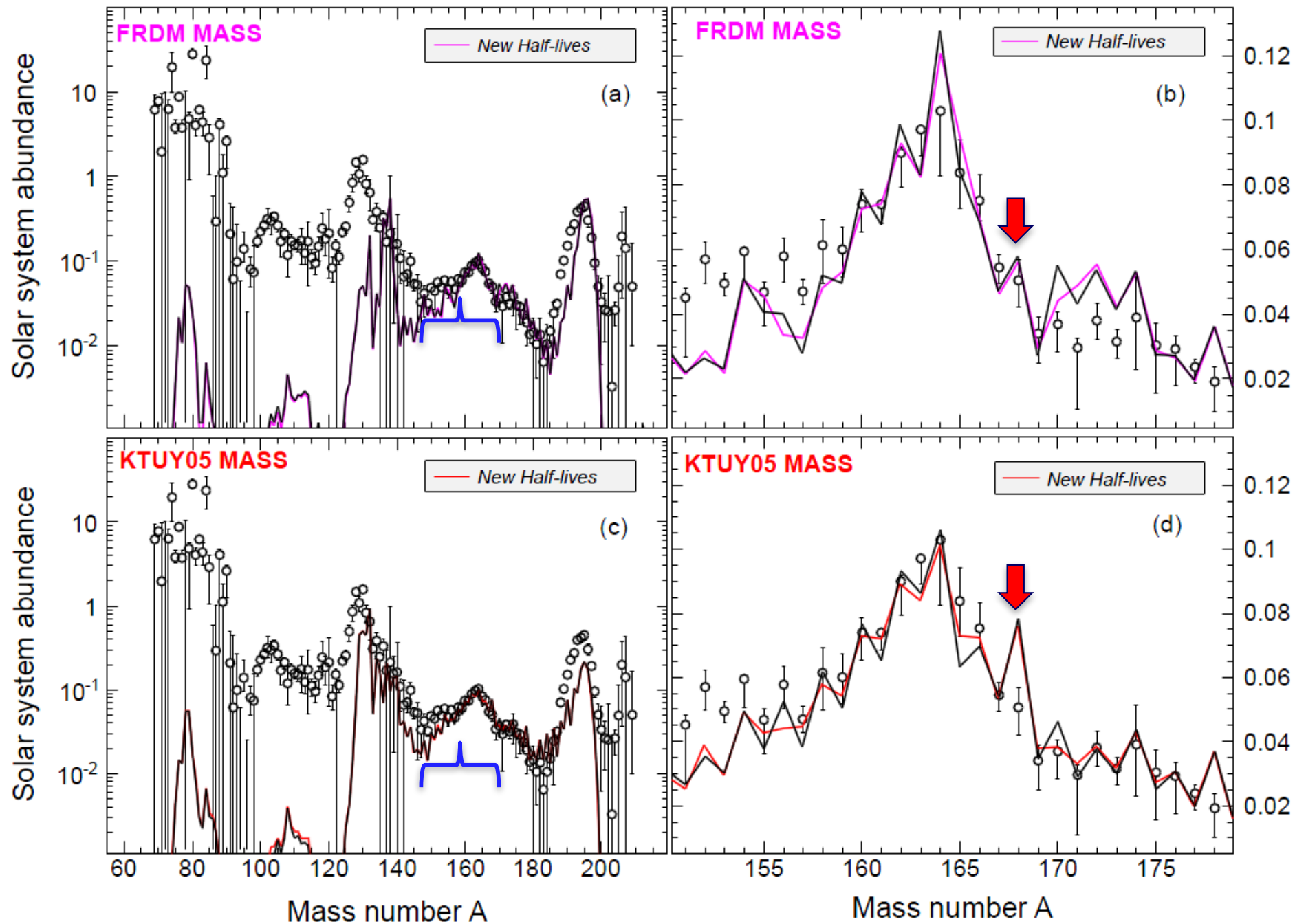
J. Wu et al. PRL (2017)



Experimental Feedback and Evaluation of the r-Process Rare-Earth Peak Formation

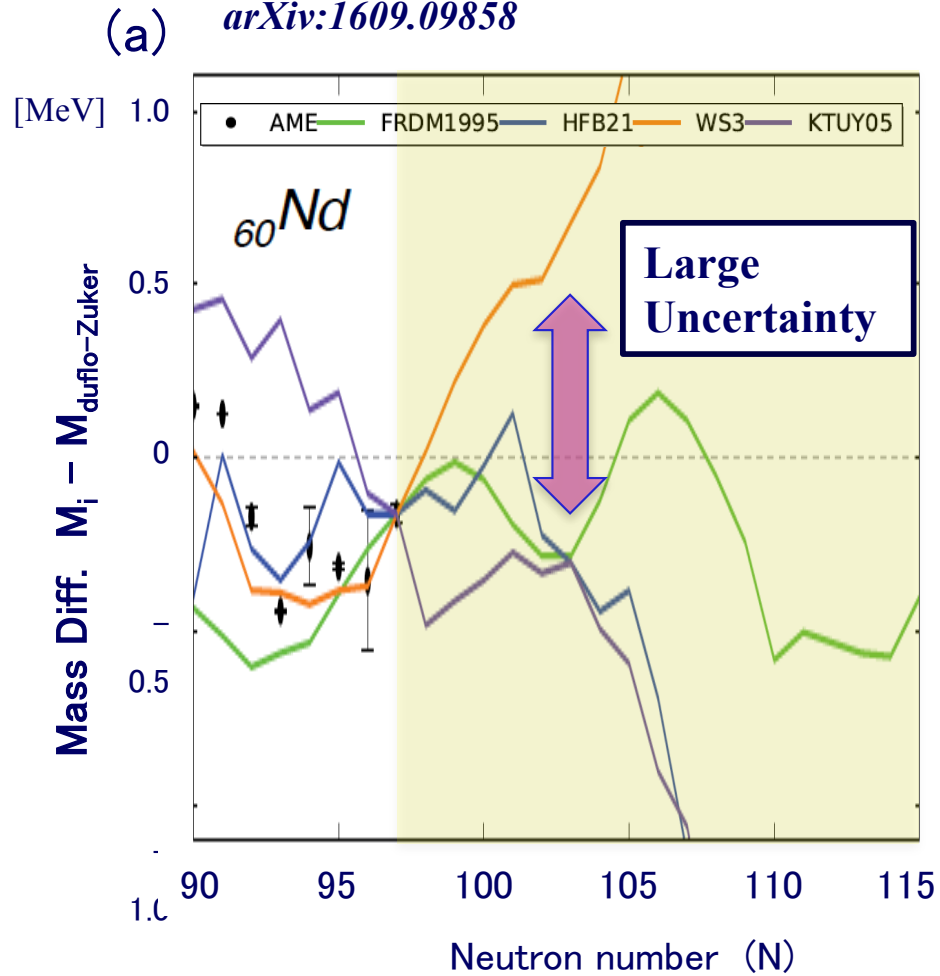
Rare-Earth Peak with Different Mass Models

