

*High-z Metal Enrichment
Traced by Extinction Curves*

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1. Extinction as a Tracer of Metals

Half of metals in interstellar medium are in dust grains.

→Dust formation is important to trace **metal production**.

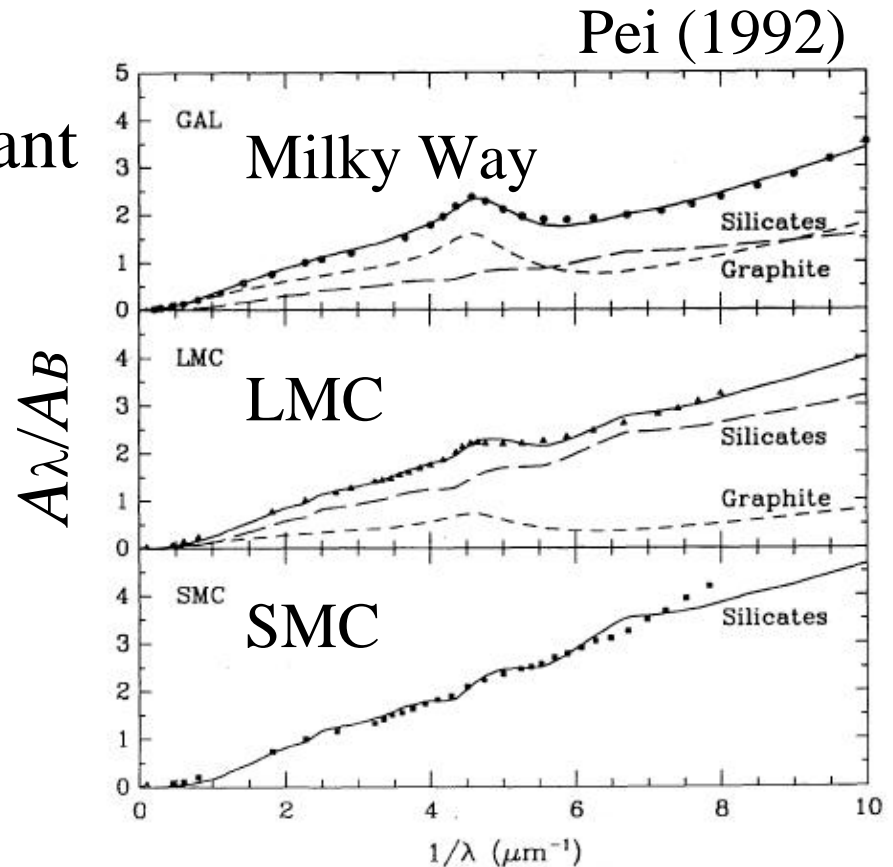
Extinction curve :

