# Division of Astrophysics and Nuclear Physics: Nuclear Physics Group (Parallel session \#1) 

Takashi Nakatsukasa

@CCS, Univ. of Tsukuba, 2014.2.19

## Members of Nuclear Theory Group

- Staff members
- Kazuhiro Yabana (3-alpha reaction, TDDFT)
- Yukio Hashimoto (TDHFB)
- Jun Terasaki (Double beta decay)
- Takashi Nakatsukasa [from April 2014]
- PD
- Yasutaka Taniguchi*
- Graduate students
- Yuta Fukuoka* (Config. mixing cal.) $\rightarrow$ T.N.
- Kazuyuki Sekizawa (Transfer reaction)
- Others (??)


## Nuclear response and dynamics in TDDFT [TDHF(B)]

- Yukio Hashimoto (TDHFB)
- Jun Terasaki (Double beta decay)
- Kazuhiro Yabana (TDDFT in Cond. Matt. Phys.)
- Kazuyuki Sekizawa (TDHF: Transfer reaction)


## Saturation properties of nuclear matter

- Constant binding energy per nucleon

$$
B / A \approx S_{n(p)} \approx 16 \mathrm{MeV}
$$

- Saturation density

$$
\rho \approx 0.16 \mathrm{fm}^{-3} \Rightarrow k_{F} \approx 1.35 \mathrm{fm}^{-1}
$$

- Naïve mean-field picture breaks down
- State-dependent effective interaction
- Density dependent interction
- Energy density functional

$$
E[\rho] \Rightarrow \quad h[\rho]\left|\varphi_{i}\right\rangle=\varepsilon_{i}\left|\varphi_{i}\right\rangle \quad h[\rho] \equiv \frac{\delta E}{\delta \rho}
$$

## Basic equation

- TDHF eq. (TDKS eq.)

$$
i \frac{\partial}{\partial t} \varphi_{i}(t)=\left\{-\frac{\hbar^{2}}{2 m} \nabla^{2}+V_{\mathrm{KS}}[\rho(t)]\right\} \varphi_{i}(t)
$$

- TDHFB eq. (TDBdGKS eq.)

$$
i \frac{\partial}{\partial t}\binom{U_{\mu}(t)}{V_{\mu}(t)}=\left(\begin{array}{cc}
h(t)-\lambda & \Delta(t) \\
-\Delta^{*}(t) & -(h(t)-\lambda)^{*}
\end{array}\right)\binom{U_{\mu}(t)}{V_{\mu}(t)}
$$

## Time-dependent DFT (TDDFT)

Time-dependent Kohn-Sham equation (1984)
$i \frac{\partial}{\partial t} \varphi_{i}(t)=\left\{-\frac{\hbar^{2}}{2 m} \nabla^{2}+V_{\mathrm{KS}}[\rho(t)]-\varepsilon_{i}\right\} \varphi_{i}(t)$

$$
V_{\mathrm{KS}}[\rho(t)]=V_{0}+\delta V_{\mathrm{KS}}(t)
$$

Induced (screening) field

$$
\delta V_{\mathrm{KS}}(t)=\frac{\delta V_{\mathrm{KS}}}{\delta \rho} \delta \rho(t)
$$



The collective motion is induced by the motion of the potential.

Complete analogue of the unified model by Bohr and Mottelson

## Small-amplitude approximation --- Linear response (RPA) equation ---

$$
\begin{gathered}
\left\{\left(\begin{array}{cc}
A & B \\
B^{*} & A^{*}
\end{array}\right)-\omega\left(\begin{array}{cc}
1 & 0 \\
0 & -1
\end{array}\right)\right\}\binom{X_{m i}(\omega)}{Y_{m i}(\omega)}=-\binom{\left(V_{\mathrm{ext}}\right)_{m i}}{\left(V_{\mathrm{ext}}\right)_{i m}} \\
A_{m i, n j}=\left(\varepsilon_{m}-\varepsilon\right) \delta_{m n} \delta_{i j}+\left\langle\phi_{m}\right| \frac{\partial h}{\partial \rho_{n j}}\left|\phi_{\rho_{0}}\right\rangle \\
B_{m i, n j}=\left\langle\phi_{i}\right| \frac{\partial h}{\partial \rho_{j n}}| |_{\rho_{0}} \left\lvert\, \begin{array}{l}
\text { • Tedious calculation of residual interactions } \\
\text { • Computationally very demanding, } \\
\text { especially for deformed systems. }
\end{array}\right.
\end{gathered}
$$

However, in principle, the self-consistent single-particle Hamiltonian should contain everything. We can avoid explicit calculation of residual interactions.