J. Wu

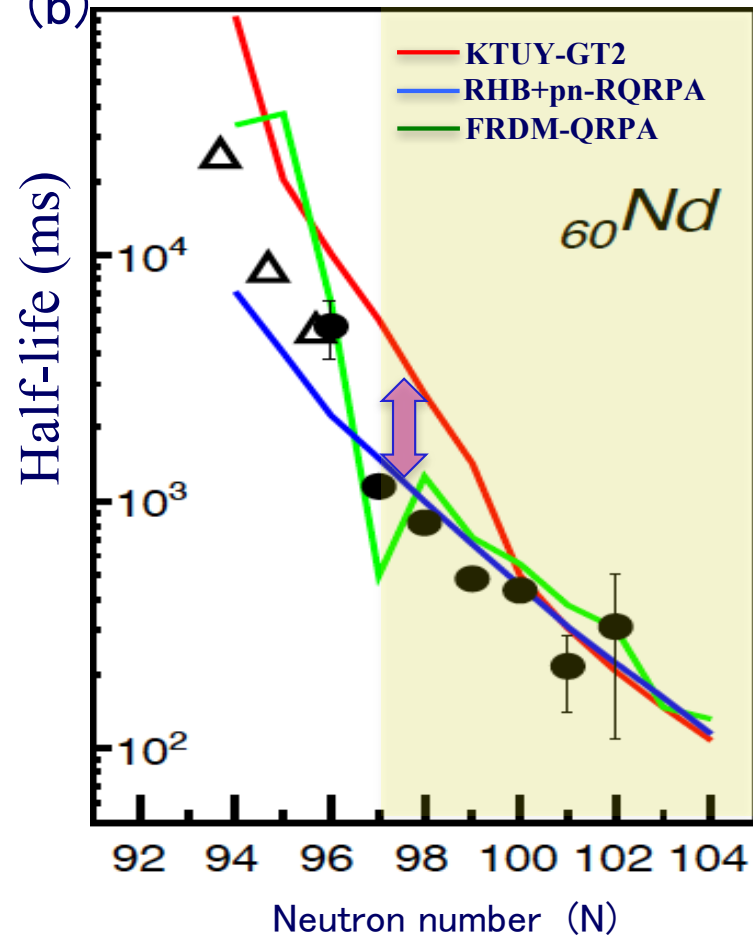


Mass & beta-decay half-lives

M.R.Mumpower, R. Surman et al.
arXiv:1609.09858