Wavelength dependence of dust cross section



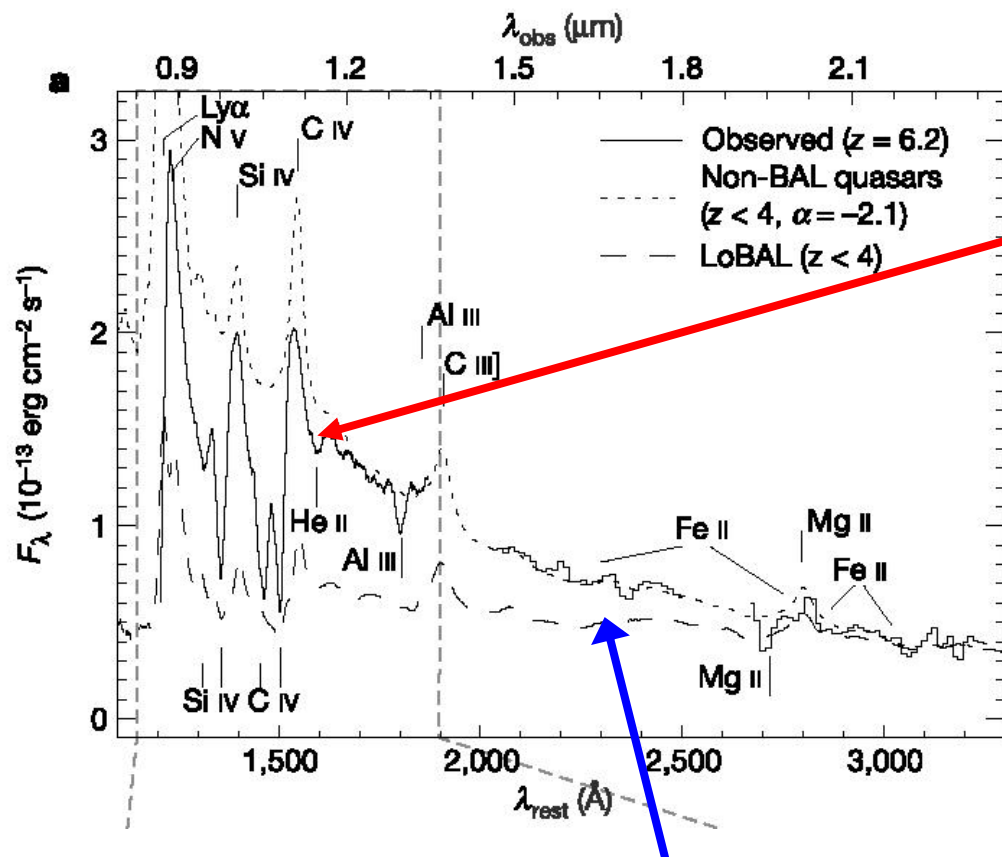
Key to understand the dust properties

(**composition**, **size distribution** etc.)

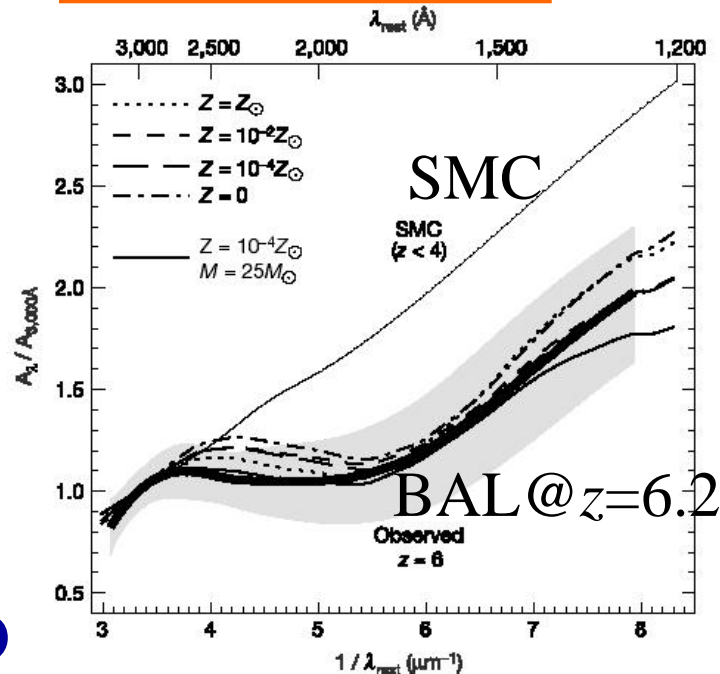


Extinction of a BAL QSO at $z = 6.2$

Maiolino et al. (2004)



BAL at $z = 6.2$



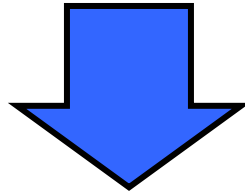
BAL at $z < 4$: Non-BAL SED + SMC extinction (reddening)

Different dust properties at high redshift.

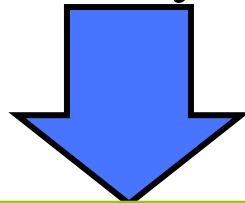
Origin of High- z Dust

Young cosmic age at $z > 5$

→ No dust supply from low-mass stars



Supernovae whose progenitors are massive stars with short lifetimes predominantly supply dust grains.