## Finite Amplitude Method

T.N., Inakura, Yabana, PRC76 (2007) 024318.

Residual fields can be estimated by the finite difference method:

$$
\begin{aligned}
& \delta h(\omega)=\frac{1}{\eta}\left(h\left[\rho_{\eta}\right]-h\left[\rho_{0}\right]\right) \\
& \rho_{\eta} \equiv \sum\left|\psi_{i}\right\rangle\left\langle\psi_{i}^{\prime}\right| \\
& \left|\psi_{i}\right\rangle=\left|\varphi_{i}\right\rangle+\eta\left|X_{i}(\omega)\right\rangle, \quad\left\langle\psi_{i}^{\prime}\right|=\left\langle\varphi_{i}\right|+\eta\left\langle Y_{i}(\omega)\right|
\end{aligned}
$$

Starting from initial amplitudes $\mathrm{X}^{(0)}$ and $\mathrm{Y}^{(0)}$, one can use an iterative method to solve the following linear-response equations.

$$
\begin{aligned}
& \omega\left|X_{i}(\omega)\right\rangle=\left(h_{0}-\varepsilon_{i}\right)\left|X_{i}(\omega)\right\rangle+\hat{Q}\left\{\delta h(\omega)+V_{\mathrm{ext}}(\omega)\right\}\left|\phi_{i}\right\rangle \\
& \omega\left\langle Y_{i}(\omega)\right|=-\left\langle Y_{i}(\omega)\right|\left(h_{0}-\varepsilon_{i}\right)-\left\langle\phi_{i}\right|\left\{\delta h(\omega)+V_{\mathrm{ext}}(\omega)\right\} \hat{Q}
\end{aligned}
$$

Programming of the RPA code becomes very much trivial, because we only need calculation of the single-particle potential, with different bras and kets.

## Step-by-step numerical procedure

1. Set the initial amplitudes $X^{(0)}$ and $Y^{(0)}$
2. Calculate the residual fields $\delta \mathrm{h}$ by the FAM formula

$$
\begin{aligned}
& \delta h(\omega)=\frac{1}{\eta}\left(h\left[\left\langle\psi^{\prime}\right|,|\psi\rangle\right]-h_{0}\right) \\
& \left|\psi_{i}\right\rangle=\left|\phi_{i}\right\rangle+\eta\left|X_{i}(\omega)\right\rangle, \quad\left\langle\psi_{i}^{\prime}\right|=\left\langle\phi_{i}\right|+\eta\left\langle Y_{i}(\omega)\right|
\end{aligned}
$$

3. Now, we can calculate the I.h.s. of the following equations:

$$
\left.\begin{array}{l}
\left(\omega-h_{0}+\varepsilon_{i}\right)\left|X_{i}(\omega)\right\rangle-\delta h(\omega)\left|\phi_{i}\right\rangle=V_{\text {ext }}(\omega)\left|\phi_{i}\right\rangle \\
\left\langle Y_{i}(\omega)\right|\left(\omega+h_{0}-\varepsilon_{i}\right)+\left\langle\phi_{i}\right| \delta h(\omega)=-\left\langle\phi_{i}\right| V_{\text {ext }}(\omega)
\end{array}\right\} \Rightarrow A \vec{x}=\vec{b}
$$

4. Update the amplitude to $\left(\mathrm{X}^{(1)}, \mathrm{Y}^{(1)}\right)$ by an iterative algorithm, such as the conjugate gradient method and its derivatives

Iterative approaches to strength functions:
Johnson et al., CPC 120, 155 (1999)
Toivanen et al., PRC 81, 034312 (2010); Carlsson et al., PRC 86, 014307 (2012)

## Finite amplitude method for superfluid systems

Avogadro and TN, PRC 84, 014314 (2011)
Residual fields can be calculated by

$$
\begin{array}{ll}
\delta h(\omega)=\frac{1}{\eta}\left\{h\left[\bar{V}_{\eta}^{*}, V_{\eta}\right]-h_{0}\right\} & V_{\eta}=V+\eta U^{*} Y, \quad \bar{V}_{\eta}^{*}=V^{*}+\eta U X \\
\delta \Delta(\omega)=\frac{1}{\eta}\left\{\Delta\left[\bar{V}_{\eta}^{*}, U_{\eta}\right]-\Delta_{0}\right\} & U_{\eta}=U+\eta V^{*} Y
\end{array}
$$