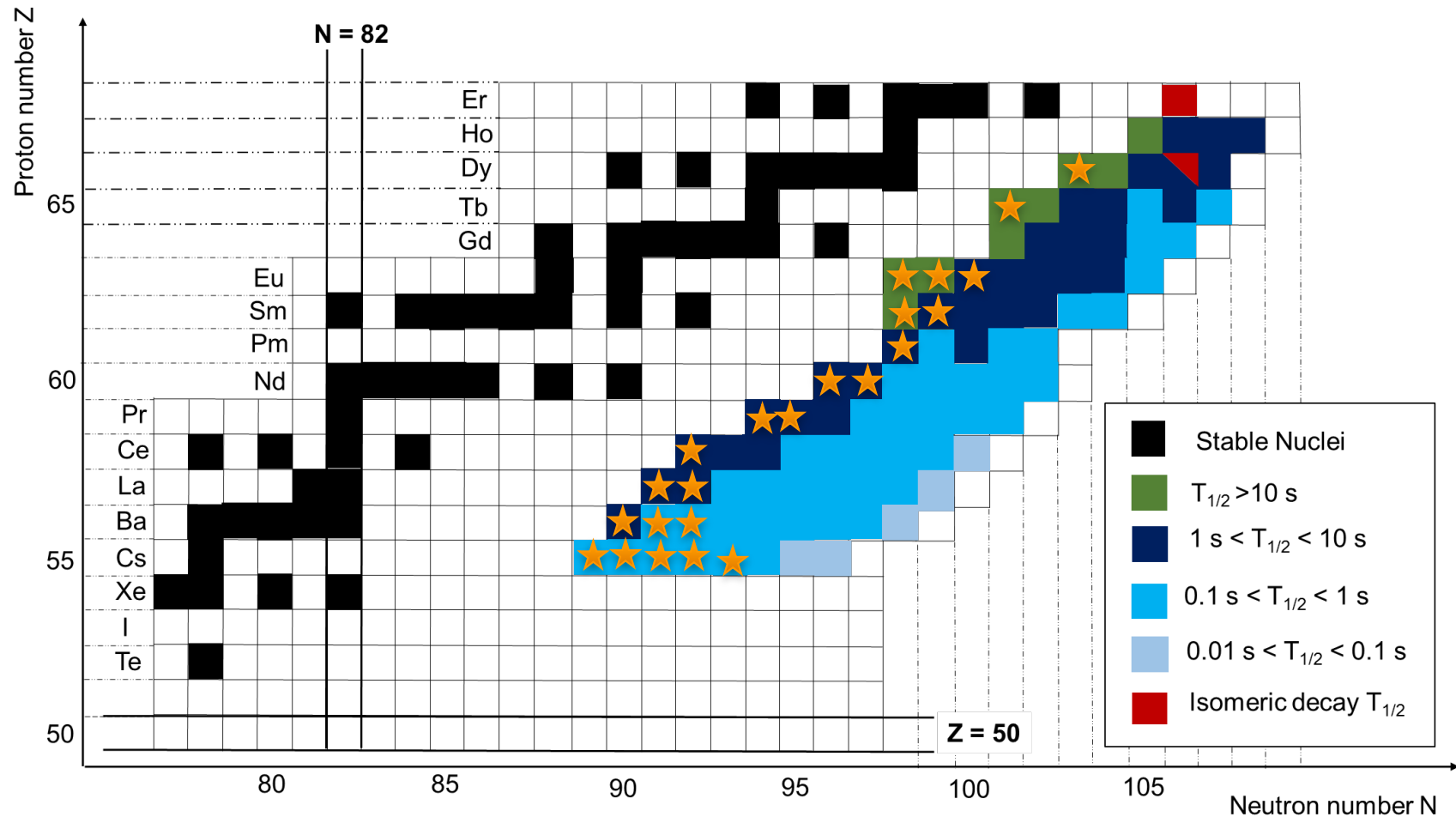


(b) *J. Wu: PRL (2017)*



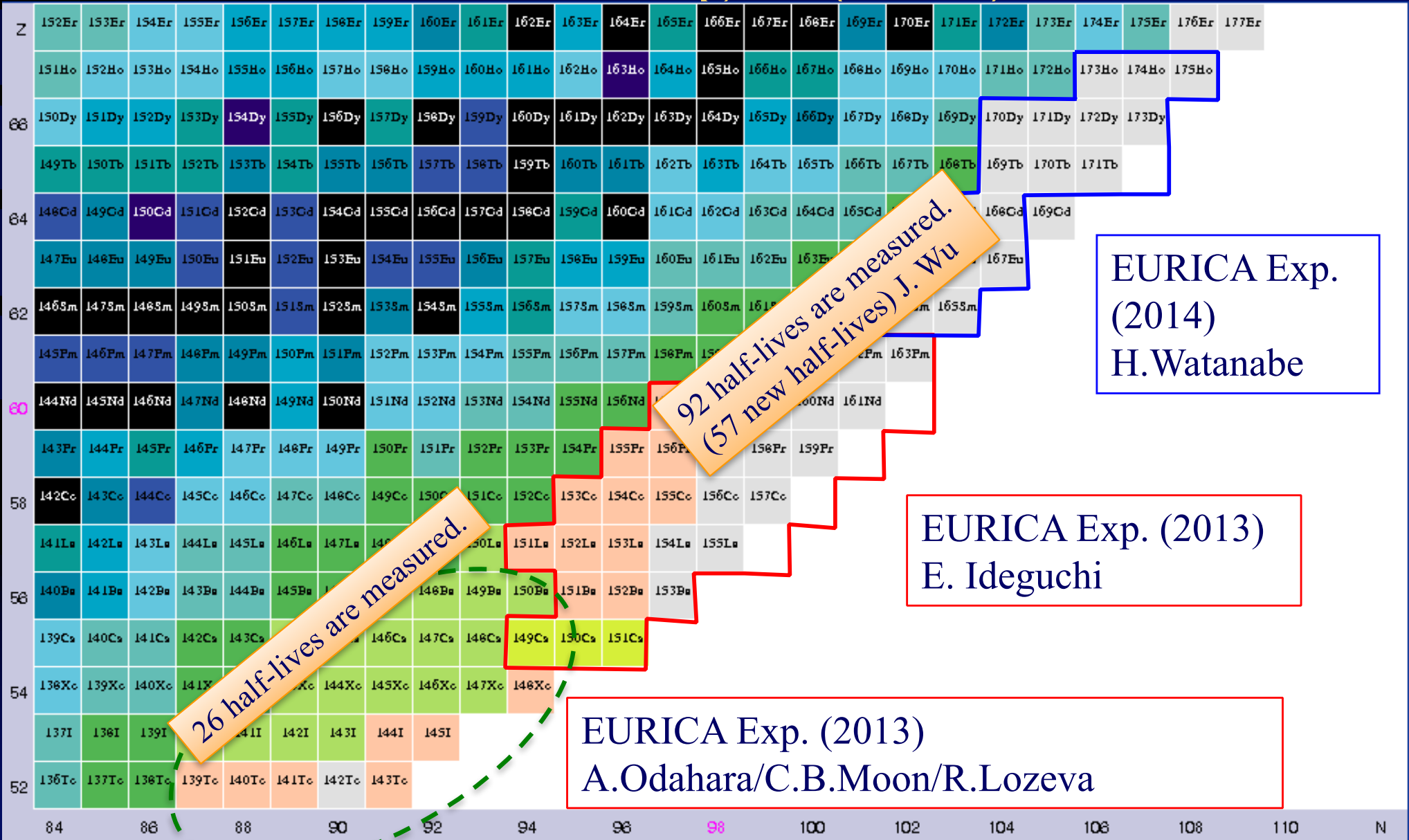
Half-lives and Mass Measurement

★ mass (known)