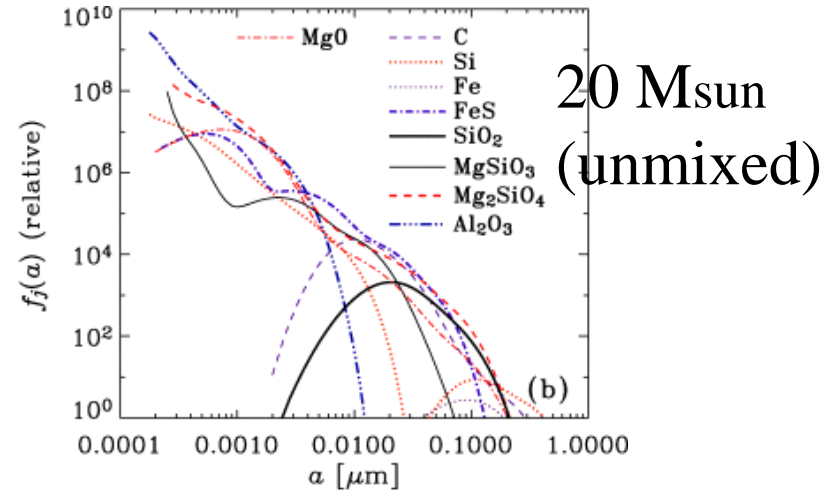
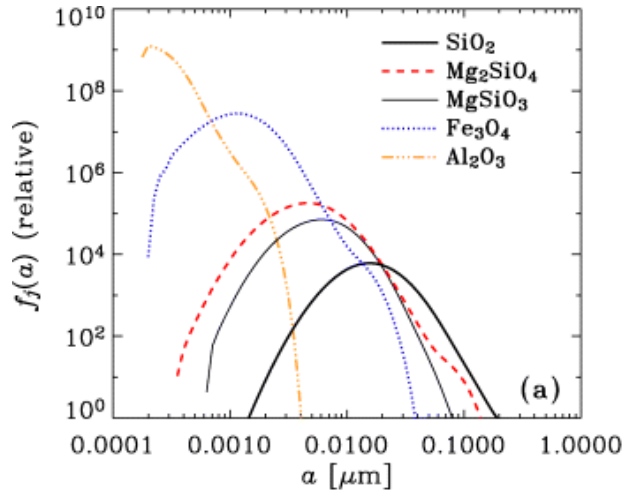


We calculate extinction curves of dust formed in SNe.
→ Compared to observations at $z = 6.2$
to constrain metal production in the early Universe.

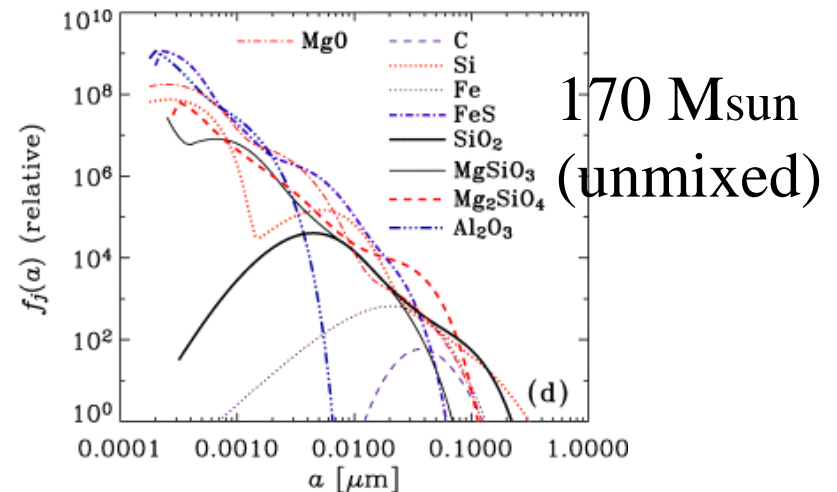
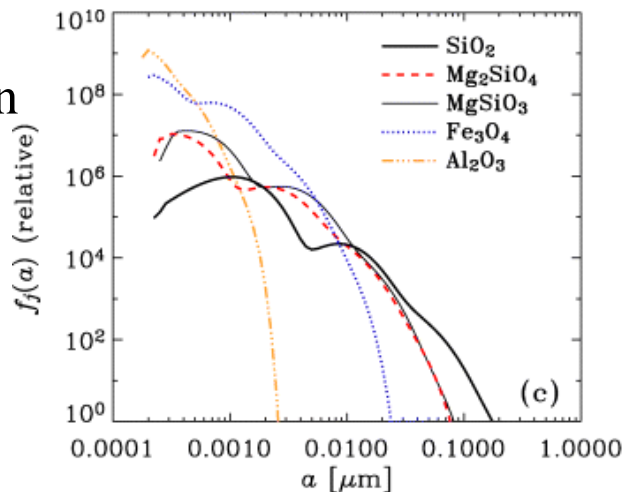
2. Dust Produced in SNe