QRPA equations are

$$
\begin{aligned}
\left(E_{\mu}+E_{v}-\omega\right) X_{\mu \nu}+\delta H_{\mu \nu}^{20}=F_{\mu \nu}^{20} & \\
\left(E_{\mu}+E_{v}+\omega\right) Y_{\mu \nu}+\delta \widetilde{H}_{\mu \nu}^{02^{*}}=F_{\mu \nu}^{02} & \left(\begin{array}{cc}
\delta H_{\mu \nu} \\
\delta \widetilde{H}_{\mu \nu} &
\end{array}\right)=W^{+}\left(\begin{array}{cc}
\delta h & \delta \Delta \\
\delta \widetilde{\Delta}^{+} & -\delta h^{+}
\end{array}\right) W \\
& W=\left(\begin{array}{cc}
U & V^{*} \\
V & U^{*}
\end{array}\right)
\end{aligned}
$$

## FAM meets HFBRAD

## Test calculation: IS monopole

Our result: Red line qp cut-off at 60 MeV

All 2qp states are included.
Calculation by Terasaki et al.
(PRC71, 034310 (2005): Green line $\stackrel{\leftrightarrows}{\omega}$


## FAM meets HFBTHO

- I discussed with Mario about the possibility of HFBTHO+FAM
- UNEDF Annual Meeting at Pack Forest, WA, USA (2009)
- A symposium in November, 2010
- Mario visited us at RIKEN after the symposium.
- Mario and Markus started working on HFBTHO+FAM.
- The first-shot result before Christmas, 2010
- The paper was published in July, 2011



# Computational advantage in FAM 

M. Stoitsov, et al., PRC 84, 041305 (2011)

GMR in ${ }^{240} \mathrm{Pu}$ (g.s. \& f. i)
(Space: 20 major shells)


| QRPA |  |  | FAM |
| :---: | :---: | :---: | :---: |
| $v_{\text {crit }}$ | Size of $A, B$ <br> matrices | Memory <br> (in GB) | Memory <br> (in GB) |
| ${ }^{40} \mathrm{Mg}$ |  |  |  |
| $10^{-3}$ | $32039 \times 32039$ | 16.4 |  |
| $10^{-4}$ | $53386 \times 53386$ | 45.6 |  |
| $10^{-5}$ | $53823 \times 53823$ | 46.35 |  |
| $10^{-10}$ | $130936 \times 130936$ | 274.31 |  |
| $10^{-15}$ | $189271 \times 189271$ | 473.18 |  |
| $10^{-20}$ | $211159 \times 211159$ | 713.41 | 0.572 |
| 100 Zr |  |  |  |
| $10^{-3}$ | $83970 \times 83970$ | 112.81 |  |
| $10^{-4}$ | $140229 \times 140229$ | 314.63 |  |
| $10^{-5}$ | $160633 \times 160633$ | 412.85 |  |
| $10^{-10}$ | $189500 \times 189500$ | 574.56 |  |
| $10^{-15}$ | $230274 \times 230274$ | 848.41 |  |
| $10^{-20}$ | $230304 \times 230304$ | 848.64 | 0.572 |

## Explicit construction of (Q)RPA matrix with FAM

- An advantageous feature in the iterative solver with FAM (i-FAM)
- No need to calculate the (Q)RPA matrix explicitly
- Computationally fast and simple
- Disadvantage in i-FAM
- Normal-mode eigenstates are missing
- (Q)RPA matrix construction with FAM (m-FAM)
- Again, it is very easy and computationally efficient!


## RPA matrix (revisited)

$$
\begin{gathered}
\left\{\left(\begin{array}{cc}
A & B \\
B^{*} & A^{*}
\end{array}\right)-\omega^{(n)}\left(\begin{array}{cc}
1 & 0 \\
0 & -1
\end{array}\right)\right\}\binom{X^{(n)}}{Y^{(n)}}=0 \\
A_{p h, p^{\prime} h^{\prime}}=\left(\varepsilon_{m}-\varepsilon\right) \delta_{p p^{\prime}} \delta_{h h^{\prime}}+\left.\frac{\partial h_{p h}}{\partial \rho_{p^{\prime} h^{\prime}}}\right|_{\rho_{0}} \\
B_{p h, p^{\prime} h^{\prime}}=\left.\frac{\partial h_{p h}}{\partial \rho_{h^{\prime} p^{\prime}}}\right|_{\rho_{0}}
\end{gathered}
$$

FAM can provide the following quantities for a given vector

$$
\delta h_{p h}=\sum_{p^{\prime} h^{\prime}}\left(\frac{\partial h_{p h}}{\left.\partial \rho_{p^{\prime} h^{\prime}}\right|_{\rho_{0}}} \delta \rho_{p^{\prime} h^{\prime}}+\left.\frac{\partial h_{p h}}{\partial \rho_{h^{\prime} p^{\prime}}}\right|_{\rho_{0}} \delta \rho_{h^{\prime} p^{\prime}}\right)
$$