Half-lives


Heavier region ($Z > 50$)



From Now:
beta-delayed neutron
emission probabilities P_n

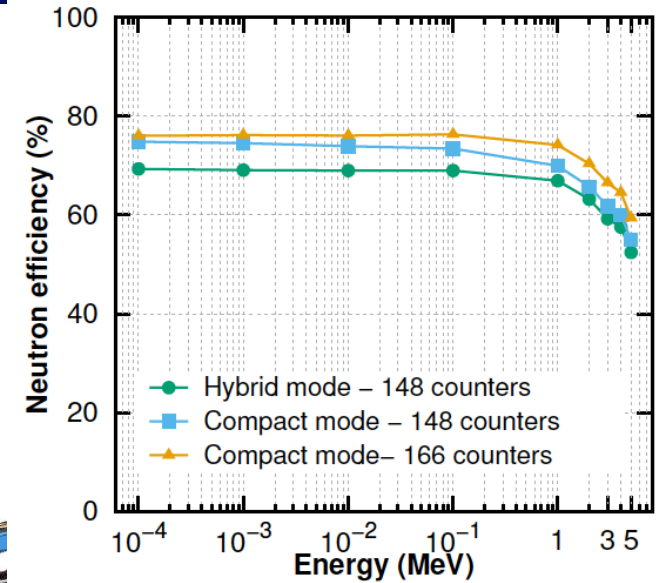
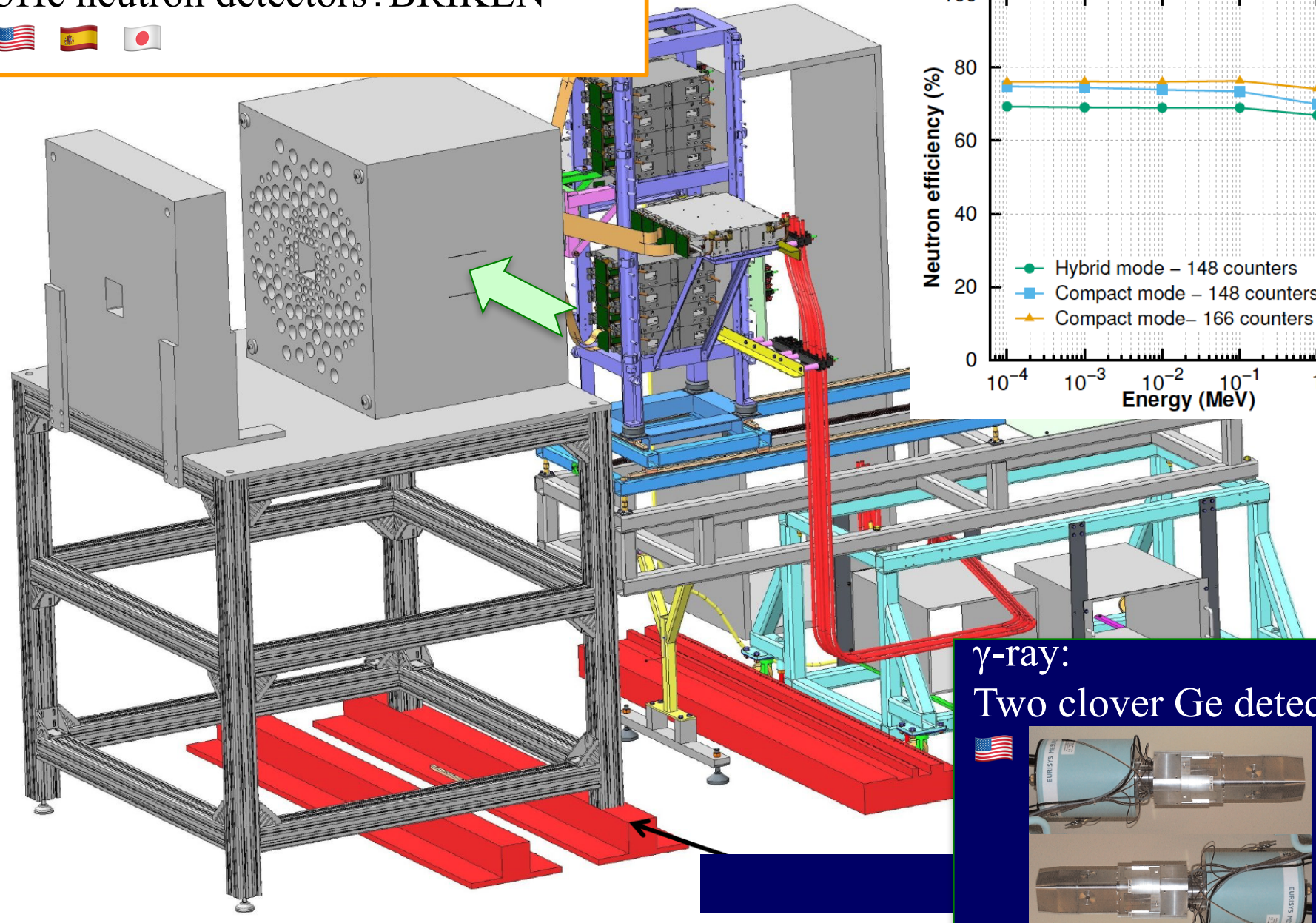
BRIKEN (2017 ~)