Nozawa et al. (2003)

20 Msun
(mixed)

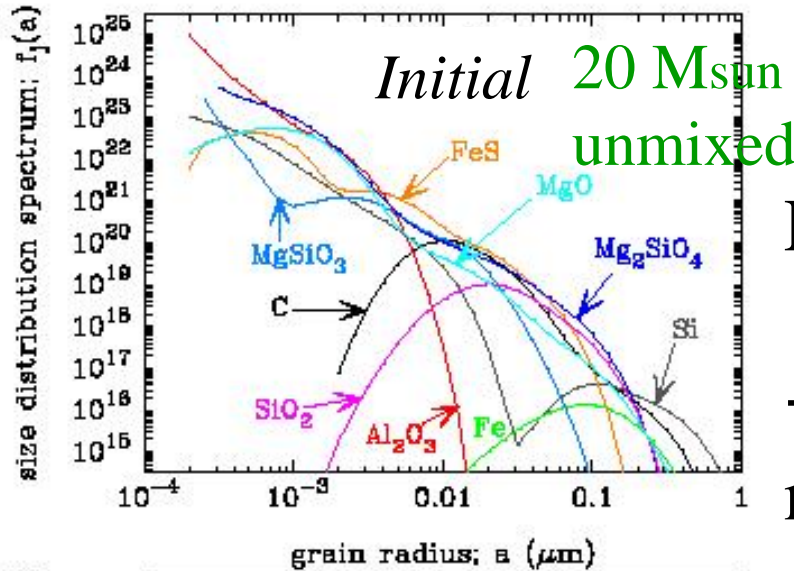


170 Msun
(mixed)

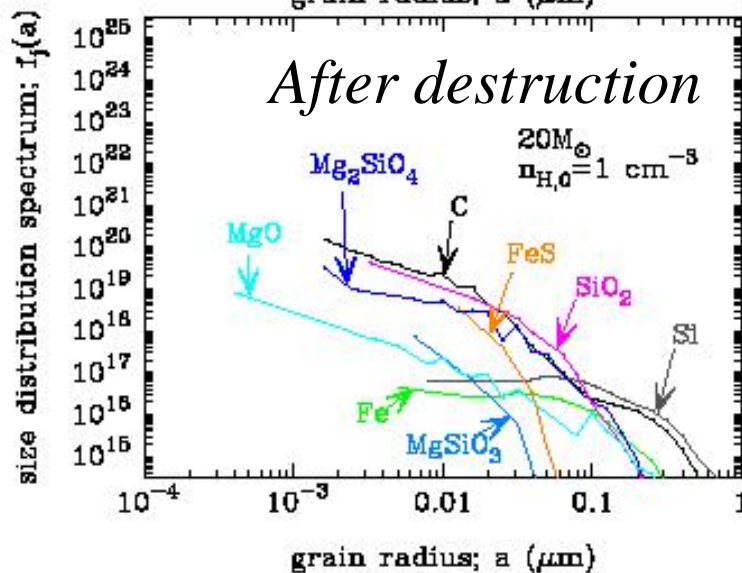


Effect of Dust Destruction

Nozawa et al. (2007)