$X_{p h}$ $Y_{p h}$

## m-FAM

$\rho_{\eta} \equiv \rho_{0}+\eta \delta \rho=\sum_{i \in h}\left|\psi_{i}\right\rangle\left\langle\psi_{i}^{\prime}\right|$
Avogadro and TN, PRC87, 014331 (2013)
$\left|\psi_{i}\right\rangle=\left|\varphi_{i}\right\rangle+\eta\left|X_{i}\right\rangle, \quad\left\langle\psi_{i}\right|=\left\langle\varphi_{i}\right|+\eta\left\langle Y_{i}\right|$


$$
\begin{aligned}
& \delta \rho_{p h}=\left(\rho_{\eta}\right)_{p h} / \eta=X_{p h} \\
& \delta \rho_{h p}=\left(\rho_{\eta}\right)_{h p} / \eta=Y_{p h}
\end{aligned}
$$

$\left|X_{i}\right\rangle=\sum_{m>A}\left|\varphi_{m}\right\rangle X_{m i}, \quad\left|Y_{i}\right\rangle=\sum_{m>A}\left|\varphi_{m}\right\rangle Y_{m i}^{*}$
Adopting the following vector leads to "A" matrix $A_{p h, m i}^{m>A}$

$$
\begin{aligned}
& X_{p^{\prime} h^{\prime}}=\delta_{p^{\prime} m} \delta_{h^{\prime} i} \\
& Y_{p^{\prime} h^{\prime}}=0
\end{aligned} \longleftrightarrow \delta h_{p h}=\sum_{p^{\prime} h^{\prime}}\left(\left.\frac{\partial h_{p h}}{\partial \rho_{p^{\prime} h^{\prime}}}\right|_{\rho_{0}} \delta \rho_{p^{\prime} h^{\prime}}+\left.\frac{\partial h_{p h}}{\left.\partial \rho_{h^{\prime} p^{\prime}}\right|_{\rho_{0}}}\right|_{h^{\prime} p^{\prime}}\right)=\left.\frac{\partial h_{p h}}{\partial \rho_{m i}}\right|_{\rho_{0}}
$$

Adopting the following vector leads to " B " matrix $\quad B$ ph,mi

$$
\begin{aligned}
& X_{p^{\prime} h^{\prime}}=0 \\
& Y_{p^{\prime} h^{\prime}}=\delta_{p^{\prime} m} \delta_{h^{\prime} i} \quad \longleftrightarrow \delta h_{p h}=\sum_{p^{\prime} h^{\prime}}\left(\frac{\partial h_{p h}}{\partial \rho_{p^{\prime} h^{\prime}}}\left|\delta \rho_{\rho_{0}}+\frac{\partial h_{p h}}{\partial \rho_{h^{\prime} p^{\prime}}| |_{\rho_{0}}}\right|_{h^{\prime} p^{\prime}} \delta \rho_{i m}\right)=\left.\frac{\partial h_{p h}}{\partial \rho_{i m}}\right|_{\rho_{0}}
\end{aligned}
$$

Repeat the calculation with all possible (m,i)-pairs.
Then, all the RPA matrix elements are explicitly calculated.

## Test numerical calculation

## HFBRAD+FAM (QRPA)

- m-FAM is efficient for small matrix.
- Computational time for the m-FAM scales like $\mathrm{N}^{2} \sim \mathrm{~N}^{3}$
- i-FAM scales like N

| E(qp) cut off | $2 \times$ N(2qp) | i-FAM | m-FAM |
| :--- | :--- | :--- | :--- |
| 60 MeV | 3482 | 1 | 0.16 |
| 80 MeV | 4656 | 1.43 | 0.38 |
| 100 MeV | 5842 | 1.93 | 0.60 |
| 160 MeV | 9528 | 4.08 | 2.56 |

## FAM-(Q)RPA

N. Hinohara, M. Kortelainen, W. Nazarewicz, Phys. Rev. C 87, 064309 (2013)

- QRPA eigenmodes by contour integration in the complex frequency plane

$$
X_{\mu \nu}^{n} \propto \frac{1}{2 \pi i} \oint_{C_{n}} X_{\mu \nu}(\omega) d \omega, \quad Y_{\mu \nu}^{n} \propto \frac{1}{2 \pi i} \oint_{C_{n}} Y_{\mu \nu}(\omega) d \omega
$$



- Test application with the HFBTHO code


## Relativistic TDMF (Covariant TDDFT)

Liang, Nakatsukasa, Niu, Meng, Phys. Rev. C 87, 054310 (2013)

- Dirac sea effects are automatically included.
- Minor extra computational cost for rearrangement terms.