BRIKEN Setup

beta counting system: AIDA 

↔ WAS3ABi 

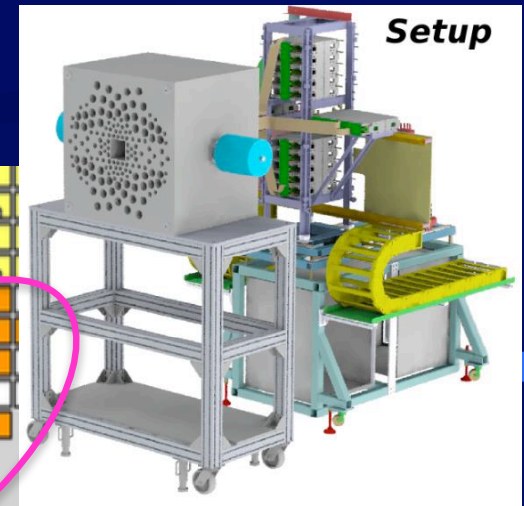
^3He neutron detectors: BRIKEN



γ -ray:
Two clover Ge detectors



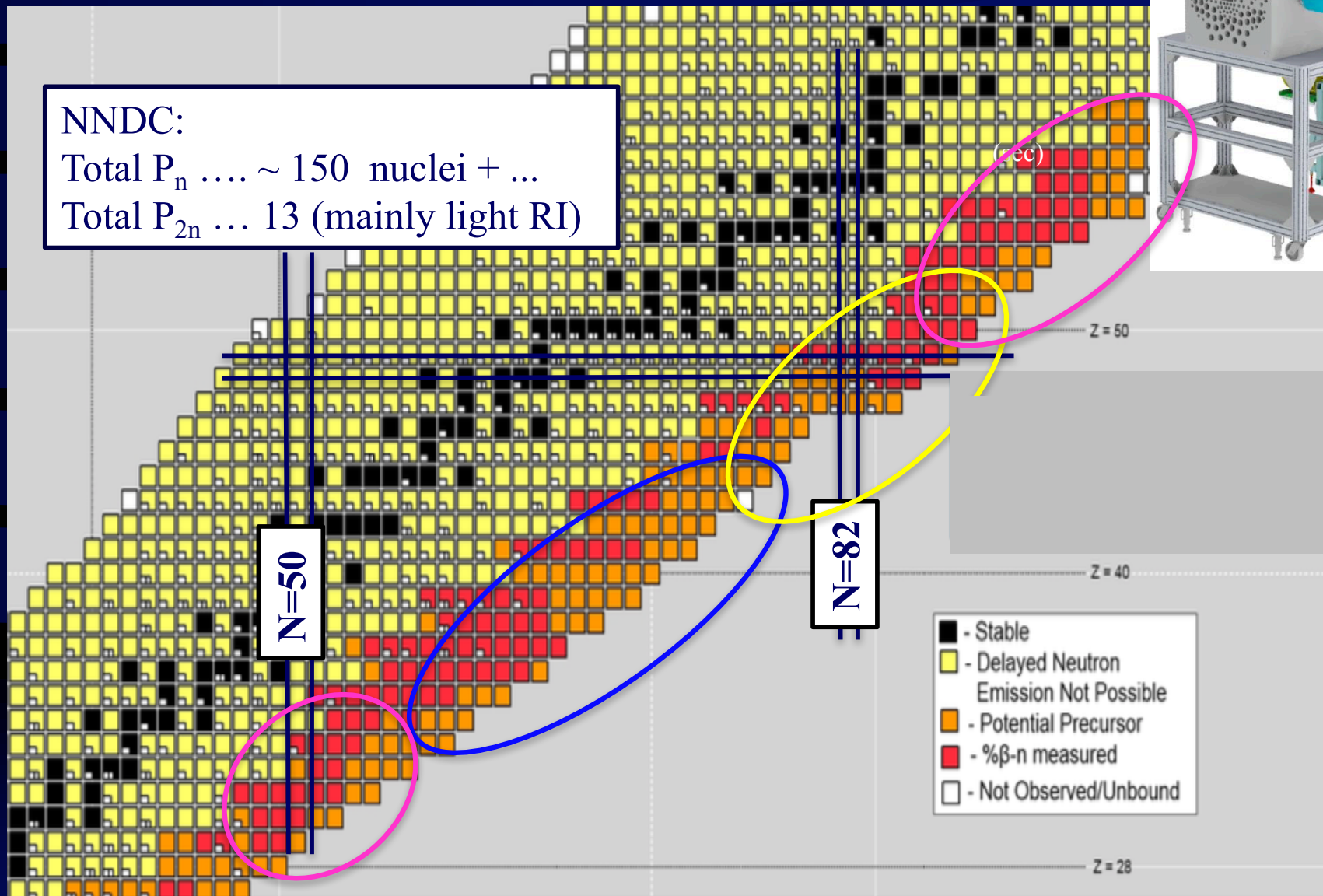
BRIKEN Project (2016 ~)



NNDC:

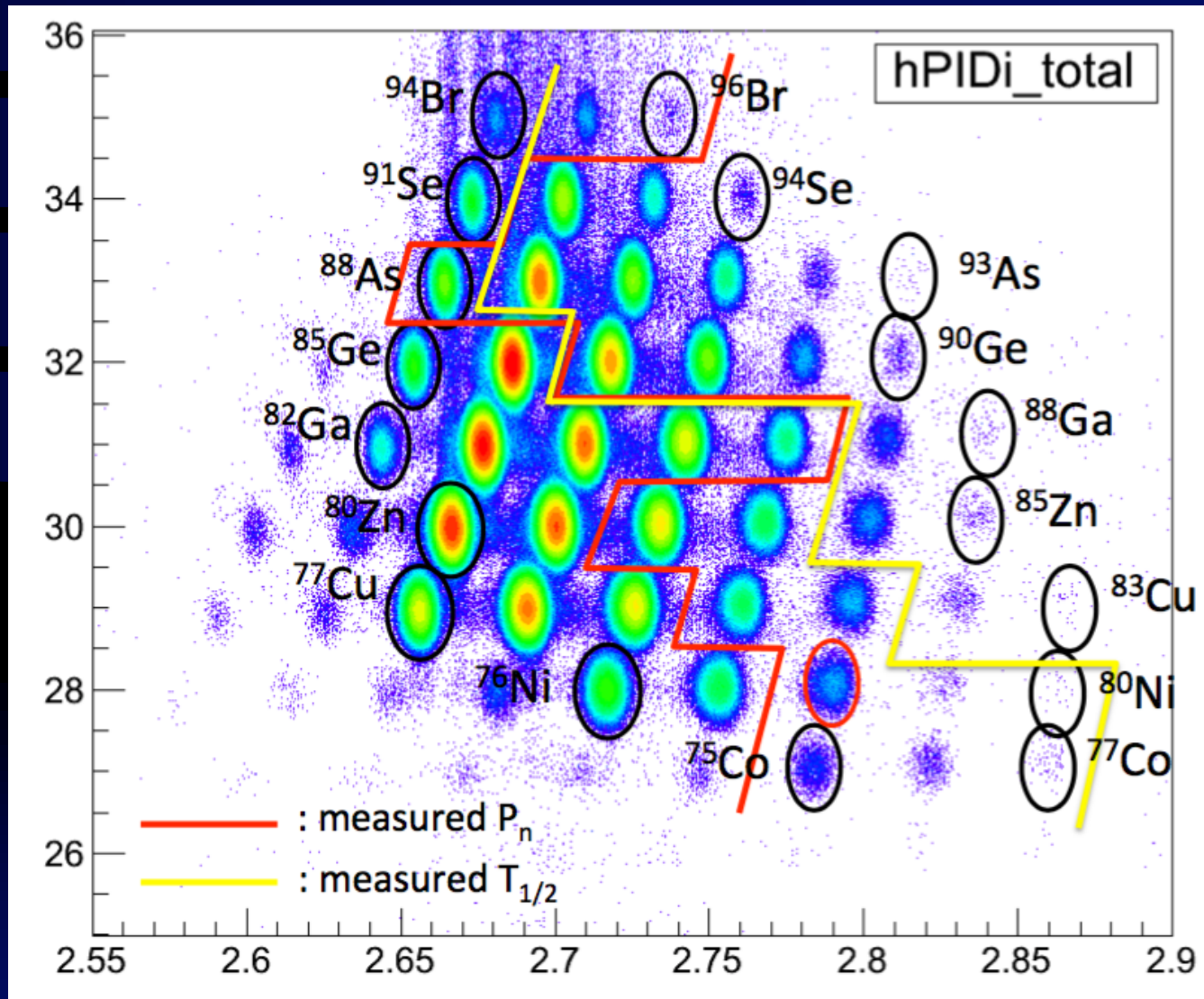
Total P_n ~ 150 nuclei + ...

Total P_{2n} ... 13 (mainly light RI)