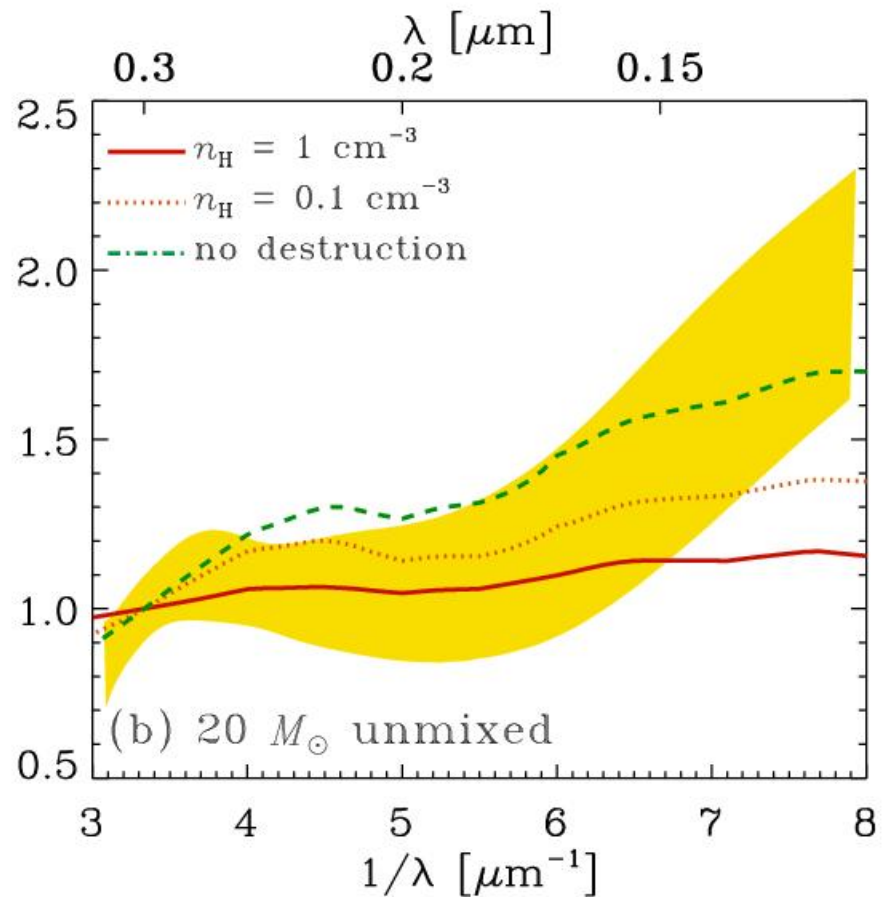
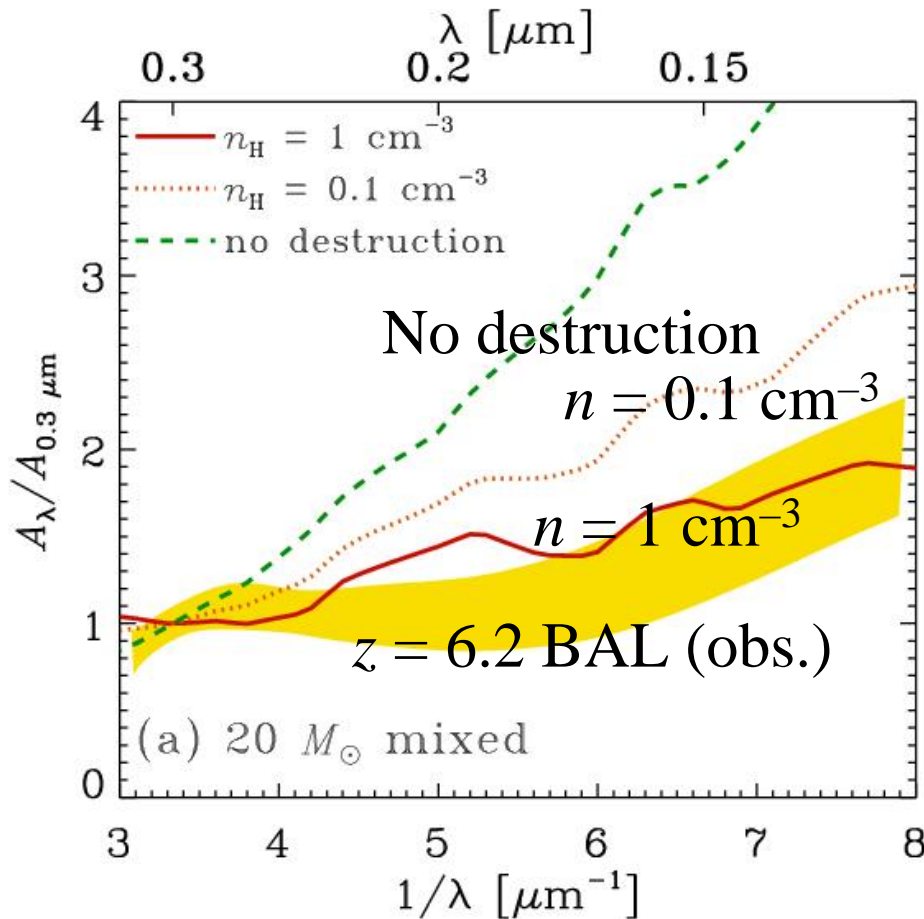
Destruction by reverse shock (Sputtering)
→ Small grains are destroyed more easily than large grains.



Extinction Curves tend to become flat.