# Magic numbers for low-energy E1 strength 

Pygmy dipole resonance (PDR)
Inakura, et al


## Development of neutron radius

Horiuchi, Inakura, Nakatsukasa, Suzuki, PRC 86, 024614 (2012)


## Photoabsorption cross sections

Inakura, T.N., Yabana, PRC 84, 021302 (R) (2011); PRC 88, 051305(R) (2013)

E1 strength functions


## Pygmy dipoles \& neutron skins

Inakura, Nakatsukasa, Yabana, PRC 88, 051305(R) (2013); 84, 021302(R) (2011)


# Low-energy E1 strength in exotic nuclei 

 Inakura, Nakatsukasa, Yabana, PRC 84, PRC 88, 051305(R) (2013) Ebata, Nakatsukasa, Inakura, in preparation.- Constrain the neutron skin thickness and the NM EOS?
- Yes, but better in very neutron rich!
- Data on ${ }^{84} \mathrm{Ni}$ are better than ${ }^{68} \mathrm{Ni}$
- Influence the r-process?
- Significantly influence the direct neutron capture process near the neutron drip line
- We need calculation with a proper treatment of the continuum.



## Shape phase transition



## Linear response and photoabsorption cross section

Yoshida, Nakatsukasa, PRC 83, 021304(R) (2011)


## 核変形と巨大共鳴（IS，IV；L＝0～3）

Yoshida，Nakatsukasa，PRC 88， 034309 （2013）



Sm isotopes：experiment at RCNP（2003）

密度汎関数の比較


$$
m^{*} / m=0.8 \sim 0.9
$$






## 非圧縮率

$K=210 \sim 230 \mathrm{MeV}$

## TDDFT simulation of nuclear fusion reaction

Fusion barrier threshold
Guo, Nakatsukasa, EPJWC 38, 09003 (2012)


## Toward a universal Energy Density Functional (EDF)

- Improvement of the EDF is essential for accurate description of nuclear properties
- Pairing energy functional
- Correlations beyond the Kohn-Sham scheme


## Pairing energy functional

Yamagami, Shimizu, Nakatsukasa, PRC 80, 064301 (2009)

$\Delta_{n}$


$$
\eta_{1}=1 / 4, \eta_{2}=5 / 2
$$

$$
H_{p a i r}(\vec{r})=\frac{V_{0}}{4} \sum_{\tau=n, p} g_{\tau}\left[\rho, \tau_{3} \rho_{1}\right]\left\{\tilde{\rho}_{\tau}(\vec{r})\right\}^{2}
$$

$$
\cdots \Delta_{\mathrm{n}, \mathrm{HFB}}(\alpha) / \Delta_{\mathrm{n}}^{(1 / 2}
$$

$$
g_{\tau}\left[\rho, \rho_{1}\right]=1-\eta_{0} \frac{\rho}{\rho_{0}}-\eta_{1} \frac{\tau_{3} \rho_{1}}{\rho_{0}}-\eta_{2}\left(\frac{\rho_{1}}{\rho_{0}}\right)^{2} 0.5
$$

$$
0.0
$$

$\frac{\Delta_{\mathrm{n}}\left({ }^{1}\right.}{0.0}$

## DFT with proton-neutron mixing

Sato, Dobaczewski, Nakatsukasa, Satula, PRC 88, 061301(R) (2013)
pn-mixing DFT code has been developed with a collaboration with a Warsaw group.

Future subjects:
Properties of $\mathrm{T}=0$ and $\mathrm{T}=1$ pairing Charge-exchange reaction


$$
E\left[\rho_{n}, \rho_{p}\right] \Rightarrow E\left[\rho_{0}, \vec{\rho}_{1}\right]
$$

$$
\hat{H}^{\prime}=\hat{H}-\vec{\lambda} \cdot \vec{T}
$$



## Related presentation

- Time-dependent density functional theory - Yukio Hashimoto (TDHFB) - Jun Terasaki (Double beta decay)
- Kazuhiro Yabana (TDDFT in Cond. Matt. Phys.)
- Kazuyuki Sekizawa (Transfer reaction)
- Multi-reference DFT
- Yukio Hashimoto (GCM with TAC)
- Yuta Fukuoka* (Stochastic config. mixing cal.)

$$
\rightarrow \text { T.N. }
$$

- Triple-alpha reaction
- Kazuhiro Yabana (Imaginary-time approach)