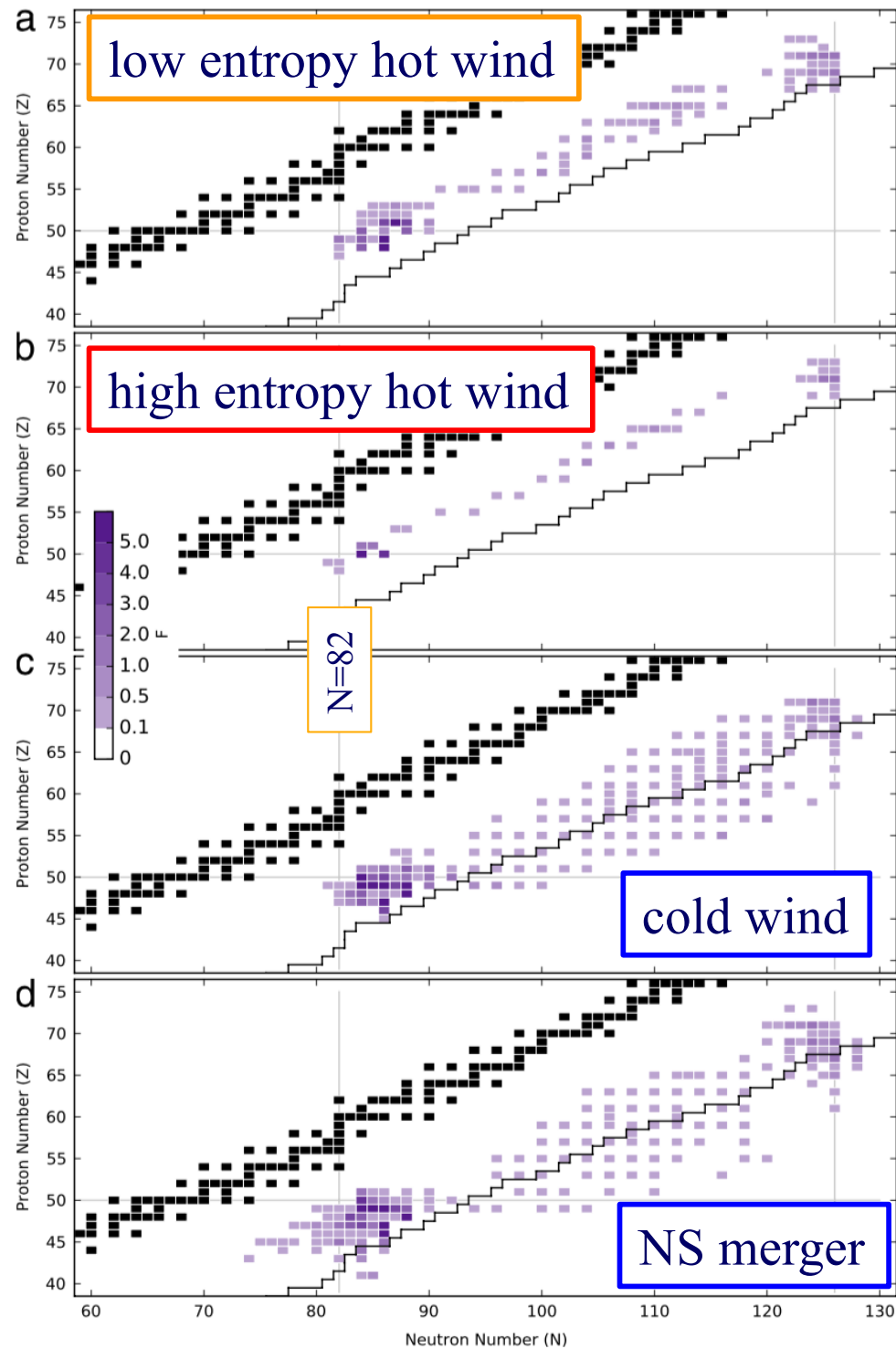


Several hundreds of beta-delayed neutron emission P_n (n) together with $T_{1/2}$ (β) & level scheme (γ)

BRIKEN: β -delayed neutron emissions



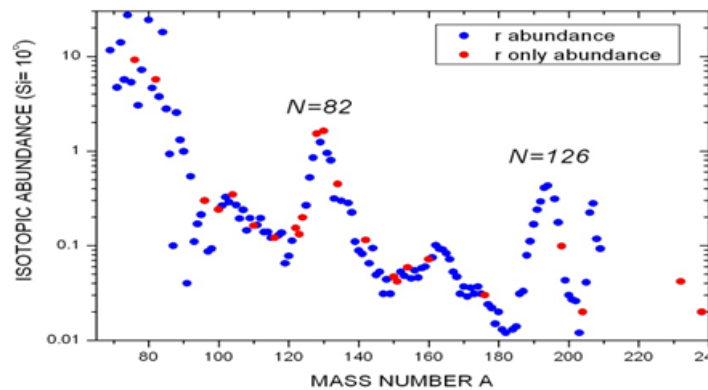
Pn values: How Sensitive in r-Process Calc.?



beta-delayed neutron emitters

M.R. Mumpower et al.
Prog. in Part. and Nucl. Phys
86 (2016) 86-126

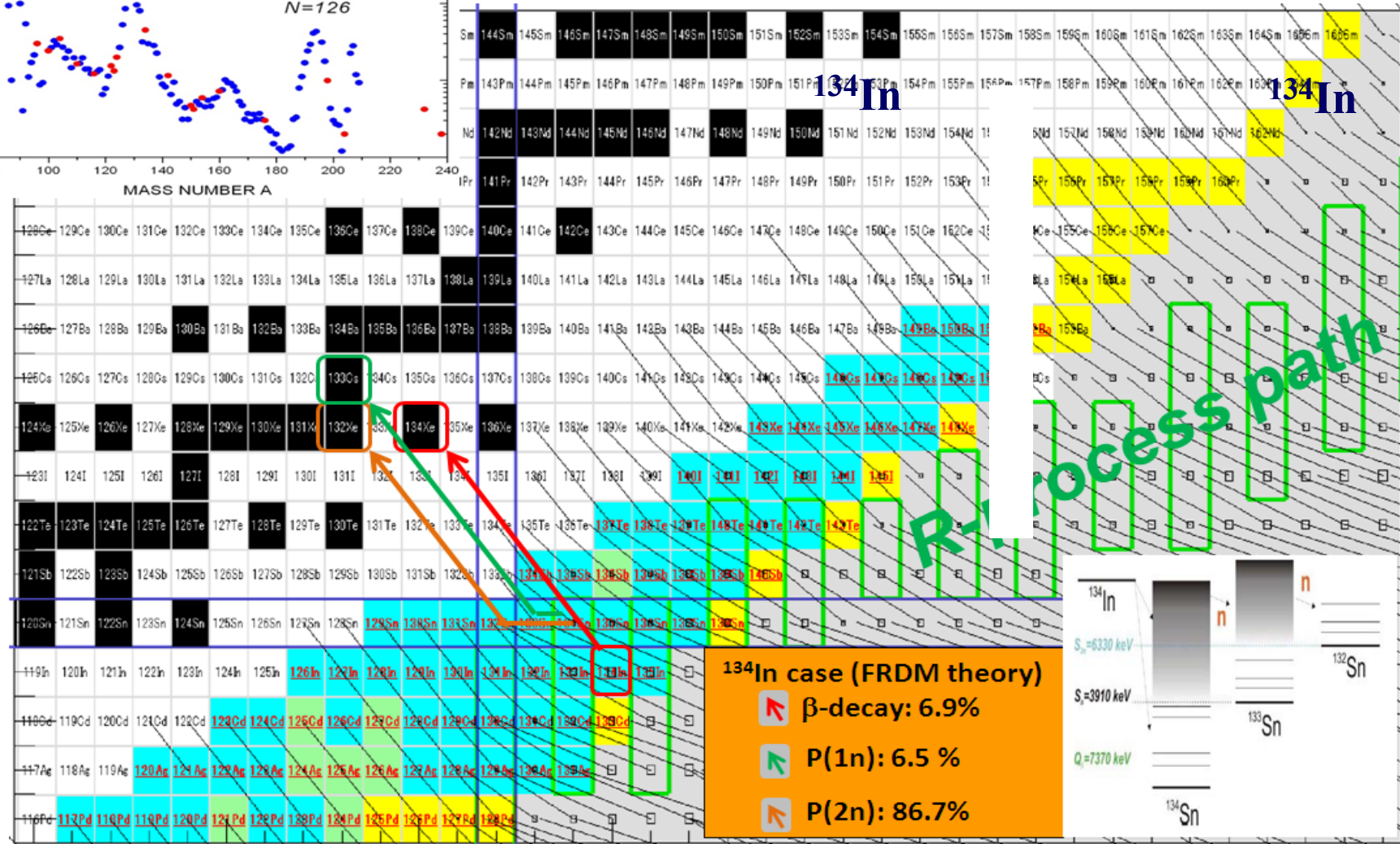
r-Process Elements : Freeze-out Time



During „Freeze-out“: detour of β -decay chains

\Rightarrow *r*-abundance changes

Proton number (Z)