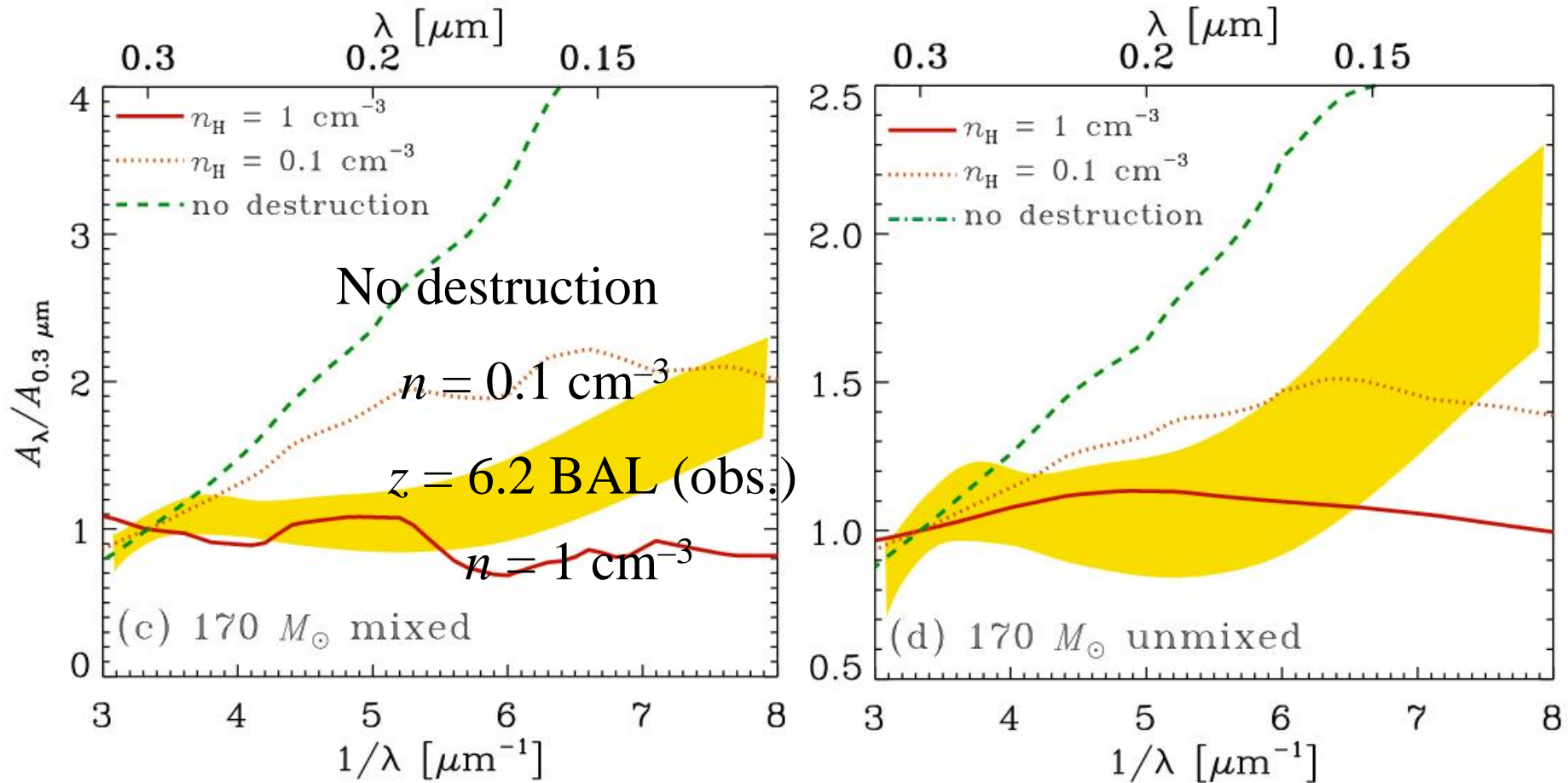
Effect of Dust Destruction

20 Msun progenitor



Effect of Dust Destruction

170 M_{\odot} progenitor



3. Summary and Discussion

Destruction by reverse shocks tend to flatten extinction curves.

The extinction curve observed for a $z = 6.2$ BAL can be explained by dust produced in SNe.