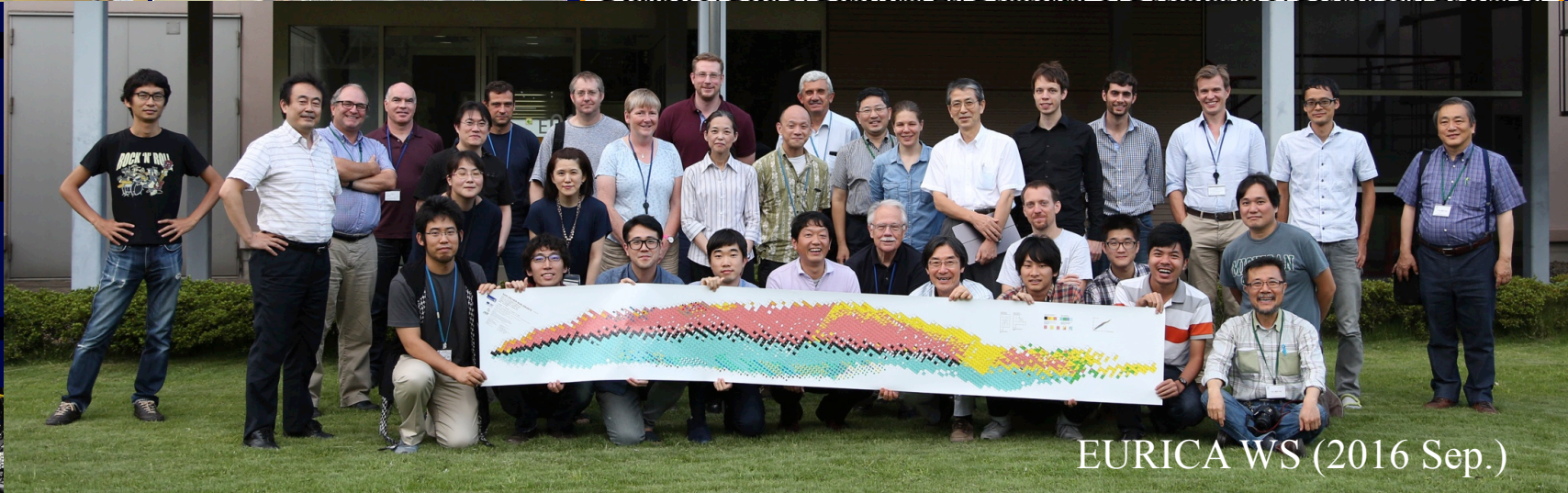
Neutron number (N)

EURICA Collaboration



19 countries : 237 collaborators

J. Agramunt, P. Aguilera, T. Alharbi, A. Algora, G. Angelis, N. Aoi, P. Ascher, R. Avigo, H.Baba, C. Borcea, A. Boso, A.M. Bruce, R.B. Cakirli, F.L.Bello Garrote, G. Benzoni, J.S.Berryman, R. Berta, B. Blank, N. Blasi, A. Blazhev, P. Boutachkov, S. Bonig, A. Bracco, F. Browne, F. Camera, R.J. Carroll, S. Ceruti, I. Celikovic, K.Y. Chae, J. Chiba, L. Coraggio, A. Covello, F.C.L. Crespi, J.-M. Daugaus, R. Daido, P. Davis, M.C. Delattre, F. Diel, F. Didieiean, Zs. Dombradi, P. Doornenbal, F. Drouet, H.J. Eberth, A.



EURICA WS (2016 Sep.)



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D.R. Napoli, F. Naqvi, M. Niikura, H. Nishizuka, S. Nishimura, T. Nishizuka, C. Nita, T. Nowacki, A. Odahara, K. Ogawa, H. Oikawa, R. Orlandi, S. Ota, T. Otsuka, H.J. Ong, S. Orrigo, M. Rajabali, J. Park, Z. Patel, A. Petrovici, F. Recchia, V. Phong, Zs. Podolyak, O.J. Rovers, L. Prochniak, P.H. Regan, S. Rice, E. Sahin, H. Sakurai, K. Sato, H. Schaffner, H.Scheit, P. Schury, C. Shand, Y. Shi, S. Shibagaki, T. Shimoda, Y. Shimizu, K. Sieja, L. Sinclair, G.S. Simpson, P.-A. Soderstrom, D. Sohler, I.G. Stefan, K. Steiger, D. Steppenbeck, K. Sugimoto, T. Sumikama, D. Suzuki, H. Suzuki, T. Tachibana, K. Tajiri, S. Takano, A. Tashima, H. Takeda, Man. Tanaka, Mas. Tanaka, Y. Takei, R. Taniuchi, J. Taprogge, K. Tajiri, T. Teranishi, S. Terashima, G. Thiamova, K. Tshoo, Zs. Vajta, J. Valiente Dobon, Y. Wakabayashi, P.M. Walker, H. Watanabe, A. Wendt, V. Werner, O. Wieland, K. Wimmer, J. Wu, Q. Wu, F.R. Xu, Z.Y. Xu, A. Yagi, S. Yagi, H. Yamaguchi, K. Yamaguchi, T. Yamamoto, M. Yalcinkaya, R. Yokoyama, S. Yoshida, K. Yoshinaga, G. Zhang

Acknowledgement: Gammapool, Prepec, IBS

BRIKEN Collaboration



Summary & Future