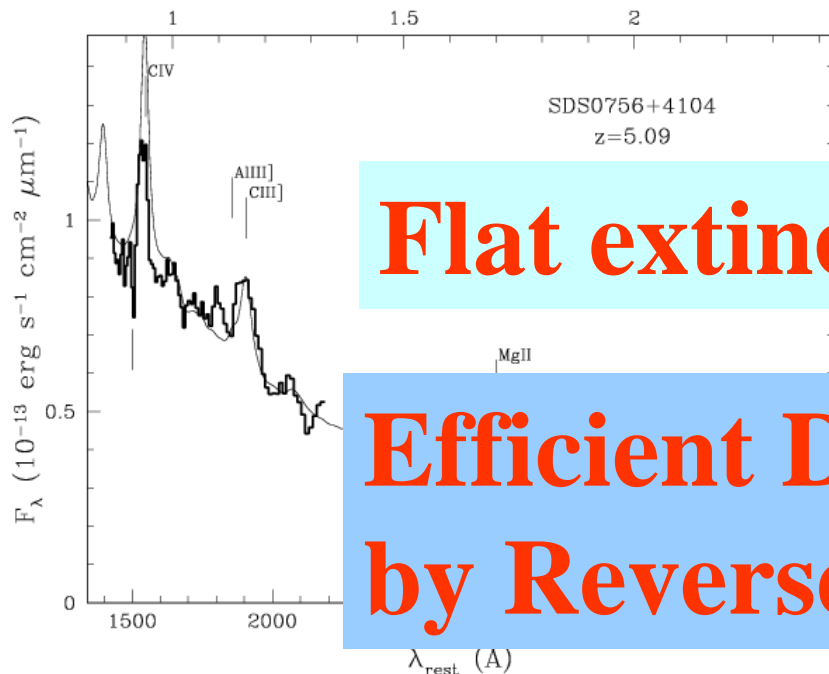
However, **too dense an environment such as $n > 1 \text{ cm}^{-3}$ flatten extinction curves too much.**

BALs without Reddening

Maiolino et al. (2004)

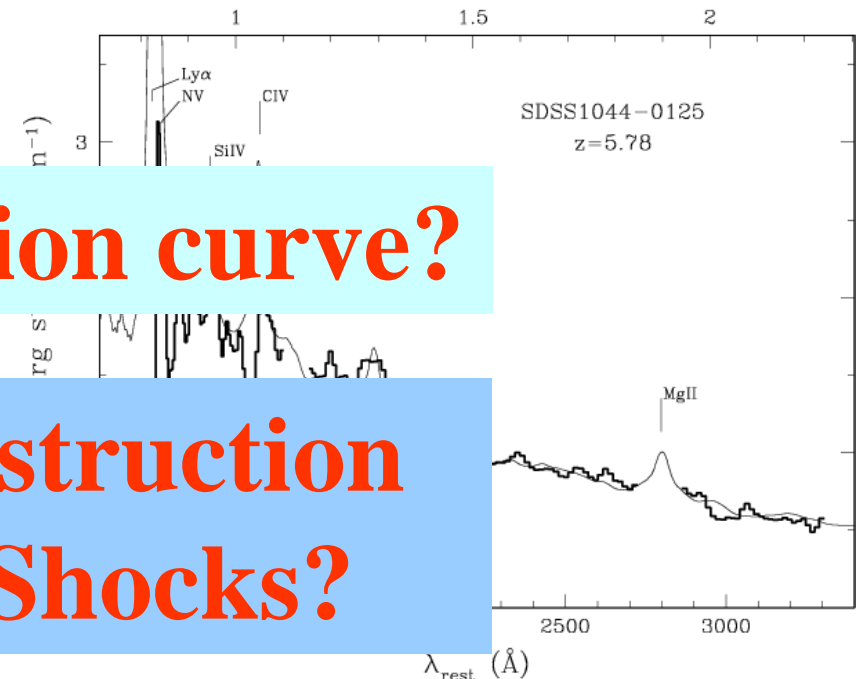
SDSS1044-0125 ($z = 5.78$)

$\lambda_{\text{obs}} (\mu\text{m})$



SDSS0756+4104 ($z = 5.09$)

$\lambda_{\text{obs}} (\mu\text{m})$



Flat extinction curve?

Efficient Destruction
by Reverse Shocks?

Consistent with non-BAL template \Rightarrow no reddening

But detected in sub-mm (10^8 – $10^9 M_{\text{sun}}$ dust: Priddey'03)